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A 5D gyrokinetic full- f global semi-lagrangian code for flux-driven ion turbulence simulations

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Abstract

This paper addresses non-linear gyrokinetic simulations of ion temperature gradient (ITG) turbulence in tokamak plasmas. The electrostatic GYSELA code is one of the few international 5D gyrokinetic codes able to perform global, full- f and flux-driven simulations. It has also the numerical originality of being based on a semi-Lagrangian (SL) method. This reference paper for the GYSELA code presents a complete description of its multi-ion species version including: (i) numerical scheme, (ii) high level of parallelism up to 500k cores and (iii) conservation law properties.

Keywords: plasma turbulence, gyrokinetic global full- f flux-driven simulations, semi-lagrangian method, high-performance computing

1. Introduction

In magnetic fusion devices, the power gain strongly increases non linearly with the energy confinement time. As a matter of fact, the quality of the plasma energy confinement largely determines the size and therefore the cost of a fusion reactor. This confinement time is mainly governed by the plasma turbulence which develops in such devices – leading to fluctuations with relative magnitude of a few percents in the hot core – and the associated transport. Understanding its origin and properties in view of a possible control is one of the critical issues in fusion science [58]. The inhomogeneities in density, temperature, and magnetic field place the plasma naturally out of thermodynamical equilibrium, and tend to excite several micro-instabilities over a wide spectral range. These plasmas exhibit a low collisionality so that conventional fluid models are not appropriate and kinetic descriptions are required. In such first-principle descriptions of plasmas, the six dimensional evolution equation for the distribution function – Vlasov or Fokker-Planck equations – is solved for each species, coupled to the self-consistent equations for the electromagnetic fields, namely Maxwell's equations. Fortunately, as far as turbulent fluctuations are concerned, they develop at much lower typical frequencies than the high frequency cyclotron motion. Therefore, this 6D problem can be reduced to a 5D one. The useful part of the distribution function then evolves in a five dimensional phase space generated by four slow variables and an adiabatic invariant. This model is known as the *gyrokinetic* model. But even with this dimensional reduction, building such a simulation tool remains quite challenging.

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The first-principle gyrokinetic codes, developed for this purpose, make already an intensive use of massively parallel supercomputers and require state-of-the-art high performance computing (HPC). Nevertheless, solving 5D non-linear gyrokinetic equations for several ion species proves so challenging that to date no code is able to treat all the physics involved. Several strategies based on different simplifications have been developed to decrease these extreme numerical costs. A first strategy is to reduce the simulation domain to a small plasma volume aligned with the magnetic field lines and sufficiently small to neglect the radial variations of macroscopic quantities such as the density and temperature fields and their gradients. Such so called *flux-tube* codes (*e.g.* GS2 [38], GYRO [19]) have the advantages to drastically reduce CPU time and memory consumption compared to the *global* approach where computing over the whole plasma domain requires huge 3D meshes. Despite this constraint, a global approach must be considered whenever the “global” scales are comparable to that of the turbulence. This allows one in particular to address some aspects of the barrier physics. Gyrokinetic models can also be split in two distinct families with respect to the representation of the distribution function: either full- f or δf models. In the δf model, only perturbations with respect to some prescribed background equilibrium (usually Maxwellian in velocity) are computed. Conversely, in full- f models, the whole distribution function is evolved. Especially, the back reaction of turbulent transport is accounted for in the time evolution of the equilibrium. In full- f simulations, the turbulent regime is evanescent if no free energy is injected in the system to prevent the inevitable relaxation of equilibrium profiles below the – linear or non-linear – threshold of the underlying instability. A heat source is mandatory in view of exploring the long time, typically on energy confinement times, behavior of turbulence and transport, leading to the so-called *flux-driven* simulations.

Finally, the codes differ also by their numerical schemes which have evolved all along the last twenty-five years, in direct link with the evolution of HPC resources. Historically, particle in cell methods (PIC) [10] have been most popular, and represent widely adopted approaches to numerical simulations of kinetic plasmas. They used to be considered as the most efficient tool to describe plasma dynamics, essentially because they are capable of describing many physical phenomena in the full dimensional case, at relatively small computational costs. So many gyrokinetic codes are PIC codes. Let us mention, Parker’s code [78], Sydora’s code [96], PG3EQ [36], GTC [72], ELMFIRE [52], GT3D [56], ORB5 [11, 60], and GTS [102]. However, it is well known that the relative numerical noise inherent to PIC methods constitutes a strong limiting factor to accurately describe the distribution function in phase space on long time scales. Moreover, the numerical noise only slowly decreases, like $1/\sqrt{N}$, when the number N of particles is increased. The main problem for non-linear gyrokinetic simulations is that the noise level *a priori* accumulates in time [77]. Even small errors in the evaluation of these moments can cause a systematic corruption of the simulation results on relatively short periods of time. Consequently, the reduction of numerical noise has been right from the start a matter of intense research, and many improvements have taken place during the past ten years. More specifically by applying different statistical methods: (i) control variate method of variance reduction (with δf scheme [5, 78] or “adjustable control variate” method [49]); (ii) importance sampling (with “optimized loading” [50]). These techniques have enabled to reduce the numerical noise by orders of magnitude. Let us specifically mention the achievements made on the ORB5 gyrokinetic PIC-code [60] on the noise issue, which are summarized in [99]. Another approach to avoid the issue of marker sampling noise is the eulerian approach. It consists in discretizing the phase space on a fixed grid, and in applying finite differences, finite volumes and/or Fourier transforms to model the differential and integral operators (see *e.g.* [18] for a review). The main drawback stems from the fact that these numerical schemes are based on explicit time integration, so they are limited by the Courant-Friedrichs-Lewy (CFL) stability condition, which slaves the maximum time step to the grid space resolution. Several gyrokinetic codes are based on this approach as proved by this non-exhaustive list of eulerian codes: GS2 [61, 38], GYRO [19], GENE [59, 74], GKV [103], GKW [81] and GT5D [55, 57].

As described in the following, one peculiarity of the GYSELA code is to be based on a semi-lagrangian method, which is a mix between PIC and eulerian approaches. Each of these three numerical methods (see [46] for a review) has advantages and drawbacks. But for sure, the wide variety of gyrokinetic codes, coming from all these different choices, is a strength for the fusion community. Indeed, due to the extremely challenging computing requirements, each gyrokinetic code runs at the limit of its applicability range. For instance significant advances like taking into account kinetic electrons and electromagnetic effects are

still difficult to achieve in global codes. Conversely, the full- f and flux-driven regimes are two necessary ingredients to investigate, among others, the mechanism underlying transport barrier creation. Benchmarks between the different approaches are therefore primordial (*e.g.* [97, 40]). Difficulties reside generally in finding common domains of validity (set of parameters, initial and boundary conditions,...). The choice of a global full- f approach has been done for the electrostatic non-linear 5D gyrokinetic code GYSELA described in this paper. Electrons are at present considered adiabatic. All the components of this semi-Lagrangian code will be detailed both in terms of physical equations and numerical methods as well as the verification and benchmarks that have been performed. Such flux-driven ITG simulations are extremely challenging and would not be possible without a high level of parallelism which will also be addressed. The code is already currently using Petaflops HPC resources and is actively preparing its evolution for the future exascale era. The paper is a comprehensive description of the multi-ion species version of the code. It is organized as follows. The physical model is presented in section 2. The numerical methods and the hybrid OpenMP/MPI parallelism are described in section 3. The code verification via invariance tests and benchmarks are detailed in section 4. The conservation law properties are analyzed in section 5. Discussion and conclusion close the paper in section 6.

2. GYSELA gyrokinetic global full- f model

GYSELA is a global code presently used with a simplified concentric circular magnetic configuration (section 2.2) similar to the Tore Supra equilibria. The new version of the code still considers adiabatic electrons but the possibility to address transport of impurities has been added. The time evolution of the full distribution function of each ion species (major species as *e.g.* Deuterium + one minor impurity) is governed by a 5D non-linear gyrokinetic Vlasov equation (section 2.3) self-consistently coupled to a 3D poisson equation (section 2.4). The required gyroaverage operator (section 2.5) that used to be approximated by a Padé expansion can be applied with a direct average on gyro-circles. Collisions are taken into account. With a linearized intra-species collision operator (section 2.7) neoclassical effects are recovered. Inter-species collisional transfers are now also considered (section 2.8). The problem of initialization and radial boundary conditions inherent to global full- f codes as GYSELA are addressed in section 2.6. Concerning boundary conditions, three modes are available in the code: (i) the *fixed-gradient* mode where the temperature profile is fixed at both radial boundaries, corresponding to decaying turbulence regimes (relaxation of equilibrium profiles below instability thresholds cannot be avoided); (ii) the *gradient-driven* mode, equivalent to the previous one but where gradient profiles are maintained by an artificial Krook-type operator and finally (iii) the *flux-driven mode* (the most often used) where temperature is still fixed at the outer boundary but can evolve freely at the inner one. The turbulence is forced with a constant-in-time incoming flux generated by a heat source independent of the distribution function (section 2.9) leading to possible long-time simulation. A simulation over several confinement times has been recently performed for comparison to Tore-Supra experiments [35].

