



# Simulation of heavy-ion slowing-down tracks with the SCENA code

Maxime Lamotte, G. de Izarra, C. Jammes

## ► To cite this version:

Maxime Lamotte, G. de Izarra, C. Jammes. Simulation of heavy-ion slowing-down tracks with the SCENA code. Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment, 2021, 993, pp.165075. 10.1016/j.nima.2021.165075 . cea-03118912

**HAL Id: cea-03118912**

**<https://cea.hal.science/cea-03118912>**

Submitted on 22 Jan 2021

**HAL** is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

# Simulation of heavy-ion slowing-down tracks with the SCENA code

M. LAMOTTE<sup>a</sup>, G. DE IZARRA<sup>a</sup>, C. JAMMES<sup>a</sup>

<sup>a</sup>CEA, DES, IRESNE, DER, Instrumentation Sensors and Dosimetry Laboratory,  
Cadarache,  
F-13108 Saint-Paul-lez-Durance, France.

---

## Abstract

In the frame of dependable neutron flux instrumentation development for Generation IV reactors, the French Atomic and Alternative Energies Commission (CEA) investigates an innovative technology based on optical signal produced within a fission chamber. In such gaseous detectors, neutrons interact with fissile material, releasing heavy ions in the MeV range, eventually leading to spontaneous photons emission in the ultraviolet to infrared range. We hereby present the space-time evolution of heavy-ion slowing-down tracks parameters in noble gases, as computed with the SCENA radiation-induced cold-plasma simulation tool. Preliminary results on excited-states noble-gas population dynamic are reported. Population of upper-lying gas levels is completed within picoseconds, only micrometers behind the projectile ionization trail.

*Keywords:* fission chambers, particle transport, gaseous detectors, gas scintillation

*PACS:* 29.85.-cAMODIF, 28.50.Dr, 28.41.Rc

---

## 1. Introduction

- 1 The French Atomic Energy and Alternative Energies Commission (CEA)
- 2 designed a new generation of neutron detector for Gen-IV reactors flux monitoring,
- 3 based on the luminescence of rare gases [1–3]. In so-called optical ionization
- 4 chamber, a thin layer of neutron-sensitive material, as  $^{235}\text{U}$  or  $^{10}\text{B}$  is deposited
- 5 on the inner detector surface, eventually releasing ions in the MeV/amu kinetic

6 energy range upon neutron field exposure. The slowing-down of heavy-ions with  
 7 a high ionization power in a rare gas —neon or argon, induces excitations and  
 8 ionizations along their tracks. Spontaneous photon emissions of the excited  
 9 gas atoms, in the ultraviolet to near-infrared spectrum, may be channelled in  
 10 an optical fibre, to be detected by a remote solid-state photon counter. Such  
 11 detectors, as CANOE (CApteur de Neutrons à Optique Expérimentale), have  
 12 been successfully assessed on research reactors for proof of concept validation,  
 13 sensitivity and linearity evaluation. To design, optimize and calibrate future  
 14 detectors, a physics simulation of light-generating processes has to be developed.  
 15 This paper starts with a historical review of radiation-induced cold-plasma  
 16 simulations, their hypothesis and limitations. SCENA, our self-developed fission-  
 17 induced plasma simulation code [4] is briefly described. A comparison between  
 18 previous analytical work and SCENA set with similar hypothesis is assessed.  
 19 Even though analogous results may be obtained by both methods with similar  
 20 hypothesis, a more realistic model coupling heavy-ion and electrons transportation  
 21 outputs substantially different parameters in terms of electrostatic field generation,  
 22 electron energy probability function and decay times. Future developments and  
 23 discussions are presented in the conclusion.

24

## 25 **2. Methods**

26 Theoretical investigations of radiation-induced cold-plasma, especially fission-  
 27 fragments generated, have been a challenge for the past 50 years. Potential  
 28 applications of such plasmas included nuclear-lasers, spaceship propulsion and  
 29 optical ionization chambers for fission and fusion reactors monitoring. Several  
 30 authors attempted to model these non-thermal plasma by postulating a Partial  
 31 Local Thermodynamic Equilibrium (PLTE), and an established steady-state  
 32 regime [5–7], finally solving a set of Boltzmann equations [8]. Electron Energy  
 33 Distribution Functions (EEDF) were found to be almost a Maxwellian around  
 34 thermal energies [9–11], with a low yield energy tail up to keV values. As

the calculated mean energy was well below gases' inelastic reactions thresholds, the high energy part of the electron spectrum was considered responsible for metastable and excited states population [12–14]. While such hypothesis can be assumed in homogeneous low-pressure plasmas with intense energy deposition rates—as in nuclear-lasers—they however cannot be retained in a low-energy deposition rate systems having spatial heterogeneity, as noble gases excited by low rates fission fragment sources.

With respect to Optical Emission Spectroscopy (EOS) analysis of alpha-particles and fission fragment induced cold-plasmas in an analytical and a prototype optical ionization chamber, no recombination continua were detected and no emissions from ionised states of rare gases indicated very low ionisation rates, incompatible with PLTE assumptions [2].