2.1. Toroidal coordinate system

Let us introduce the notations used in the paper. We consider a set of coordinates labelled $\{x^i\}$, the metric tensor $\{g_{ij}\}$ is the product of the transposed Jacobian matrix J^T and the Jacobian matrix J , *i.e.* $\{g_{ij}\} = J^T J$. For a set of cartesian coordinates X^i , the elements J_{ij} of the Jacobian matrix are defined as $J_{ij} = \partial_{x^j} X^i$. Let g represents the determinant of the metric tensor (*i.e.* $g = \det\{g_{ij}\}$), then the Jacobian in space J_x is defined as $J_x = \sqrt{g}$ and is equal to $J_x = [(\nabla x^1 \times \nabla x^2) \cdot \nabla x^3]^{-1}$, *i.e.* the volume element is $J_x d^3\mathbf{x}$. The tensor $\{g^{ij}\}$ is the inverse of the tensor $\{g_{ij}\}$. The element of the contravariant metric tensor verifies the relation $g^{ij} = \nabla x^i \cdot \nabla x^j$. With these notations, each vector $\mathbf{A}$ can be defined in terms of its covariant components A_i as $\mathbf{A} = A_i \nabla x^i$ and the equivalent norm is given by $\|\mathbf{A}\| = \sqrt{(A_1)^2 g^{11} + (A_2)^2 g^{22} + (A_3)^2 g^{33}}$. At present, in the code, the coordinate system used is the toroidal one, *i.e.* the set of coordinates (x^1, x^2, x^3) is equal to (r, θ, φ) where r is the radial position, θ is the poloidal angle and φ the toroidal angle. Therefore, $g^{rr} = 1$, $g^{\theta\theta} = 1/r^2$, $g^{\varphi\varphi} = 1/R^2$ and $g^{ij} = 0$ for all $i \neq j$. $R(r, \theta) = R_0 + r \cos \theta$ with R_0 the major radius of the torus at the magnetic axis. The Jacobian J_x is then equal to $J_x = rR$.

2.2. Simplified magnetic configuration and current

Consistently with the chosen coordinate system, the magnetic topology is fixed and consists of concentric toroidal magnetic surfaces with circular poloidal cross-sections. Indeed, the magnetic field $\mathbf{B}$ is defined as $\mathbf{B} = (B_0 R_0 / R) [\zeta(r) \mathbf{e}_\theta + \mathbf{e}_\varphi]$ with $\zeta(r) = r / (q R_0)$. B_0 corresponds to the magnetic field on the magnetic axis. The vectors $\mathbf{e}_\theta = r \nabla \theta$ and $\mathbf{e}_\varphi = R \nabla \varphi$ are the unit vectors in the poloidal and toroidal periodic directions. The safety factor profile $q(r)$ is defined by three parameters q_1 , q_2 , and q_3 as $q(r) = q_1 + q_2 \exp(q_3 \log(r/a))$. With this choice of angles, $\mathbf{B} \cdot \nabla \varphi / \mathbf{B} \cdot \nabla \theta = q(r) R_0 / R$, namely the local field line pitch depends not only on r but also on θ . The current is decoupled from the field and the magnetic field is assumed to satisfy the Ampère equation, but not the force balance equation. Then the Ampère equation leads to a current of the form $\mu_0 \mathbf{J} = \mu_0 J_T R \nabla \varphi$ with $\mu_0 J_T = \frac{B_0 R_0}{R} \frac{\zeta}{r} \left(1 + \frac{r}{\zeta} \frac{d\zeta}{dr} - \frac{r}{R} \cos \theta \right)$.

2.3. Full- f gyrokinetic Vlasov equation

Let us consider the gyro-center coordinate system $(\mathbf{x}_G, v_{G\parallel}, \mu)$ where $\mathbf{x}_G$ corresponds to 3D space coordinates. Let F_s be the particle distribution function of species s and $\bar{F}_s$ the one associated to the guiding-centers. The global gyrokinetic code GYSELA models for each species s , the time evolution of the guiding-center distribution function $\bar{F}_s$, with no separation between equilibrium and perturbation. The non-linear time evolution of $\bar{F}_s$ is governed by the 5D collisional gyrokinetic equation described by Brizard and Hahm [14]

$$\partial_t \bar{F}_s - [H, \bar{F}_s]_{\text{GC}} = \mathcal{R}hs(\bar{F}_s) \quad (1)$$

where H the Hamiltonian of the system is defined as $H = \frac{1}{2} m_s v_{G\parallel}^2 + \mu_s B + q_s \bar{U}$ and where $[\cdot, \cdot]_{\text{GC}}$ are the gyrokinetic Poisson brackets expressed as (see equation (150) in [14])

$$[X, Y]_{\text{GC}} = \frac{\mathbf{B}_s^*}{m_s B_{\parallel s}^*} \cdot (\nabla X \partial_{v_{G\parallel}} Y - \partial_{v_{G\parallel}} X \nabla Y) - \frac{\mathbf{b}}{q_s B_{\parallel s}^*} \cdot (\nabla X \times \nabla Y) \quad (2)$$

with $\mathbf{b} = \mathbf{B}/B$ the unit vector along the magnetic field line at the guiding-center position. The scalar $B_{\parallel s}^*$ corresponds to the volume element in guiding-center velocity space. It is simply $B_{\parallel s}^* = \mathbf{B}_s^* \cdot \mathbf{b}$ with $\mathbf{B}_s^*$ defined as

$$\mathbf{B}_s^* \equiv \mathbf{B} + \frac{m_s}{q_s} v_{G\parallel} \nabla \times \mathbf{b} \quad (3)$$

i.e.

$$B_{\parallel s}^* \equiv B + \frac{m_s}{q_s} v_{G\parallel} \mathbf{b} \cdot (\nabla \times \mathbf{b}) \quad (4)$$

In our case, the right hand side $\mathcal{R}hs$ of the previous boltzmann equation (1) is given by

$$\mathcal{R}hs(\bar{F}_s) = B_{\parallel s}^* (\mathcal{D}_r(\bar{F}_s) + \mathcal{K}(\bar{F}_s) + \mathcal{C}(\bar{F}_s) + \mathcal{S})$$

where $\mathcal{D}_r$ and $\mathcal{K}$ are respectively a diffusion term and a Krook operator applied on a radial buffer region (see section 2.10), $\mathcal{C}$ corresponds to a collision operator (see section 2.7) and $\mathcal{S}$ refers to source terms which are detailed in section 2.9.

Let us define the Poisson bracket as $[F, G] = \mathbf{b} \cdot (\nabla F \times \nabla G)$, and the parallel projection as $\nabla_{\parallel}^* F = \mathbf{b}_s^* \cdot \nabla F$. Let b_k represent the covariant components of the unitary magnetic field vector $\mathbf{b} = \mathbf{B}/\|\mathbf{B}\|$ and J_x the Jacobian in space of the system. Then it can be easily checked that $[F, G] = J_x^{-1} \epsilon^{ijk} \partial_i F \partial_j G b_k$ and $\mathbf{b}_s^* \cdot \nabla F = \mathbf{b}_s^{*i} \partial_i F$. By using this formalism, it can be proven that the Hamiltonian form (1) is equivalent to the following conservative form

$$B_{\parallel s}^* \frac{\partial \bar{F}_s}{\partial t} + \nabla \cdot \left(B_{\parallel s}^* \frac{d\mathbf{x}_G}{dt} \bar{F}_s \right) + \frac{\partial}{\partial v_{G\parallel}} \left(B_{\parallel s}^* \frac{dv_{G\parallel}}{dt} \bar{F}_s \right) = \mathcal{R}hs(\bar{F}_s) \quad (5)$$

where the evolution of the gyro-center coordinates of species s are described by:

$$\frac{dx_G^i}{dt} = v_{G\parallel} \mathbf{b}_s^* \cdot \nabla x_G^i + \mathbf{v}_{E \times B_s} \cdot \nabla x_G^i + \mathbf{v}_{D_s} \cdot \nabla x_G^i \quad (6)$$

$$m_s \frac{dv_{G\parallel}}{dt} = -\mu_s \mathbf{b}_s^* \cdot \nabla B - q_s \mathbf{b}_s^* \cdot \nabla \bar{U} + \frac{m_s v_{G\parallel}}{B} \mathbf{v}_{E \times B_s} \cdot \nabla B \quad (7)$$

where $\mathbf{b}_s^*$ is defined as¹

$$\mathbf{b}_s^* = \frac{\mathbf{B}}{B_{\parallel s}^*} + \frac{m_s v_{G\parallel}}{q_s B_{\parallel s}^* B} \nabla \times \mathbf{B} \quad (8)$$

The i -th contravariant coordinates of the ‘ $\mathbf{E} \times \mathbf{B}$ ’ drift are given by

$$\mathbf{v}_{E \times B_s} \cdot \nabla x_G^i = \mathbf{v}_{E \times B_s}^i = \frac{1}{B_{\parallel s}^*} [\bar{U}, x_G^i] \quad (9)$$

and the i -th contravariant coordinates of the ‘grad- B ’ and ‘curvature’ drifts reads (at low β limit)

$$\mathbf{v}_{D_s} \cdot \nabla x_G^i = \mathbf{v}_{D_s}^i = \left(\frac{m_s v_{G\parallel}^2 + \mu_s B}{q_s B_{\parallel s}^* B} \right) [B, x_G^i] \quad (10)$$

Here, $\bar{U} = \mathcal{J}_\mu U$ is the gyro-average of the fluctuating electrostatic potential U . This gyro-average operator $\mathcal{J}_\mu$ will be discussed in more detail in section 2.5. Besides, using the fact that the axi-symmetric equilibrium is determined by three of the ideal MagnetoHydroDynamic (MHD) equations $\nabla p = \mathbf{J} \times \mathbf{B}$, $\nabla \times \mathbf{B} = \mu_0 \mathbf{J}$ and $\nabla \cdot \mathbf{B} = 0$, the i -th contravariant coordinates of $\mathbf{b}_s^*$ appearing in equation (6) reads

$$\mathbf{b}_s^* \cdot \nabla x_G^i = \mathbf{b}_s^{*i} = \frac{\mathbf{B} \cdot \nabla x^i}{B_{\parallel s}^*} + \frac{m_s v_{G\parallel}}{q_s B_{\parallel s}^* B} \mu_0 \mathbf{J} \cdot \nabla x_G^i \quad (11)$$

2.4. Self-consistent coupling with the quasi-neutrality equation

The electron density n_e is supposed to follow an adiabatic (Boltzmann) response on a flux surface, namely $n_e(\mathbf{x}, t) = n_{e0}(r) \exp(e [U(\mathbf{x}, t) - \langle U \rangle_{\text{FS}}(r, t)] / T_e(r))$. $\langle U \rangle_{\text{FS}}$ represents the flux surface average of the electrostatic potential U , i.e

$$\langle U \rangle_{\text{FS}}(r) = \frac{\int U(r, \theta, \varphi) J_x \, d\theta \, d\varphi}{\int J_x \, d\theta \, d\varphi}$$

Let us define, for each species s , its concentration $c_s \equiv n_s/n_{e0}$ and its initial radial density profile n_{s0} . We also assume that $n_{e0} = \sum_s Z_s n_{s0}$ so that $\sum_s c_{s0} Z_s = 1$. Under these assumptions, the quasi-neutrality equation, self-consistently coupled to the gyrokinetic equation (5), reads

$$-\frac{1}{n_{e0}} \sum_s Z_s \nabla_\perp \cdot \left(\frac{n_{s0}}{B_0 \Omega_s} \nabla_\perp U \right) + e \left(\frac{U - \langle U \rangle_{\text{FS}}}{T_e} \right) = \frac{1}{n_{e0}} \sum_s Z_s (n_{G_s} - n_{G_s, \text{eq}}) \quad (12)$$

with $\Omega_s = q_s B_0 / m_s$. The gyro-center density n_{G_s} of species s is defined by $n_{G_s}(\mathbf{x}, t) = \int J_v d\mu dv_{G\parallel} \mathcal{J}_\mu \bar{F}_s(\mathbf{x}, \mathbf{v}, t)$ –where $J_v = (2\pi B_{\parallel s}^* / m_s)$ stands for the Jacobian in the velocity space– similarly $n_{G_s, \text{eq}}$ is the gyro-center density for $\bar{F}_s = \bar{F}_{s, \text{eq}}$. In practice, the right hand side of (12) the charge density of guiding-centers ρ , is computed as

$$\rho(\mathbf{x}, t) = \frac{1}{n_{e0}} \sum_s Z_s \int d\mu \mathcal{J}_\mu \left[\int J_v dv_{G\parallel} (\bar{F}_s(\mathbf{x}, \mathbf{v}, t) - \bar{F}_{s, \text{eq}}(r, \theta, v_{G\parallel})) \right] \quad (13)$$

¹Let us mention that in this definition $\mathbf{b}_s^*$ is not equal to $\mathbf{B}_s^* / B_{\parallel s}^*$ but equal to $\mathbf{b}_s^* = \frac{\mathbf{B}_s^*}{B_{\parallel s}^*} - \frac{m_s v_{G\parallel}}{q_s B_{\parallel s}^* B} \left(\mathbf{b} \times \frac{\nabla B}{B} \right)$.

To avoid the problem of the singularity in $1/r$, the problem is solved within a ring $r_{\min} \leq r \leq r_{\max}$ where $r_{\min} \geq 10^{-5}$. One difficulty with equation (12) is to deal with the flux surface average term $\langle U \rangle_{\text{FS}}$. This term is non linear in θ , because the space Jacobian depends on θ . Therefore, it does not allow one projecting simply in 2D Fourier space. To overcome this problem, solving (12) is performed in two steps. Let us define the differential operator $\mathcal{L} = -\frac{1}{n_{e0}} \sum_s Z_s \nabla_{\perp} \cdot \left(\frac{n_{s0}}{B_0 \Omega_s} \nabla_{\perp} \cdot \right)$ and $\langle \cdot \rangle_{\theta, \varphi}$ the average on θ and φ directions. Then, the first step consists in finding $\mathcal{U}$ solution of the following 1D radial differential equation,

$$\left(\mathcal{L} + \frac{e}{T_e} \right) \mathcal{U} = \rho - \langle \rho \rangle_{\theta, \varphi} \quad \text{with} \quad \mathcal{U}(r, \theta, \varphi) = U(r, \theta, \varphi) - \langle U \rangle_{\theta, \varphi}(r) \quad (14)$$

The second step consists in solving the following differential system

$$\mathcal{L} \langle U \rangle_{\text{FS}} = \langle \rho \rangle_{\theta, \varphi} + \left(\mathcal{L} + \frac{e}{T_e} \right) \langle \mathcal{U} \rangle_{\text{FS}} \quad (15)$$

In this equation, φ plays the role of a parameter. A Fourier projection is performed in the θ direction. In the radial direction, for both differential systems (14) and (15), finite differences are used. Finally, the electrostatic potential is reconstructed with the formula $U = \mathcal{U} - \langle U \rangle_{\text{FS}} + \langle \mathcal{U} \rangle_{\text{FS}}$. For both equation systems (14) and (15) Dirichlet boundary conditions are applied at $r = r_{\max}$. Concerning the inner boundary if $r_{\min} > 10^{-2}$ boundary conditions are Dirichlet otherwise Neumann. Let us notice that applying Dirichlet conditions to both matrix systems does not exactly correspond to apply Dirichlet boundary conditions to $\mathcal{U}$ but $\mathcal{U}|_{r_{\min}} = \langle U \rangle_{\theta, \varphi}|_{r_{\min}}$ and $\mathcal{U}|_{r_{\max}} = \langle U \rangle_{\theta, \varphi}|_{r_{\max}}$. On the other hand, for Neumann conditions as $r_{\min} \leq 10^{-2}$, $J_x(r_{\min}, \theta)$ can be assumed constant. This implies $\langle U \rangle_{\text{FS}}(r_{\min}) = (1/4\pi) \langle U \rangle_{\theta, \varphi}(r_{\min})$, then $\partial_r \langle U \rangle_{\theta, \varphi}(r_{\min}) = \partial_r \langle U \rangle_{\text{FS}}(r_{\min}) = 0$ and thus $\partial_r \mathcal{U}(r_{\min}) = 0$. The fact that this solution is equivalent to solving directly equation (12) and the detailed numerical scheme associated are explained in Appendix A.

2.5. Gyro-average operator

The gyro-radius ρ_s is transverse to $\mathbf{b} = \mathbf{B}/B$ and depends on the gyrophase angle φ_c , i.e $\rho_s = (\mathbf{b} \times \mathbf{v})/\Omega_s = \rho_s (\cos \varphi_c \mathbf{e}_{\perp 1} + \sin \varphi_c \mathbf{e}_{\perp 2})$. Here $\mathbf{e}_{\perp 1}$ and $\mathbf{e}_{\perp 2}$ are the unit vectors of a basis in the plane perpendicular to the magnetic field direction $\mathbf{b}$. Let $\mathbf{x}_G$ be the guiding-center radial coordinate and $\mathbf{x}$ the position of the particle in real space. These two quantities differ by a Larmor radius ρ_s as $\mathbf{x} = \mathbf{x}_G + \rho_s$. The gyro-average $\bar{g}$ of any function g depending on the spatial coordinates corresponds to the following operation

$$\bar{g}(\mathbf{x}_G, v_{\perp}) = \oint_0^{2\pi} \frac{d\varphi_c}{2\pi} g(\mathbf{x}) = \left\{ \oint_0^{2\pi} \frac{d\varphi_c}{2\pi} \exp(\rho_s \cdot \nabla) \right\} g(\mathbf{x}_G) \quad (16)$$

The operator $\exp(\rho_s \cdot \nabla)$ corresponds to the change of coordinates $(\mathbf{x}, \mathbf{p}) \rightarrow (\mathbf{x}_G, \mathbf{p}_G)$. This gyro-average process consists in computing an average on the Larmor circle. It weakens fluctuations that develop at sub-Larmor scales. Introducing $\hat{g}(\mathbf{k})$ the Fourier transform of g , with $\mathbf{k}$ the wave vector, it is simple to prove that the gyro-average operation reads

$$\bar{g}(\mathbf{x}_G, v_{\perp}) = \int_{-\infty}^{+\infty} \frac{d^3 \mathbf{k}}{(2\pi)^3} J_0(k_{\perp} \rho_s) \hat{g}(\mathbf{k}) e^{i\mathbf{k} \cdot \mathbf{x}_G} \quad (17)$$

where $k_{\perp}$ is the norm of the transverse component of the wave vector $\mathbf{k}_{\perp} = \mathbf{k} - (\mathbf{b} \cdot \mathbf{k}) \mathbf{b}$, and J_0 is the Bessel function² of first order. Considering the expression (17), in Fourier space the gyro-average reduces to the multiplication by the Bessel function of argument $k_{\perp} \rho_s$. This operation is straightforward in simple geometry with periodic boundary conditions, such as in local codes. Conversely, in the case of global codes,

²The Bessel function of the first kind are defined as $J_n(z) = \frac{i^{-n}}{\pi} \int_0^{\pi} \exp(iz \cos \theta) \cos(n\theta) d\theta$.

the use of Fourier transform is not applicable for two main reasons: (i) radial boundary conditions are non periodic, and (ii) the radial dependence of the Larmor radius has to be accounted for. Several approaches have been developed to overcome this difficulty.

2.5.1. A Padé approximation for the gyro-average operator

The first one, currently used in the code, consists in simplifying the treatment of the gyro-average operator by approximating the Bessel function with a Padé expansion $J_{\text{Padé}}(k_{\perp}\rho_s) = 1/[1 + (k_{\perp}\rho_s)^2/4]$ (e.g. see [87]). Such a Padé representation requires the solving of a Laplacian operator $\nabla_{\perp}^2$ in real space. Indeed, using the equivalence $i\mathbf{k}_{\perp} \leftrightarrow \nabla_{\perp}$, the gyro-average operation of any g function is defined such that each m Fourier mode of $\bar{g}$ is solution of the equation

$$\left[1 - \frac{1}{2\Omega_s^2} \frac{B_0}{m_s} \mu_s \left(\frac{\partial^2}{\partial r^2} + \frac{1}{r} \frac{\partial}{\partial r} - \frac{m^2}{r^2} \right)\right] \bar{g}^m(r, \varphi) = g^m(r, \varphi) \quad (18)$$

where at first approximation $B(r, \theta)$ has been replaced by B_0 to be consistent with the quasi-neutrality equation. In this differential equation, first and second derivatives are computed using a Taylor expansion of second order leading to a tridiagonal matrix system. This Padé approximation gives the correct limit in the large wavelength limit $k_{\perp}\rho_c \ll 1$, while keeping $J_{\text{Padé}}$ finite in the opposite limit $k_{\perp}\rho_s \rightarrow \infty$. The drawback is a filtering of small scales: in the limit of large arguments $x \rightarrow \infty$, $J_{\text{Padé}}(x) \rightarrow 4/x^2$, whereas $J_0 \rightarrow (2/\pi x)^{1/2} \cos(x - \pi/4)$.