We chose, as Budnik et al. [15] to study individual slowing down tracks mechanisms along a space and time evolution in an attempt to build future radiation-plasma macro-model, considering the plasma as an assembly of independent ionization tracks. Population states of this problem is achieved by a Monte-Carlo solved particle transport computer code developed previously, the Simulation of Collisions Electrons-Neutrals in Atmospheres (SCENA) [4]. SCENA allows a 3-dimensions and time coupled analysis of both projectile ions and delta-electrons interactions with a monoatomic buffer gas, as found in fission chambers.

### 2.1. *Physics coupling in heavy-ion slowing-down mechanism*

Typical times encountered in the generation and die-away of a heavy-ion track may differ by several orders of magnitudes, allowing uncoupled mathematical simulation of some physical processes. We hereby review main time scales of interactions mechanisms.

Fission fragments generated in a fissile coating may be released at kinetic energies up to 99 MeV, corresponding to an initial velocity of about  $14\text{E}+6$  m/s for an ideal Light-Fission Fragment (LFF, atomic mass = 96). Heavy ions' stopping powers, computed from HKS semi-empirical model revised by Stolterfoht [16] or from ICRU 73 tables [17, 18], input in a Continuously Slowing Down Approximation

65 (CSDA) solved with explicit Euler integration scheme provide an estimation of  
 66 their ranges, slowing-down times and secondary-electrons energy spectra.  
 67 LFF in 1 atm neon crosses the first millimetre from emission layer's surface in  
 68 72 pico-seconds but completely transfers its kinetic energy in about 45 mm, a  
 69 distance covered in 11 ns.  
 70 Energetic electrons originating from the heavy-ion/buffer gas interaction, with  
 71 an average kinetic energy of about 40 to 50 eV, may ionize or excite the gas  
 72 until falling below its first excitation threshold, in pico-seconds.  
 73 The similar lifetimes in a plasma sheath of heavy-ions and super-threshold  
 74 secondary electrons accordingly implies necessity for a coupled simulation of  
 75 both fission fragment transport and ionization tracks die-away. Conversely,  
 76 thermalization of subthreshold electrons, sterile for luminescence production in  
 77 the ultraviolet to near-infrared spectrum, takes significantly longer times scales,  
 78 in the order of a microsecond. Such low-energy, long-lasting electrons may be  
 79 discarded from the simulation if no external energy input, as an electric field, is  
 80 present.  
 81 These times can be put in perspective with the tens of nanoseconds of atomic  
 82 spontaneous emission and of the most probable atomic collision time in a 1 atm  
 83 gas at room temperature, 1 ns (for a  $1.5\text{\AA}$  diameter Argon atom at a thermal  
 84 mean velocity of 400 m/s).

85

## 86 *2.2. Heavy-ion transport*

87 The present section sums-up physical models implemented in SCENA for  
 88 simulation of processes considered in excited levels population induced by heavy-  
 89 ions-to-gas energy transfer. The following hypothesis were applied;

- 90 • A single heavy-ion is emitted in an homogeneous medium (gas) with a  
 91 discrete initial energy and charge
- 92 • Heavy-ions travel in straight paths following Continuously Slowing-Down  
 93 Approximation (CSDA)

- 94 • Heavy-ions only transfer energy to the gas through ionisations along their  
95 tracks, no direct-excitation is considered, neither nuclear collisions
- 96 • Heavy-ions having potential ranges of less than a millimetre are discarded  
97 from simulation
- 98 • Ejected delta electrons may ionize, excite or heat-up the gas
- 99 • No recombinations nor metastable states electron-impact de-excitations  
100 are considered

101 An external electric field, normal to the heavy-ion track may be applied, as this  
102 function was already implemented in SCENA to simulate standard ionisation  
103 chambers, estimate metastable states population kinetics and validate numerical  
104 models against other Boltzmann equation solvers in reference cases. Reduced  
105 fields to be found in fission chambers are well below 5 Td, thus no Lorentz-force  
106 is applied on the swift heavy-ion.

### 107 *2.3. Electron generation and transport*

108 The heavy-ion slowing-down profile in a buffer gas, subdivided in millimetre-  
109 long segments with averaged parameters such as charge, velocity or stopping-  
110 power allows uncoupled computation of optical spectra in regions of interest ie.  
111 around the emission layer, or at the end of heavy-ion track.

112 The Singly Differential Cross-Section (SDCS), obtained with HKS formula revised  
113 by Stolterfoht [16], calculated along each millimetre characterises the initial  
114 delta-electron emission energy spectrum and population.

115 During their free-flight, electrons move given the resulting Lorentz-force due to  
116 an external or the self-induced electrostatic field —if considered— and their own  
117 inertia. Excitations, ionisations and elastic collisions probabilities are computed  
118 and recorded at each random census-time —in the order of tens of femto-seconds  
119 for analysis, as well as the Electron Energy Distribution Function (EEDF).

120

#### 121 2.4. *Electrostatic field*

122 Local electrostatic field within a heavy-ion plasma sheath is generated by  
 123 charged particles production along the slowing down track. Free electrons and  
 124 ions of several generations increase swarm's size, resulting in a local difference  
 125 of potential, to evolve with time, gas composition and speed of seed electrons,  
 126 moving substantially faster than ions.