2.5.2. Integration on gyro-circles using Hermite interpolation

A second widespread method for this gyro-averaging process is to use a quadrature formula. In this context, the integral over the gyro-ring is usually approximated by a sum over four points or more on the gyro-ring [70]. This is rigorously equivalent to considering the Taylor expansion of the Bessel function at order two in the small argument limit, namely $J_0(k_{\perp}\rho_s) \simeq 1 - (k_{\perp}\rho_s)^2/4$, and equivalent to computing the transverse Laplacian at second order using finite differences. This method has been extended to achieve accuracy for large Larmor radius [50], i.e the number of points (starting with four) is linearly increased with the gyro-radius to guarantee the same number of points per arclength on the gyro-ring. In this approach –used e.g. in [56] and [60]– the points that are equidistantly distributed over the ring are rotated for each particle (or marker) by a random angle calculated every time step. This is performed on a finite element formalism and enables therefore high order accuracy by keeping the matricial formulation. In [26] the influence of the interpolation operator (which is of great importance when the quadrature points do not coincide with the grid points) has been studied and it is shown that the cubic splines are appropriate. The direct integration on gyro-circles proposed in [26] has been recently generalized to arbitrary coordinates [92] and implemented in the code. For the distribution function the gyro-average operator is applied on $\bar{F}_s - \bar{F}_{s,\text{eq}}$ to deal with values close to 0 at the domain boundaries. Two interpolations –cubic splines and Hermite polynomial– have been tested on analytical cases and basic gyrokinetic simulations with a 4D drift-kinetic model, one Larmor radius and the standard linear Cyclone benchmark case (see [92] for more details and comparison with Padé approximation). Both appear to give the same results. However, the Hermite interpolation is slightly faster and its local character is more favourable for parallelization. The number of points per circle is an input parameter comprised between 4 and 16. There is no adaptive number of points depending on the Larmor radius value because the CPU time is in fact determined by the maximum number. As shown in [92] the method converges with the number of points, so there is no interest in decreasing the number of points at small radius. It is shown in [83] that 16 points is a good compromise between accuracy and CPU time consumption (2 times slower than the previous Padé approximation due to its higher algorithmic complexity). All numerical results presented in the following are performed with the Padé approximation.

2.6. Initial and boundary conditions

Boundary conditions are periodic in θ and φ directions. Concerning the radial direction, GYSELA is a *global* code, i.e it considers a large fraction of the plasma radius. This is in contrast to flux-tube codes which

focus on a small volume around magnetic field lines by proceeding from a scale separation assumption, the fluctuation scale length being smaller than that of the equilibrium. In such codes, periodicity is assumed along the radial direction. Conversely, global codes as GYSELA face the delicate problem of radial boundary conditions. Non-axisymmetric fluctuations of the electric potential and of the distribution function –*i.e.* $(m, n) \neq (0, 0)$ modes, with m and n the poloidal and toroidal wave numbers– are forced to zero at both radial boundaries of the simulated domain. As far as the axisymmetric component is concerned, the value of the potential is prescribed at the outer boundary, while the radial electric field is set to zero at the inner boundary. Such boundary conditions may be subject to numerical instabilities if turbulence reaches the radial borders. Buffer regions with overdamped fluctuations (see section 2.10) have been added to avoid possible numerical instabilities at radial boundaries.

Initial conditions consist of an equilibrium distribution function $\bar{F}_{s,\text{eq}}$ perturbed by a sum of accessible (m, n) Fourier modes (m and n being the poloidal and toroidal wave numbers, respectively). That means, $\bar{F}_s = \bar{F}_{s,\text{eq}} + \delta\bar{F}_s$ where the perturbation part $\delta\bar{F}_s$ reads $\delta\bar{F}_s = \bar{F}_{s,\text{eq}} g(r) h(v_{G\parallel}) \delta p(\theta, \varphi)$ with $\delta p(\theta, \varphi) = \sum_{m,n} \epsilon \cos(m\theta + n\varphi + \delta_{mn})$ where the amplitude ϵ is fixed and the phases δ_{mn} have arbitrary values. The radial function $g(r)$ (resp. $h(v_{G\parallel})$) has a polynomial dependence and vanishes at both radial (resp. $v_{G\parallel}$) boundaries. Concerning the initialization of the equilibrium distribution function two choices are possible: (i) the first one is a local conventional Maxwellian, (ii) and the second one is a *canonical* Maxwellian, *i.e.* depends on the motion invariants. The Maxwellian distribution function is defined as

$$\bar{F}_{s,\text{eq}}(r, E) = n_{s_0}(r) \times [2\pi T_s(r)/m_s]^{-\frac{3}{2}} \exp\left(-\frac{E}{T_s(r)}\right) \quad (19)$$

where E stands for the kinetic energy $m_s v_{G\parallel}^2/2 + \mu B(r, \theta)$ which is the second invariant of the system (at vanishing electric potential). The initial radial profiles of the ion temperature and density (respectively $T_s(r)$ and $n_{s_0}(r)$) are deduced by numerical integration of their gradient profiles given by the two parameters κ and Δr : $d \log T_s(r)/dr = -\kappa_{T_s} \cosh^{-2}((r - r_p)/\Delta r_{T_s})$ with r_p corresponding to the middle of the radial box. $\bar{F}_{s,\text{eq}}$ is constant on a magnetic surface labelled by the radial coordinate r . As shown in [32], such initial state does not constitute an equilibrium of the system solved by GYSELA at vanishing electric field. Indeed, a stationary equilibrium of the collisionless equations of the code must depend on the three motion invariants, namely the adiabatic invariant μ , the total energy E and the toroidal kinetic momentum $P_\varphi = q_s \psi + m_s R v_\varphi$ with ψ the poloidal flux and v_φ the toroidal fluid velocity. In GYSELA, a convenient choice for this equilibrium is provided by the canonical Maxwellian (19) in which the radial coordinate r is replaced –as proposed in [4]– by an effective radial coordinate $\bar{r}$, with the dimension of a length, derived from P_φ

$$\bar{r} = r_p - \frac{q_p}{r_p} \left[\psi(r) - \psi(r_p) \right] - \frac{m_s q_p}{e B_0 r_p} \left[R v_{G\parallel} - R_0 v_{G\parallel} \right] \quad (20)$$

where $\psi(r) = -B_0 \int_0^r r' / q \, dr'$. The last term $v_{G\parallel}$ in (20) is defined as,

$$v_{G\parallel} = \text{sign}(v_{G\parallel}) \sqrt{2/m_s} \sqrt{E - \mu B_{\text{max}}} \mathcal{H}(E - \mu B_{\text{max}})$$

with $\mathcal{H}$ the Heaviside function and B_{max} the maximum of the magnetic field on the whole simulation box. It has been chosen to minimize parallel flows. With this expression the difference between $\bar{r}$ and r is of order ρ_* , the ratio of ρ_s the Larmor radius of species s on the minor radius a . In the case of a decaying turbulence it appears that it is important to choose $\bar{F}_{s,\text{eq}}$ to be a function of the motion invariants, especially for studying zonal flows. Indeed, it had been observed in [56, 4] that breaking this rule leads to the development of large scale steady flows, which can prevent the onset of turbulence. This phenomena has also been observed in GYSELA simulations where a study of the difference between both equilibrium initialization has been performed in details [32, 34]. In [34] it has also been shown that when the system is driven by an external source of free energy, the choice of a canonical equilibrium is less crucial. Indeed, the turbulence onset is only delayed and its ultimate nature is unaltered; namely characterized by the same level of parallel and axisymmetric flows, the same level of transport and the same correlation times and lengths.

2.7. Simplified collision operator recovering neoclassical effects

Although fusion plasmas are weakly collisional, so that a kinetic approach is mandatory, the effect of collisions cannot be fully neglected. Indeed collisional transport plays an important role in regions where turbulent transport is low, such as transport barriers. Even more importantly, collisional friction damps low frequency flows, and hence controls the equilibrium radial electric field. Collisions also regularize fine structures in velocity space. In tokamak configurations, collisions depend on Coulomb binary interactions between particles (see Landau [62] for calculations). Then it is shown in [51] that the collision operator applied to the distribution functions can take the classical Fokker-Planck operator form:

$$\mathcal{C}_{ss'}(F_s, F_{s'}) = \frac{\partial}{\partial \mathbf{v}} \cdot \left\{ \left(\mathcal{D}(F_{s'}) \cdot \frac{\partial}{\partial \mathbf{v}} + \mathcal{V}(F_{s'}) \right) F_s \right\} \quad (21)$$

where $\mathcal{V}$ is a dynamical friction term and $\mathcal{D}$ is a diffusion term of the order of $\mathcal{D} \sim v_{T_s}^2 \nu_{ss'}$ with the collision frequency $\nu_{ss'}$ being defined as

$$\nu_{ss'} \sim \frac{\mathcal{D}}{v_{T_s}^2} \sim \frac{n_{s'}}{m_s} \left(\frac{1}{m_s} + \frac{1}{m_{s'}} \right) \frac{q_s^2 q_{s'}^2 \log \Lambda}{(1 + \epsilon_0)^2} (v_{T_s}^2 + v_{T_{s'}}^2)^{-3/2} \quad (22)$$