127 This self-induced electrical field may shift EEDF, eventually impacting the  
 128 excited gas levels population. To analytically estimate an electrostatic field  
 129 around the heavy-ion track, Budnik et al. proposed the following hypothesis;

- 130 • Heavy-ion travel time is considered negligible in regard of the primary  
 131 electrons die-away time
- 132 • Plasma track is cylindrical and radially symmetrical
- 133 • Ionization events do not create another generation of electrons
- 134 • No diffusion of resonance radiation are modelled
- 135 • Attachment occur if an electron is below a given distance of an atom
- 136 • Initial electrons are set along a corona after few (not considered) interactions,  
 137 with a Gaussian radius distribution
- 138 • Initial electron and ion populations in helium are set to  $\langle r_e^2 \rangle = 4\text{E-}8 \text{ cm}^2$
- 139 • Energy degradation by atomic impact, molecular ion formation and recombination

The first hypothesis being questionable with respect to typical process times encountered in a plasma track calculated in Sect.2.1, will solely be implemented in SCENA for comparison purposes. A Particle In Cell (PIC) is set up to discretize a semi-infinite cylinder of 100  $\mu\text{m}$  radius, representing the plasma sheath, into 50 cylindrical subshells.

To estimate the electric field inside an thin subshell of radius ranging between  $r$  and  $r + dr$ , with a charge density  $\rho_e$  computed by counting particles within shell's boundaries, we set-up a classical electrostatic approach.

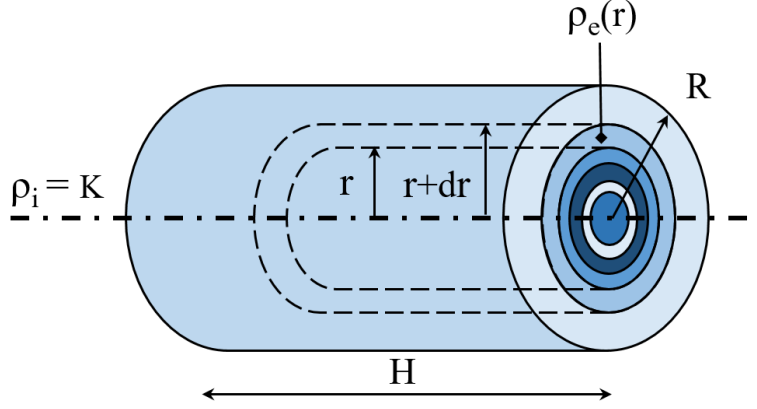


Figure 1: Electrostatic model along the heavy-ion slowing-down track of length  $H$  and radius  $R$ . Charge densities varying with respect to distance from centre of the plasma sheath is denoted  $\rho_e(r)$ , ion linear charge density produced by heavy-ion impact along the track is referred as  $\rho_i$ .

It is noteworthy that we considered positive ions standstill regarding their thermal motion velocity against the tens of eV of free electrons. In the picoseconds studied in our simulations, where gas may be excited by electrons, ions remain in their birth subshells.

Fig. 1 shows the electrostatic field model set in a new SCENA function to compute analytically electric field around the semi-infinite heavy-ion track. Two sets of equations define electric field generated by charges in cylindrical shells, and by gas ions along a central line. Track-length segment  $H$  of the heavy-ion being considered infinitely longer (1 mm) than its radius  $R$ , of about a micron, by applying Gauss theorem, we define:

$$\frac{\sum Q}{\varepsilon_0} = \oint \vec{E} \cdot d\vec{A} \quad (1)$$

The volume of a sheath located at radius  $r$  is simply:

$$V_s = 2 H \pi r dr \quad (2)$$



electric field generated by free charges  $E_e$  at a radius  $r_s$  is:

$$\frac{\int_0^{r_s} \rho H 2 \pi r dr}{\varepsilon_0} = 2 E_e \pi r_s H \quad (3)$$

finally;

$$E_e(r_s) = \frac{\int_0^{r_s} \rho r dr}{\varepsilon_0 r_s} \quad (4)$$

Gas ionisations due solely to heavy-ion impacts generate positive charges along a linear track, inducing an electric field  $E_i$  described by Eq 5.

$$E_i(r_s) = \frac{\bar{N}_\delta}{2 \pi \varepsilon_0 r_s} \quad (5)$$

140 where  $\bar{N}_\delta$  is the ionization events number on the heavy-ion track's segments, as  
 141 computed by SCENA's slowing-down module.

### 142 3. Results

143 Results on population of excited states, their locations and Electron Energy  
 144 Distribution Functions (EEDF) were computed as function of the time elapsed  
 145 since departure of the projectile heavy-ion during its first millimetre journey  
 146 in a buffer gas. As a matter of comparison with other authors, local electron  
 147 densities, electric fields and EEDF were assessed, assuming an instantaneous  
 148 transport of the projectile and a cylindrical symmetry around the semi-infinite  
 149 heavy-ion trajectory, for various radius and census times, neglecting transport  
 150 of the latter. All results presented below were computed with parameters to  
 151 be found in a CANOE fission chamber as exploited on ORPHÉE and CABRI  
 152 research reactors. Table 1 sums-up simulation inputs of our particle transport  
 153 in the case of excitation produced by an ideal Heavy Fission Fragment (HFF).

154

	Projectile	HFF
	$E_0$ (MeV)	68
	A (amu)	130
	Z	54
	$Z_{\text{eff}_0}$	13.8 [4]
	Target	Neon
	Density ( $\text{cm}^{-3}$ )	2.68E+19
	Temperature (K)	300
	HFF-impact ionisations (first mm)	7E+4
	Mean electron energy (first mm, eV)	45

Table 1: Input values for electron-slowng down simulation in a CANOE optical ionization chamber.

### 155 3.1. Evolution of secondary electrons in a semi-infinite sheath

156 Initial HFF-impact ionizations develop continuous distributions of electrons  
157 described by the corresponding SDCS, with an average energy —45 eV— well  
158 above neon’s first ionization threshold —21.56 eV—. As depicted Fig. 2, EEDF  
159 in the centre of the plasma sheath shortly after projectile’s passage is mainly  
160 composed of slow first-generation electrons produced by heavy-ion impact ionizations.  
161 At further locations, EEDF present harder spectra, consisting of fast electrons  
162 that could travel the long distance in such short time, and low-energy secondary  
163 electrons produced by fast super-threshold first-generation electrons.

164 Similarly, on a longer time scale, super-threshold electrons being moderated  
165 down to sub-ionization level, energy spectra reflect a swarm being mostly scattered  
166 elastically. Fast super-threshold electrons, able to produce ionizations and  
167 excitations induce a shoulder on the energy spectrum visible around 16.62 eV,  
168 neon’s first excitation level. At this given time, no further increase of swarm’s  
169 size is excepted and only few excitations up to lower-lying  $^1S_x$  levels can be  
170 provoked until mitigation of electron’s energy down to thermal motion.

171

172 A close-up on electron swarms at shorter distances from projectile track’s

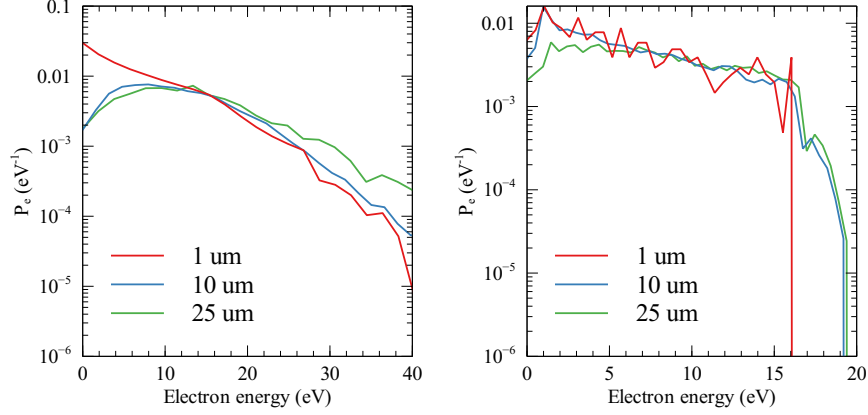


Figure 2: EEDF of electrons generated by both ion-impact and electron impact ionization on neon at 1 atm at various sheath radius, 10 ps (left) and 100 ps (right) after a projectile instantaneous passage.

173 centre, in the micrometre range as on Fig. 3, highlights secondary electrons  
 174 emission mechanism. Initial distribution, characterized by SDCS with a tail up  
 175 to keV values at HFF first millimetre, quickly falls below ionization threshold  
 176 value. As super-threshold electrons may travel long distances, up to hundreds  
 177 of micrometers, high energy tail remains visible for extended census times at  
 178 high radius, to be shifted down to thermal energies with a delay function of  
 179 inner radii.  
 180 Near the end of the heavy-ion track, where only low-energy electrons are initially  
 181 emitted, expansion of the sheath takes longer time and does not contain subthreshold  
 182 electrons past few micrometers, resulting in a lower gas excitation probability.

### 183 3.2. Evolution of electrostatic field a semi-infinite sheath

184 Electron densities around the heavy-ion track are subjected to two opposite  
 185 mechanisms. Fast expansion of the swarm initially generated along a line, tends  
 186 to decrease electron density as function of radius and time. On the other hand,  
 187 as ionizations events occur, generation of ions and free electrons increases charge  
 188 densities in some shells.