where $n_{s'}$ is the density of species s' , q_s (resp. $q_{s'}$) is the particle charge of species s (resp. s'), m_s is the particle mass of species s , ϵ_0 the permittivity of free space, $\log \Lambda \approx 17$ the Coulomb logarithm, v_{T_s} is the thermal velocity of species s and $v_{T_{>}}$ is the maximum value between the two thermal velocities v_{T_s} and $v_{T_{s'}}$. Let us first underline that the frequency $\nu_{ss'}$ is different from $\nu_{s's}$. Besides, let us compare the collision frequencies for the different following particle populations: (i) ion-ion collisions: $\nu_{ii} \propto Z^4 n_i / \sqrt{m_i} T_i^{3/2}$, (ii) electron-electron collisions: $\nu_{ee} \propto n_e / \sqrt{m_e} T_e^{3/2}$, (iii) electron-ion collisions: $\nu_{ei} \approx Z \nu_{ee}$ and (iv) ion-electron collisions: $\nu_{ie} \approx \frac{m_e}{m_i} Z^2 \nu_{ee} \ll \nu_{ii} \ll \nu_{ee}$. Therefore, the ion-electron collisions can be neglected. At the moment in the code, the electrons are considered adiabatic, therefore only the ion-ion collisions and impurity-ion collisions (see section 2.8.1) are taken into account. The full gyro-averaged and linearized Landau operator has been derived in [104] but such a full Coulomb collisional operator is difficult to implement in GYSELA without severe loss of parallelisation efficiency (see section 3.2 for more details). Therefore, the operator implemented in the code, is a simplified version of the Lenard-Bernstein operator [71]. This simplified version has been derived in [43] where it is especially shown that it recovers the exact neoclassical transport in the banana and plateau regimes³ (see Helander's book [53] p.149 for complete neoclassical transport regime description). This generic energy and momentum-conserving collision operator has been implemented and successfully tested in the code (see [2], [31]). A new version also valid for the Pfirsch-Schüller regime is under development. The current collision operator is expressed as a simplified Lorentz-type operator where only the $v_{G\parallel}$ contribution is taken into account:

$$\mathcal{C}_s(\bar{F}_s) = \frac{1}{B_{\parallel s}^*} \frac{\partial}{\partial v_{G\parallel}} \left\{ B_{\parallel s}^* \mathcal{D}_{\parallel s} \bar{F}_{M_s} \frac{\partial}{\partial v_{G\parallel}} \left(\frac{\bar{F}_s}{\bar{F}_{M_s}} \right) \right\} \quad (23)$$

where $\bar{F}_{M_s}$ is a shifted Maxwellian distribution –chosen such that $\mathcal{C}_s(\bar{F}_{M_s}) = 0$ – defined as:

$$\bar{F}_{M_s} = \frac{n_{s0}}{(2\pi T_{s,\text{coll}}/m_s)^{3/2}} \exp \left(-m_s \frac{(v_{G\parallel} - V_{\parallel s,\text{coll}})^2}{2T_{s,\text{coll}}} - \frac{\mu_s B}{T_{s,\text{coll}}} \right) \quad (24)$$

³The physics of neoclassical transport depends on the collisionality ν_* . If the collisionality is low, $\nu_* < 1$, the particle orbits are completed by a typical thermal particle. In this so-called *banana* regime, trapped particles almost fully determine the transport coefficients. In the opposite limit, $\nu_* > \epsilon^{-3/2}$, the particle orbit is not fully completed because its motion is disturbed by collisions before. This high collisionality regime is called the *Pfirsch-Schüller* regime or fluid regime. In-between, the *plateau* regime is characterized by a weak dependance of the transport on the collisionality of the plasma.

The collision term $\mathcal{D}_{\parallel s}$ reads for each μ_s :

$$\mathcal{D}_{\parallel s}(r, v) = \mathcal{A}_s(r) \left(\frac{\Phi(v) - G(v)}{2v} \right) \quad \text{with} \quad \mathcal{A}_s(r) = 3 \frac{\sqrt{\pi}}{2} \frac{v_{T_{s,\text{coll}}}^3 \epsilon^{3/2}}{qR_0} \nu_{*s} \quad (25)$$

where $v(r, v_{G\parallel}) = \sqrt{E/T_{s,\text{coll}}}$. The scalar ν_{*s} , associated to the main s species, is a dimensionless ion-ion collisionality parameter depending on the ion-ion collision frequency ν_{ss} such that:

$$\nu_{*s} = \frac{qR_0}{v_{T_{s,\text{coll}}} \epsilon^{3/2}} \nu_{ss} \quad \text{with} \quad \nu_{ss} = \frac{4\sqrt{\pi}}{3} \frac{n_s e^4 \log \Lambda}{(4\pi\epsilon_0)^2 m_s^2 v_{T_{s,\text{coll}}}^3} \quad (26)$$

where $v_{T_{s,\text{coll}}}$ denotes the initial thermal velocity $v_{T_{s,\text{coll}}} = (T_{s,\text{coll}}/m_s)^{1/2}$, q the safety factor, $R = R_0 + r \cos \theta$ the major radius, $\epsilon = r/R_0$ the inverse aspect ratio and n_s the density of ion species s . The explicit expression (25) also involves the error function Φ and the Chandrasekhar function G defined as

$$G(v) = \frac{\Phi(v) - v\Phi'(v)}{2v^2} \quad \text{with} \quad \Phi(v) = \frac{2}{\sqrt{\pi}} \int_0^v e^{-x^2} dx \quad \text{and} \quad \Phi'(v) = \frac{2}{\sqrt{\pi}} e^{-v^2} \quad (27)$$

Considering that the ν_{*s} scalar which is given (as input data) in the code corresponds to the main species, the collision frequencies $\nu_{s's'}$ for minority species s' are deduced from this value as

$$\nu_{s's'} = \frac{v_{T_{s',\text{coll}}} \epsilon^{3/2}}{qR_0} \nu_{*s'} \quad \text{with} \quad \nu_{*s'} = \left(\frac{n_{s'}}{n_s} \right) \left(\frac{Z_{s'}}{Z_s} \right)^4 \left(\frac{T_{s,\text{coll}}}{T_{s',\text{coll}}} \right)^2 \nu_{*s} \quad (28)$$

Let us express for each μ_s , the operator $\mathcal{V}_{\parallel s}$ as

$$\mathcal{V}_{\parallel s}(r, v) = - \frac{(v_{G\parallel} - V_{\parallel s,\text{coll}})}{v_{T_{s,\text{coll}}}^2} \mathcal{D}_{\parallel s}(r, v) \quad \text{with} \quad v_{T_{s,\text{coll}}} = \sqrt{T_{s,\text{coll}}/m_s} \quad (29)$$

then the collision operator (23) can be expressed with a more classical Fokker-Planck structure as:

$$\mathcal{C}_s(\bar{F}_s) = \frac{1}{B_{\parallel s}^*} \frac{\partial}{\partial v_{G\parallel}} \left[B_{\parallel s}^* \left(\mathcal{D}_{\parallel s} \frac{\partial \bar{F}_s}{\partial v_{G\parallel}} - \mathcal{V}_{\parallel s} \bar{F}_s \right) \right] \quad (30)$$

where the operators $\mathcal{D}_{\parallel s}$ and $\mathcal{V}_{\parallel s}$ respectively model a diffusion and a drag in the parallel velocity direction. The conservation properties of parallel momentum and energy are ensured by constraining $\mathcal{D}_{\parallel s}$ to depend on μ_s only and defining the local fluid velocity $V_{\parallel s,\text{coll}}$ and ion temperature $T_{s,\text{coll}}$ as follows (see Appendix [AppendixB](#) for more details)

$$V_{\parallel s,\text{coll}} = P^{-1} \left[\left\langle \frac{m_s}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s} v_{G\parallel}) \right\rangle \langle m_s \mathcal{D}_{\parallel s} v_{G\parallel} \rangle - \left\langle \frac{1}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s}) \right\rangle \langle m_s^2 \mathcal{D}_{\parallel s} v_{G\parallel}^2 \rangle \right] \quad (31)$$

$$T_{s,\text{coll}} = P^{-1} \left[\langle m_s \mathcal{D}_{\parallel s} \rangle \langle m_s^2 \mathcal{D}_{\parallel s} v_{G\parallel}^2 \rangle - \langle m_s \mathcal{D}_{\parallel s} v_{G\parallel} \rangle \langle m_s^2 \mathcal{D}_{\parallel s} v_{G\parallel} \rangle \right] \quad (32)$$

where

$$P = \langle m_s \mathcal{D}_{\parallel s} \rangle \left\langle \frac{m_s}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s} v_{G\parallel}) \right\rangle - \langle m_s^2 \mathcal{D}_{\parallel s} v_{G\parallel} \rangle \left\langle \frac{1}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s}) \right\rangle \quad (33)$$

with the brackets $\langle \cdot \rangle$ corresponding to the velocity space integral $\langle \cdot \rangle = \int \cdot J_v d\mu_s dv_{G\parallel} F_s$. The impact of the collisions on the evolution of the distribution function $\bar{F}_s$ is taken into account by stepping the evolution of the distribution function $\partial_t \bar{F}_s = C_s(\bar{F}_s)$ with a Crank-Nicolson scheme. This collision operator forces the system to relax towards the Maxwellian distribution function, calculated from the instantaneous and local parallel flow $V_{\parallel s, \text{coll}}$ and the isotropic temperature $T_{s, \text{coll}} \sim T_{s, \text{coll} \parallel}$.

2.8. Collisional transfer between two species

The inter-species operator currently implemented in the code is highly simplified. It only ensures the moment transfers and energy transfers between species. A more complete version, satisfying neoclassical results for impurity transport, is under development.