189 Fig. 4 presents electron densities around a HFF slowing-down track in neon,  
 190 along its first millimetre, from which electrostatic field calculations have been

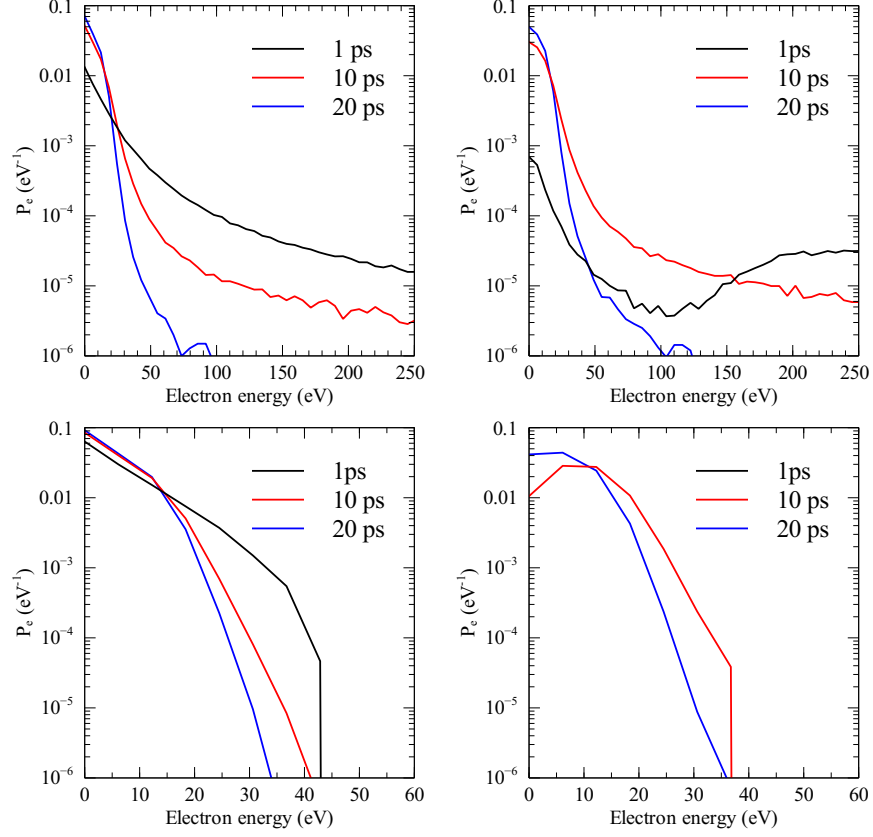


Figure 3: Up: EEDF of electrons generated by both ion-impact and electron impact ionization on neon at 1 atm at various census time, along the first mm crossed by HFF, in the shell of radius 0 to  $1\text{E-}4$  m (left), and  $1\text{E-}5$  to  $1\text{E-}4$  m (right). Down: same parameter but at the last computed heavy-ion track millimetre segment, with an initial kinetic energy of 1.4 MeV. No electrons had time to reach the external subshell within 1 ps.

191 performed.  
 192 Heavy-ions displaying higher stopping power in denser gases, associated electrostatic  
 193 field substantially increases charged particles production per unit length, and  
 194 so reduced electric field in the first picoseconds of the electron dying-away.  
 195 Self-induced electric-fields up to several MV/m may generate an electrostatic  
 196 force, oriented radially and quickly decreasing after ion's passage. In 1 atm  
 197 xenon, this electrostatic field located 0.15  $\mu\text{m}$  from track's centre, 0.2 ps after  
 198 ion's passage rises a local reduced field of 68 Td, resulting in a radial acceleration  
 199 of  $3\text{E}+17 \text{ m.s}^{-2}$ . Despite their high values, self-induced electric forces don't  
 200 induce significant changes in global EEDF and buffer gas excitation level population  
 201 apportionment because of a fast decay.

202

### 203 *3.3. Global yield and position of inelastic events*

204 Activation of the heavy-ion time-resolved transport routine and electron  
 205 seeding of SCENA allows complementary investigation of the electron swarm  
 206 parameters from a more realistic point of view. In such case, new electrons with  
 207 random direction vectors and initial velocity computed from SDGS are added  
 208 at each simulation census-time.

209 Fig. 5 plots the total excitation rate of neon's level  $^2P_6$  (18.64 eV) responsible for  
 210 692.95 nm line emission and  $^1S_5$  lowest metastable level (16.62 eV) as function  
 211 of time in the first millimetre long sheath of plasma. A maximum swarm size  
 212 is obtained 112 ps after passage of the projectile. This value is coherent with  
 213 the sum of time for HFF to cross the cell, 105 ps, and time for SDGS-described  
 214 electron swarm to fall below neon's ionisation threshold,  $\sim 3$  ps.

215 In the presence of an external electric field, such as in standard fission chambers,  
 216 a drift of electrons and so ionisations events' positions is noticeable, as depicted  
 217 Fig. 6. If the applied reduced electric field is above breakdown voltage (about  
 218 1 Td for 1 atm neon), inelastic events keep increasing after disappearance of the  
 219 projectile from the cubic-sided millimetre cell, as well as a general deviation of  
 220 swarm's position from the track centre.