2.8.1. Collisional energy transfer

The energy exchange between two species is approached by the following reduced collision operator

$$\frac{d\bar{F}_s}{dt} = -\frac{\eta_{ss'}^{\mathcal{E}}}{(2\pi\bar{T}/m_s)^{3/2}} \frac{\Delta T_{ss'}}{\bar{T}} \left(\frac{\mathcal{E}_{ss'}}{\bar{T}} - \frac{3}{2} \right) \exp\left(-\frac{\mathcal{E}_{ss'}}{\bar{T}}\right) \equiv C_{ss'}^{\mathcal{E}}(\mathcal{E}_{ss'}) \quad (34)$$

$$\frac{d\bar{F}_{s'}}{dt} = -\frac{\eta_{s's}^{\mathcal{E}}}{(2\pi\bar{T}/m_{s'})^{3/2}} \frac{\Delta T_{s's}}{\bar{T}} \left(\frac{\mathcal{E}_{s's}}{\bar{T}} - \frac{3}{2} \right) \exp\left(-\frac{\mathcal{E}_{s's}}{\bar{T}}\right) \equiv C_{s's}^{\mathcal{E}}(\mathcal{E}_{s's}) \quad (35)$$

where d/dt stands for the phase space Lagrangian derivative and the following definitions have been adopted

$$\bar{T} = \frac{T_s + T_{s'}}{2} \quad ; \quad \Delta T_{ss'} = T_s - T_{s'} = -\Delta T_{s's} \quad (36)$$

$$\bar{V} = \frac{V_{\parallel s} + V_{\parallel s'}}{2} \quad ; \quad \Delta V_{ss'} = V_{\parallel s} - V_{\parallel s'} \quad (37)$$

$$\mathcal{E}_{ss'} = \frac{m_s (v_{G\parallel} - \bar{V})^2}{2} + \mu_s B \quad ; \quad \mathcal{E}_{s's} = \frac{m_{s'} (v_{G\parallel} - \bar{V})^2}{2} + \mu_{s'} B \quad (38)$$

The temperatures and fluid velocities which enter these definitions are flux surface averaged, *i.e.* by considering $\langle \cdot \rangle_{\text{FS}} = \int \cdot J_x d\theta d\varphi / \int J_x d\theta d\varphi$. Then velocities correspond to $V_{\parallel s}(r) = \langle \int v_{G\parallel} \bar{F}_s d^3\mathbf{v} \rangle_{\text{FS}} / N_s(r)$ with density $N_s(r) = \langle \int \bar{F}_s d^3\mathbf{v} \rangle_{\text{FS}}$, the volume element being defined as $d^3\mathbf{v} = J_v dv_{G\parallel} d\mu_s$. Temperatures are defined as $T_s(r) = 2/3 \langle \int \mathcal{E}_s \bar{F}_s d^3\mathbf{v} \rangle_{\text{FS}} / N_s(r)$ where $\mathcal{E}_s = (m_s/2) (v_{G\parallel} - V_{\parallel s})^2 + \mu_s B$. The parameter $\eta_{ss'}^{\mathcal{E}}$ has been designed so that particles and parallel momentum are conserved

$$\eta_{ss'}^{\mathcal{E}} = \frac{8 \varepsilon^{3/2}}{3\sqrt{\pi}} \frac{m_s}{m_{s'}} \left(\frac{Z_{s'}}{Z_s} \right)^2 \frac{n_{s'} \sqrt{T_{s, \text{coll}}/m_s}}{qR_0} \left(1 + \frac{v_{T_{s'}}^2}{v_{T_s}^2} \right)^{-3/2} \nu_{*s} \quad (39)$$

where the normalized ion-ion collision frequency ν_{*s} is given by eq.(26) (for detailed calculation see Appendix [Appendix C.1](#)). Numerically, equation (34), is solved as $\bar{F}_s(t - \Delta t) = \bar{F}_s(t) + \Delta t C_{ss'}^{\mathcal{E}}(\mathcal{E}_{ss'}, t)$.

2.8.2. Parallel momentum transfer

Parallel momentum exchange between two species can be modelled by the following approximate collision operator:

$$\frac{d\bar{F}_s}{dt} = -\eta_{ss'}^{v_{G\parallel}} \Delta V_{ss'} v_{G\parallel} \exp\left(-\frac{E_s}{\bar{T}}\right) \equiv C_{ss'}^{v_{G\parallel}}(E_s) \quad \text{with } E_s = \frac{1}{2} m_s v_{G\parallel}^2 + \mu_s B \quad (40)$$

$$\frac{d\bar{F}_{s'}}{dt} = -\eta_{s's}^{v_{G\parallel}} \Delta V_{s's} v_{G\parallel} \exp\left(-\frac{E_{s'}}{\bar{T}}\right) \equiv C_{s's}^{v_{G\parallel}}(E_{s'}) \quad (41)$$

where $\bar{T}$ is defined by equation (36). $\Delta V_{ss'}$ is equivalent to eq.(37). The parameters $\eta_{s's}^{v_{G\parallel}}$ and $\eta_{ss'}^{v_{G\parallel}}$ are designed such that only parallel momentum exchanges are induced by this operator, thus leading to

$$\eta_{ss'}^{v_{G\parallel}} = n_{s'} m_s^2 \left(\frac{\varepsilon}{2\pi} \right)^{3/2} \frac{1}{qR_0} \frac{\sqrt{T_{s,\text{coll}}}}{\bar{T}^{5/2}} \left(\frac{Z_{s'}}{Z_s} \right)^2 \frac{v_{T_s}}{v_{T_{>}}} \nu_{*s} \quad \text{and} \quad \eta_{s's}^{v_{G\parallel}} = \left(\frac{m_{s'}}{m_s} \right)^{3/2} \eta_{ss'}^{v_{G\parallel}} \quad (42)$$

where the velocity $v_{T_{>}}$ corresponds to the maximum value between v_{T_s} and $v_{T_{s'}}$ (for more details see Appendix [AppendixC.2](#)).

2.9. Flux-driven code with source terms

2.9.1. Sources of heating, momentum and vorticity

As introduced previously GYSELA is a *full-f* code, namely the back reaction of turbulent transport is accounted for in the time evolution of the equilibrium. In such a framework, the turbulence regime is evanescent if no free energy is injected in the system. Indeed, turbulent transport results in the flattening of the temperature profile, which would ultimately reach marginal stability in the absence of any forcing. A heat source is mandatory in view of exploring the long time –on energy confinement times– behaviour of turbulence and transport. In GYSELA, the possibility to drive the system by a prescribed source was added in 2009 [84]. This source consists of the sum of the product of Hermite and Laguerre polynomials in $v_{G\parallel}$ and μ_s , respectively, in the spirit of the pioneering work by Darmet *et al* [29]. It is versatile enough to allow for separate injection of heat, parallel momentum and vorticity. Such a versatility imposes serious constraints on the expression of the source in phase space. The separation between these three kinds of sources is achieved using projections onto the bases of orthogonal Hermite polynomials in $v_{G\parallel}$ and orthogonal Laguerre polynomials in μ_s . The retained expression for the source is the following (see Appendix [AppendixD](#) for detailed description)

$$\frac{d\bar{F}_s}{dt} = S_{\text{heat}}(r, \theta, v_{G\parallel}, \mu_s) + S_{\text{moment}}(r, \theta, v_{G\parallel}, \mu_s) + S_{\text{vorticity}}(r, \theta, v_{G\parallel}, \mu_s) \quad (43)$$

where the pure heating source is defined as:

$$S_{\text{heat}} = \left[\bar{v}_{G\parallel s}^2 - \frac{1}{2} - \frac{J_{\parallel B}}{2 - J_{\parallel B}^2} (2 - \bar{\mu}_s) (2\bar{v}_{G\parallel s} - J_{\parallel B}) \right] \frac{2S_0^\mathcal{E}}{(2\pi T_{s,\text{srce}}/m_s)^{3/2} T_{s,\text{srce}}} S_r^\mathcal{E} e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s} \quad (44)$$

with $\bar{\mu}_s = \mu_s B / T_{s,\text{srce}}$, $\bar{v}_{G\parallel s} = v_{G\parallel} / \sqrt{2T_{s,\text{srce}}/m_s}$, $J_{\parallel B} \equiv \sqrt{2m_s T_{s,\text{srce}}}/(q_s B^2) \mu_0 J_{\parallel}$ and $\alpha = (m_s/q_s^2) \times T_{s,\text{srce}}/(2B^2)$. The moment source S_{moment} is a pure momentum source expressed as

$$S_{\text{moment}} = [2\bar{v}_{G\parallel s}(2 - \bar{\mu}_s) - J_{\parallel B} (1 + 2\bar{v}_{G\parallel s} - \bar{\mu}_s)] \frac{S_0^{v_{G\parallel}}}{4\pi^{3/2} (T_{s,\text{srce}}/m_s)^2} S_r^{v_{G\parallel}} e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s} \quad (45)$$

while $S_{\text{vorticity}}$ is a pure source of vorticity

$$S_{\text{vorticity}} = - \left[2\bar{v}_{G\parallel s}^2 - \bar{\mu}_s \right] \frac{S_0^\Omega}{\alpha (2\pi T_{s,\text{srce}}/m_s)^{3/2}} S_r^\Omega e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s} \quad (46)$$

The prescribed radial envelopes $S_r^\mathcal{E}$, $S_r^{v_{G\parallel}}$ and S_r^Ω are the sum of two hyperbolic tangents under the constraint that the integral over the minus radius is normalized to 1. They are usually localized close to the inner boundary of the simulation domain [84]. The heat source is an isotropic source that takes particles at a vanishing velocity and accelerates them up to $v \approx 1.5v_{T_h}$. See schematic view with figure 1. As a remark,

there is another choice possible for the heat source S_{heat} in the code, which corresponds to

$$S_{\text{heat}} = \frac{S_0^\mathcal{E}}{3\sqrt{2} \left(\frac{\pi T_{s,\text{srce}}}{m_s} \right)^{3/2} T_{s,\text{srce}}} \left(\frac{E}{T_{s,\text{srce}}} - \frac{3}{2} \right) \exp \left(-\frac{E}{T_{s,\text{srce}}} \right) S_r^\mathcal{E} \quad (47)$$

There are two main advantages in dealing with a prescribed heat source: (i) the forcing of turbulence can mimic that in experiments, in contrast to simulations where the mean gradient is prescribed and (ii) the sum of the spatially and time (on the τ_E time scale) averaged turbulent and neoclassical heat fluxes must balance the prescribed driving flux. In this case, the response is the temperature gradient, which ultimately governs the internal energy and therefore the performance of the discharge. Flux driven simulations then allow investigating the impact of heating power on the energy confinement time [85]. The source of vorticity described before has been efficiently used to polarize the plasma [94] inducing the development of sheared electric fields in the turbulent core. The creation of ion transport barriers by these externally induced sheared $E \times B$ flows has been studied in details in [95, 94].