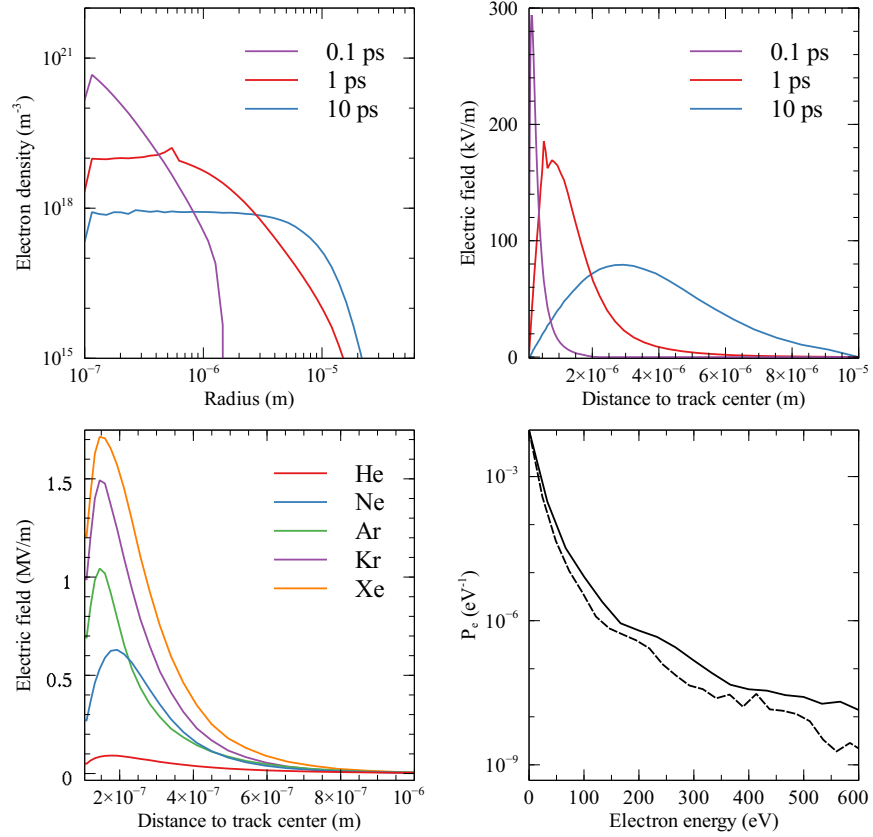


Figure 4: Up: Electron densities at various radii and census-times in 1 atm neon bombarded by HFF (left) and associated electric field (right) in the assumption of instantaneous projectile journey over 1 mm.

Down: Impact of the buffer noble-gas specie on the electric field in the corona around the projectile, 0.2 ps after seed-electron release (left). EEDF in 1 atm Xe 100 ps after seed-electron release with (solid line) and without (dashed line) considering self-induced electrostatic field (right).

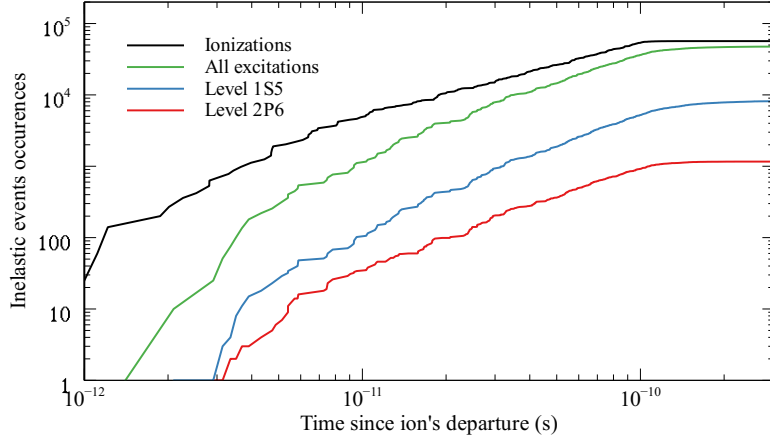


Figure 5: Evolution of accumulated excitation and ionization events of 1 atm neon by secondary HFF-induced electrons as function of time after projectile's passage.

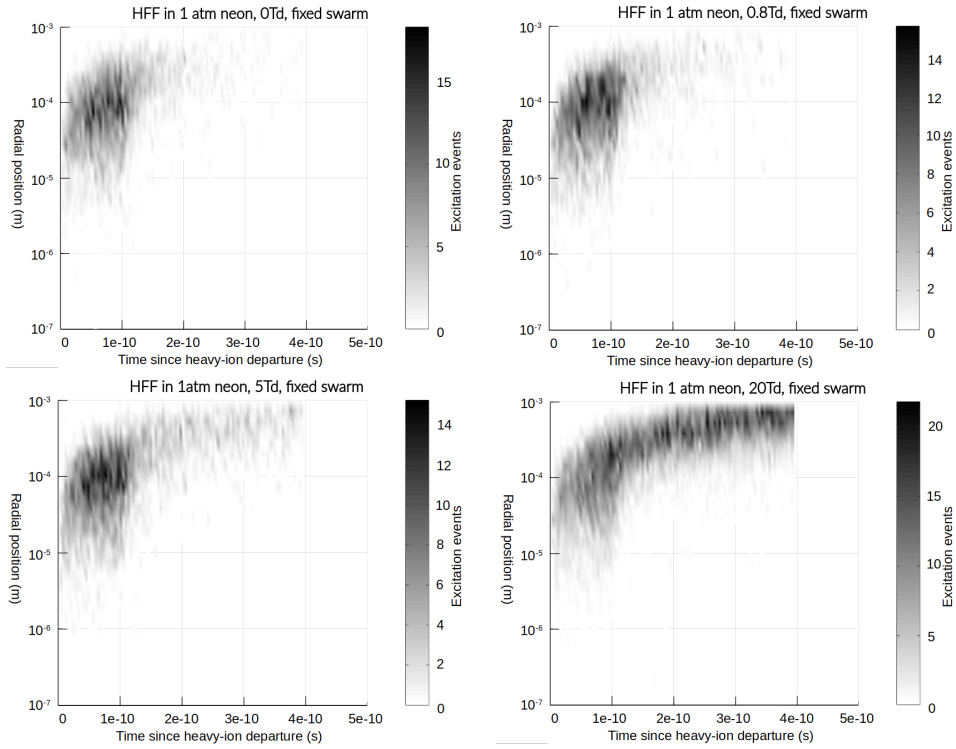


Figure 6: Radial position of excitation events from the heavy-ion track-centre in a 1 mm-sided neon cubic cell as function of time since ion's departure, in the presence of an external electric field of: up left 0 Td, up right 0.8 Td, down left 5 Td, down right 20 Td. In all cases, the HFF exits the cell after 105 ps.

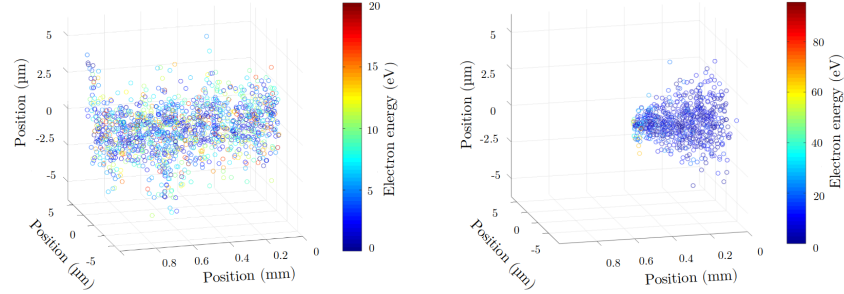


Figure 7: 3-Dimensional views of the electron swarm energy distribution (colour circles) produced after a 50 ps simulation. Left: with instantaneous projectile transport hypothesis. Right: with coupled transportation of the HFF. In both cases, about 1000 electrons are displayed.

### 3.4. Space and time evolution of the global electron swarm

A 3-dimensional representation of the electron swarm generated around a HFF track, with only 1/200 th of electron population to be found in real cases is presented Fig. 7. Fast degradation of secondary electron's energy, as detailed in section 3.1 and clearly visible on 3-D views lead to an overall low EEDF over the complete plasma track. Fig. 8 shows the global EEDF of a 1 mm plasma track as function of time. Such distribution, while different from thermal motion distribution remains relatively energetic for a long time, in the order of microseconds without other diffusion mechanisms, but is sterile for further excitation level population.

Inelastic events production, useful for luminescence generation appears to be located close around the projectile heavy-ion (in micrometres radii), and mostly along the first millimetres of its track, as further steps induce low-energy deposition rates.

## 4. Conclusion

A simulation of heavy-ion slowing-down and energy-transfer to noble gas has been performed in configurations to be found in optical ionization chambers. Secondary electron swarms have been analysed by their space and time energy spectra evolutions. Fast degradation of electrons energy implies population of



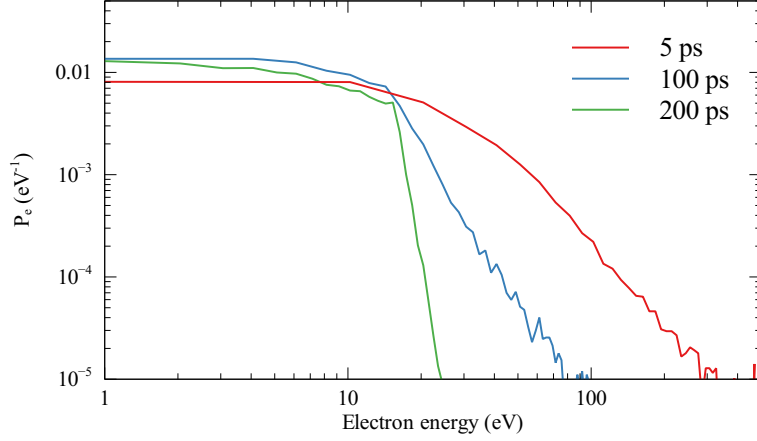


Figure 8: EEDF of electrons generated by both ion-impact and electron impact ionization on neon at 1 atm along the projectile's passage in the first 1 mm-segment. Times are given since ion departure from fissile layer. HFF is expected to leave the mm-sided cell after about 105 ps.