2.9.2. Energetic particle source

A source of energetic particles (EPs) has been also implemented in the code to study the interaction between EPs and turbulence. EPs are characterized by energies larger than the thermal energy. The excitation by EPs of the geodesic acoustic modes (GAMs) –corresponding to the oscillatory component of large scale $E \times B$ zonal flows– creates modes which are called energetic geodesic acoustic modes (EGAMs). For more details on the impact of EGAMs on turbulence see Zarzoso’s papers [105, 106]. In practice this source is coupled to the heating source, such that $d_t \bar{F}_s = S_{\text{heat}} + S_{\text{EP}}$ where S_{EP} is also built, as for the previous sources, by using projection onto the Laguerre and Hermite polynomial bases, with the constraint to inject only parallel energy. For symmetry reason, the energetic particle source, is built as

$$S_{\text{EP}}(r, \theta, v_{G\parallel}, \mu_s, t) = S_0^{\text{EP}}(t) S_r^{\text{EP}}(r) (S_+ + S_-) \quad \text{with} \quad S_-(\theta, v_{G\parallel}, \mu_s) = S_+(\theta, -v_{G\parallel}, \mu_s)$$

where S_0^{EP} is the source amplitude and S_r^{EP} the radial profile is normalized such that $\int r dr S_r^{\text{EP}} = 1$. After the same kind of calculation than for the previous sources, the final expression reads

$$S_\pm = \left[(\bar{v}_{G\parallel s} \pm \bar{v}_0)^2 - \frac{1}{2} - \mathcal{Q}_{\text{EP}}(2 - \bar{\mu}_s)(2(\bar{v}_{G\parallel s} \pm \bar{v}_0) - J_{\parallel B}) \right] e^{-(\bar{v}_{G\parallel s} \pm \bar{v}_0)^2} e^{-\bar{\mu}_s} \quad (48)$$

where $\bar{\mu}_s = \mu_s B(r, \theta) / T_{S\perp}$ and $\mathcal{Q}_{\text{EP}} = J_{\parallel B} / (2 - J_{\parallel B}^2(1 + 2\bar{v}_0^2))$ with $\bar{v}_0 = v_0 / \sqrt{2T_{S\parallel}}$ an arbitrary normalized velocity. The expressions of the parallel current $J_{\parallel B}$ and of the velocity $\bar{v}_{G\parallel s}$ are the same than in equation (44) where $T_{s,\text{srce}}$ is replaced by $T_{S\parallel}$. $T_{S\parallel}$ and $T_{S\perp}$ correspond to the normalized parallel and perpendicular temperatures of the energetic particle source. Each of the terms S_+ and S_- does not inject neither mass –which is essential because electrons are considered adiabatic in the code– nor vorticity. See figure 1 for a representative view of the source in parallel velocity direction. The source mimics the effects of two tangential neutral beam injectors, oriented in the co-and counter-current directions. It is localized around the mid position $r = 0.5(r_{\min} + r_{\max})$ and brings the distribution function out of the equilibrium by creating a positive slope in energy. As explained in [106] v_0 and $T_{S\parallel}$ are both critical parameters in view of exciting EGAMs.

2.9.3. Source of impurities

Finally, it is also possible to add a source of impurities s' of the form

$$S_{n_{s'}} = \frac{S_0^{n_{s'}} S_r}{\left(\frac{2\pi T_{s',\text{srce}}}{m_{s'}} \right)^{3/2}} \left(\frac{5}{2} - \frac{\mu_{s'} B}{T_{s',\text{srce}}} - \frac{m_{s'} v_{G\parallel}^2}{2T_{s',\text{srce}}} \right) \exp \left(-\frac{m_{s'} v_{G\parallel}^2}{2T_{s',\text{srce}}} - \frac{\mu_{s'} B}{T_{s',\text{srce}}} \right) \quad (49)$$

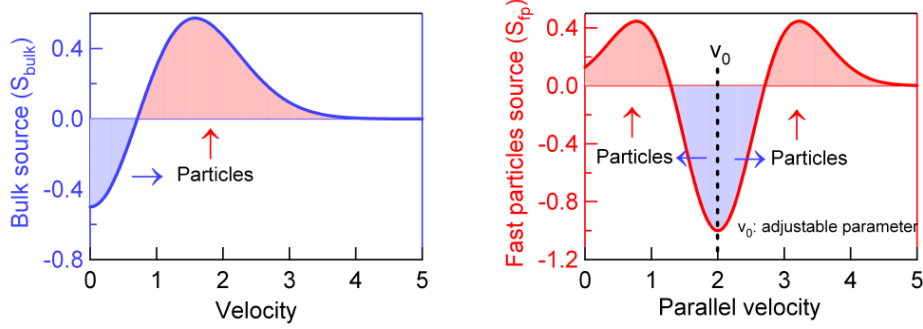


Figure 1: Schematic view of the energetic particle source as a function of the parallel velocity

Let us notice that at this source of matter is not a pure source, due to the fact that it also injects some amount of vorticity. This injection is balanced by a sink for the major species s , such that

$$Z_s \int S_{n_s} J_{v_s} dv_{G\parallel} d\mu_s + Z_{s'} \int J_{v_{s'}} S_{n_{s'}} dv_{G\parallel} d\mu_{s'} = 0$$

2.10. Diffusion terms in the buffer regions

Finally, to avoid strong gradients at the boundaries, radial diffusion and an artificial damping can be added in buffer regions. These buffer regions are defined at each side of the radial domain $r \in [r_{\min}, r_{\max}]$ as a sum of hyperbolic tangents:

$$H_{\text{buff}}(r) = 1 + \frac{1}{2} \left[\tanh \left(\frac{r - r_{\max} + B_L L_r}{B_S L_r} \right) - \tanh \left(\frac{r - r_{\min} - B_L L_r}{B_S L_r} \right) \right] \quad (50)$$

where L_r is the length of the radial domain. B_L and B_S are respectively the location and the stiffness of the buffer regions. The function H_{buff} plays the role of a mask which is equal to 1 in the buffer regions and 0 elsewhere. The diffusion term which is applied in the buffer regions is of the form

$$\mathcal{D}_r(\bar{F}_s) = \frac{1}{B_{\parallel s}^*} \left[\frac{1}{r} \frac{\partial}{\partial r} \left(r \chi(r) B_{\parallel s}^* \frac{\partial}{\partial r} \bar{F}_s \right) \right] \quad (51)$$

with $\chi(r) = \chi_0 H_{\text{buff}}(r)$ (χ_0 being the diffusion coefficient). The equation $\partial_t \bar{F}_s = \mathcal{D}_r(\bar{F}_s)$ is solved by using a Crank-Nicolson scheme (see Appendix E). An artificial damping term ν_0 is introduced in the buffer regions by defining a Krook operator $\mathcal{K}_r(\bar{F}_s) = -\nu(r)(\bar{F}_s - \bar{F}_{s,\text{eq}})$ with $\nu(r) = \nu_0 H_{\text{buff}}(r)$ and solving $\partial_t \bar{F}_s = \mathcal{K}_r(\bar{F}_s)$. Let Δt be the time step, then an analytic solution of the previous equation is given by $\bar{F}_s(t + \Delta t) = \bar{F}_{s,\text{eq}} + (\bar{F}_s(t) - \bar{F}_{s,\text{eq}}) \exp(-\nu(r)\Delta t)$. This mechanism restores the distribution function to its initial equilibrium state $\bar{F}_{s,\text{eq}}$, in the buffer regions, by slowly damping all the turbulent modes of the system. It also plays the role of a heat sink by effectively coupling the plasma with the outer thermal baths.

3. A highly parallel semi-lagrangian code

Solving the set of gyrokinetic equations (5)-(12) is very challenging. Indeed, it consists of (i) one evolution equation (5) of the distribution function for each ion species (namely a major species as Deuterium or Tritium and an impurity species as Helium, Carbon or Tungsten) in the 4D phase space parametrized by the adiabatic invariant μ ; (ii) the set of four coupled ordinary differential equations (ODE) for the trajectories (6)-(7); and (iii) 3D integro-differential equations for the field, namely the quasi-neutrality equation (12). This set of equations is nonlinear, the dominant quadratic nonlinearity being due to the $\mathbf{E} \times \mathbf{B}$ advection term. The quasi-neutrality equation is generally solved in most of the gyrokinetic codes by using Fourier projection

in all the periodic directions and finite differences or finite elements in the others. Concerning the global algorithm, the challenge consists in finding numerical schemes which preserve the first principles such as the conservation of Casimir invariants, the phase space volume and the total energy. Various numerical schemes, classifiable as PIC, Eulerian or semi-Lagrangian, have been explored until now. In the following, we focus on the semi-Lagrangian approach which is specific of the GYSELA code.