240 near-infrared emitting gas excitation levels to be complete picoseconds after  
 241 heavy-ion passage, in the very first millimetres of its track, and around a  
 242 radius of few micrometres. Electrostatic field and associated force generated  
 243 by free flying charges around a heavy-ion track have been estimated and display  
 244 values up to kV/cm in helium, and up to MV/cm in xenon, modifying slightly  
 245 electron energy spectrum distribution in the corona around swift heavy-ion.  
 246 Nevertheless, such narrow-located phenomenon cannot be observed in a more  
 247 realistic simulation, considering swift heavy-ion time of flight as non null. Typical  
 248 reaction times found within an ion-generated plasma track unveil requirement of  
 249 a coupling of projectile and secondary electron swarm dynamic equations to be  
 250 representative of such non-equilibrium transient plasma. Future improvements  
 251 of SCENA may include collisionnal de-excitation model, as pressures in fission  
 252 chambers induce atomic collision times in the order of excited-states radiative  
 253 decay constants.

## 254 References

- 255 [1] M. Lamotte, G. De Izarra, C. Jammes, Development and first use  
256 of an experimental device for fission-induced spectrometry applied to  
257 neutron flux monitoring, Nuclear Instruments and Methods in Physics  
258 Research Section A: Accelerators, Spectrometers, Detectors and Associated  
259 Equipment 953 (2019) 163236.
- 260 [2] M. Lamotte, G. De Izarra, C. Jammes, Heavy-ions induced scintillation  
261 experiments, Journal of Instrumentation 14 (09) (2019) C09024.
- 262 [3] M. Lamotte, G. De Izarra, C. Jammes, Design and irradiation test of  
263 an innovative optical ionization chamber technology, Nuclear Instruments  
264 and Methods in Physics Research Section A: Accelerators, Spectrometers,  
265 Detectors and Associated Equipment 968 (2020) 163945.
- 266 [4] M. Lamotte, G. De Izarra, C. Jammes, SCENA: A simulation tool for  
267 radiation-induced gas scintillation, Nuclear Instruments and Methods in  
268 Physics Research Section A: Accelerators, Spectrometers, Detectors and  
269 Associated Equipment 982 (2020) 164576.
- 270 [5] J. Guyot, G. Miley, J. Verdeyen, Application of a two-region heavy charged  
271 particle model to noble-gas plasmas induced by nuclear radiations, Nuclear  
272 Science and Engineering 48 (4) (1972) 373–386.
- 273 [6] D. Rees, C. Leffert, D. Rose, Electron density in mixed gas plasmas  
274 generated by fission fragments, Journal of Applied Physics 40 (4) (1969)  
275 1884–1896.
- 276 [7] J. E. Deese, H. Hassan, Analysis of nuclear induced plasmas, AIAA Journal  
277 14 (11) (1976) 1589–1597.
- 278 [8] H. Hassan, J. E. Deese, Electron distribution function in a plasma generated  
279 by fission fragments, The Physics of Fluids 19 (12) (1976) 2005–2011.

- 280 [9] B. S. Wang, G. H. Miley, Monte Carlo simulation of radiation-induced  
281 plasmas, *Nuclear Science and Engineering* 52 (1) (1973) 130–141.
- 282 [10] R. H. Lo, G. H. Miley, Electron energy distribution in a helium plasma  
283 created by nuclear radiations, *IEEE Transactions on Plasma Science* 2 (4)  
284 (1974) 198–205.
- 285 [11] D. Auphelle, F. Euvé, M. Fitaire, A. Pointu, M. Vialle, et al.,  
286 Caractéristiques électroniques d’un plasma créé par un faisceau d’ions  
287 accélérés, *Physica B+ C* 97 (2-3) (1979) 235–243.
- 288 [12] G. Russell, Feasibility of a nuclear laser excited by fission fragments  
289 produced in a pulsed nuclear reactor, *NASA SP-236* (1971) 53–62.
- 290 [13] A. Pointu, D. Auphelle, F. Euve, M. Fitaire, M. Vialle, Calculation of the  
291 electron distribution function of a rare gas nuclear induced plasma, *Journal*  
292 *de Physique* 41 (10) (1980) 1101–1108.
- 293 [14] N. Peyraud, Energy transfer theory in particle-beam generated plasmas,  
294 *Physics Letters A* 106 (1-2) (1984) 37–42.
- 295 [15] A. Budnik, Y. V. Sokolov, A. Vakulovskiy, Mathematical simulation  
296 of the space-time evolution of fission-fragment plasma tracks, *Hyperfine*  
297 *Interactions* 88 (1) (1994) 185–192.
- 298 [16] N. Stolterfoht, R. D. DuBois, R. DuBois, R. D. Rivarola, Electron emission  
299 in heavy ion-atom collisions, Vol. 20, Springer Science & Business Media,  
300 1997.
- 301 [17] P. Sigmund, R. Bimbot, H. Geissel, H. Paul, A. Schinner, ICRU - report  
302 73, Stopping of ions heavier than helium, *J. ICRU* 5, 1 (2005).
- 303 [18] P. Sigmund, A. Schinner, H. Paul, Errata and addenda for IRCU report  
304 73, Stopping of ions heavier than helium, *J. ICRU* 5, 1 (2009) 1–10.