3.1. Specificity of the GYSELA code: the semi-Lagrangian scheme

Semi-Lagrangian (SL) schemes have been first used for the advection of vorticity in simplified models of large scale flows. It has gained maturity when the discretization approach was introduced in the relevant context of atmospheric flows. A comprehensive review of semi-Lagrangian methods in this meteorological context until 1990 is due to Staniforth [91]. It is also applied to geophysical fluid dynamics (cf. [39]). In magnetized plasma turbulence area, the SL method has been first applied to calculate a turbulence driven by passing ions in 2D (1D in space, 1D in velocity) [90] and trapped ions in 3D (2D in space, 1D in velocity) [30, 87]. This method was then extended to the 4D model (3D in space and $v_{\parallel}$ (with $\mu = 0$)) of Ion Temperature Gradient (ITG) driven turbulence in cylindrical geometry with the development of the GYSELA code (for GYrokinetic SEmi-LAgrangian code) [45] and the CYGNE code [17]. The 4D drift-kinetic slab-ITG version of the GYSELA code has shown good properties of energy conservation in non-linear regime [45] as well as accurate description of fine spatial scales [86]. In the CYGNE code the standard Taylor expansion is replaced by a Bürlisch-Stoer scheme (for the 2D advection) to increase the spatial accuracy and the logarithmic interpolation technique is used to ensure the positivity of the distribution function. Brunetti et al. [17] have shown that (i) the positivity can be preserved but at the cost of larger diffusion and (ii) that non-equidistant meshes in radial and parallel velocity directions are a key tool for obtaining accurate results. Due to the good conservation property obtained in 4D, the gyrokinetic 5D version of the GYSELA code [47] has been based on the same numerical scheme. The global algorithm for the new 5D gyrokinetic multi-ion species version of the code will be described in the following both in terms of numerics (section 3.1.3) and parallelisation (section 3.2). Advantages and drawbacks of such a 5D semi-lagrangian code will be discussed. The purpose of the semi-Lagrangian method is to take advantage of both the Lagrangian and Eulerian approaches, with an accurate description of the phase space, in particular regions where the density is low, as well as an enhanced numerical stability. It is based on the fact that the most precise way to solve convection (or advection) hyperbolic PDE is to use their characteristics along which the distribution function remains constant. This method was primarily developed by Cheng and Knorr [23]. It has been cast in more general framework of SL by Sonnendrücker in 1998 [90] and Nakamura in 1999 [76]. In this approach, the phase-space mesh grid is kept fixed in time (Eulerian method) and the Vlasov equation is integrated along the trajectories (Lagrangian method) using the invariance of the distribution function along the trajectories. The GYSELA code is based on this standard semi-Lagrangian approach [90]. This approach has been recently renamed *Backward semi-Lagrangian approach (BSL)* by its author to make the distinction with the emergence of new schemes: (i) *Forward semi-Lagrangian approach (FSL)* firstly proposed in 2009 [28] and (ii) *Conservative semi-Lagrangian approach (CSL)* proposed in [27, 13] through Parabolic Spline Method [107]. The main difference between BSL and FSL approaches is that the advection equations are solved backward in the first case and forward in the second one. They are both based on solving the advective form of Vlasov equation while the CSL methods deal with the conservative form of the Vlasov equation. FSL and CSL schemes have both been tested in GYSELA [67, 12] but the actual version of the code is the standard BSL approach as described in the following.

3.1.1. Backward Semi-Lagrangian (BSL) concept

Let us consider the 5D gyrokinetic Vlasov equation

$$B_{\parallel s}^* \frac{\partial \bar{F}_s}{\partial t} + \nabla \cdot \left(B_{\parallel s}^* \frac{d\mathbf{x}_G}{dt} \bar{F}_s \right) + \frac{\partial}{\partial v_{G\parallel}} \left(B_{\parallel s}^* \frac{dv_{G\parallel}}{dt} \bar{F}_s \right) = 0 \quad (52)$$

which corresponds to equation (5) without collisions and source terms. By using the incompressibility property of the gyrocenter flow in the 4D phase space, *i.e.* $\nabla \cdot \left(B_{\parallel s}^* d_t \mathbf{x}_G \right) + \partial_{v_{G\parallel}} \left(B_{\parallel s}^* d_t v_{G\parallel} \right) = 0$, the previous

conservative Vlasov equation (52) is equivalent to its advective form

$$\frac{\partial \bar{F}_s}{\partial t} + \frac{d\mathbf{x}_G}{dt} \cdot \nabla \bar{F}_s + \frac{dv_{G\parallel}}{dt} \frac{\partial \bar{F}_s}{\partial v_{G\parallel}} = 0 \quad (53)$$

Let $\mathbf{\Gamma} = (\mathbf{x}_G, v_{G\parallel})$ be a position vector in the phase space solution of the characteristic equations (6)-(7). Then equation (53) leads to $d_t \bar{F}_s(\mathbf{\Gamma}(t), t) = 0$. The semi-lagrangian method uses this invariance of the distribution function $\bar{F}_s$ along its characteristics. Let us consider the computational 4D domain $[x_{G\min}^1, x_{G\max}^1] \times [x_{G\min}^2, x_{G\max}^2] \times [x_{G\min}^3, x_{G\max}^3] \times [v_{G\parallel\min}, v_{G\parallel\max}]$ and the associated grid (fixed in time) defined by the finite set of mesh points $\mathbf{\Gamma}_{ijkl} = (x_{G_i}^1, x_{G_j}^2, x_{G_k}^3, v_{G\parallel_l})$ with $x_{G_q}^p = q\Delta x_G^p$ for all $q = 0, \dots, N_p$ with $p = 1, 2, 3$ and $v_{G\parallel_l} = l\Delta v_{G\parallel}$ for all $l = 0, N_{v\parallel}$. N_1, N_2, N_3 are the number of cells in each spatial directions and $N_{v\parallel}$ the number of cells in $v_{G\parallel}$ direction. Let us also assume that $\bar{F}_s$ is known at each point $\mathbf{\Gamma}_{ijkl}$ of this grid at time t_n . Therefore the distribution function can be computed at the next time $t_n + \Delta t$, with Δt the time step, on the same grid by using the invariance property $\bar{F}_s(\mathbf{\Gamma}_{ijkl}(t_n + \Delta t), t_n + \Delta t) = \bar{F}_s(\mathbf{\Gamma}(t_n, \mathbf{\Gamma}_{ijkl}, t_n + \Delta t), t_n)$ where $\mathbf{\Gamma}(t_n, \mathbf{\Gamma}_{ijkl}, t_n + \Delta t)$ corresponds to the solution of the characteristic at time step t_n which is equal to $\mathbf{\Gamma}_{ijkl}$ at time $t_n + \Delta t$. The method consists in (i) first, finding the foot of the characteristic at the time t_n : $\mathbf{\Gamma}(t_n, \mathbf{\Gamma}_{ijkl}, t_n + \Delta t)$ by solving backward in time the advection equations (6)-(7) and (ii) second, computing $\bar{F}_s(\mathbf{\Gamma}(t_n, \mathbf{\Gamma}_{ijkl}, t_n + \Delta t), t_n)$ by interpolation, using the fact that at this time t_n the distribution function is known over the whole fixed grid. Cubic spline interpolations are used in the code, because it offers a good compromise between accuracy (small diffusivity) and simplicity (numerical cost) [41, 7].

3.1.2. Time-splitting

In low-dimensional systems, the semi-Lagrangian method is very efficient. When applied to higher dimensional problems, one faces the problem of multidimensional interpolation, which is extremely expensive for high dimensional problems. However, this problem has been partially cured by using the time-splitting idea of Cheng and Knorr [23]. Using the incompressibility property, Strang's operator decomposition into space and velocity can be applied, replacing equation (5) by a set of two conservative equations. Besides, to avoid dealing with a 3D space operator, the latter is also divided into two parts. Let us denote $\mathcal{X}_G = (x_G^1, x_G^2) = (r, \theta)$ and remind that in our case $x_G^3 = \varphi$ then the Boltzmann equation (5) is solved by applying a splitting of Strang [93] as

$$B_{\parallel s}^* \frac{\partial \bar{F}_s}{\partial t} + \nabla_{\perp} \cdot \left(B_{\parallel s}^* \frac{d\mathcal{X}_G}{dt} \bar{F}_s \right) = 0 \quad \text{at } (\varphi, v_{G\parallel}) \text{ fixed} \quad (54)$$

$$B_{\parallel s}^* \frac{\partial \bar{F}_s}{\partial t} + \frac{\partial}{\partial \varphi} \left(B_{\parallel s}^* \frac{d\varphi}{dt} \bar{F}_s \right) = 0 \quad \text{at } (\mathcal{X}_G, v_{G\parallel}) \text{ fixed} \quad (55)$$

$$B_{\parallel s}^* \frac{\partial \bar{F}_s}{\partial t} + \frac{\partial}{\partial v_{G\parallel}} \left(B_{\parallel s}^* \frac{dv_{G\parallel}}{dt} \bar{F}_s \right) = 0 \quad \text{at } \mathbf{x}_G \text{ fixed} \quad (56)$$

This splitting into three equations was introduced in the 4D version of the code. Indeed, as explained in [45], in the 4D drift-kinetic slab case the conservative and advective forms of the equations are equivalent in the (r, θ) direction and z direction separately due to the independent vanishing divergence property, see eqs (9)-(10) in [45] due to $\nabla v_{GC} = 0$ and $\partial_z v_{\parallel} = 0$. In the 5D gyrokinetic case $\nabla_{\perp} (B_{\parallel s}^* d_t \mathcal{X}_G) \neq 0$ and $\partial_{\varphi} (B_{\parallel s}^* d_t \varphi) \neq 0$ so that these terms should be taken into account as source terms of the advective form of equations (54) and (55). However, they are presently set to zero. As shown in section 5, this simplification does not prevent obtaining good conservation properties but likely has an impact on the maximal value acceptable for the discretization time step. A solution to overcome this problem could be to use a conservative scheme instead of the BSL one but the first tests we have performed [12] are not conclusive. This constraint on the numerical value of Δt is acceptable for ion turbulence simulations but could become problematic when addressing kinetic electrons. The development of more efficient semi-lagrangian schemes

is still an active axis of research. An idea currently under evaluation is to separate and to treat differently the linear and non-linear parts. Encouraging results have been presented in paper [66].

In the current version, the advections in φ and $v_{G\parallel}$ directions are straightforward, but that in the $\mathcal{X}_G$ direction requires more attention. Indeed, if we consider the action of the $\tilde{\mathcal{X}}_G$ operator between times t and $t + \Delta t$, the value of the electric field $\mathbf{E}$ at time $t + \Delta t/2$ is required in second order time scheme. This value is calculated by using a predictor-corrector method. Besides, computing these 2D trajectories is equivalent to solving $d\mathcal{X}_G/dt = \mathbf{V}(\mathcal{X}_G, \varphi, t)$, $\mathbf{V}$ being the advection field. This system is solved by using the parabolic assumption developed in [90]. Let $\mathcal{X}_{G_{ij}}$ be the position of $\mathcal{X}_G(t_n + \Delta t)$ at time $t_n + \Delta t$, then there exists a displacement $\mathbf{d}_{ij} = (\alpha_{ij}, \beta_{ij})$ tangent to the parabola such that $\mathcal{X}_G(t_n) = \mathcal{X}_{G_{ij}} - \mathbf{d}_{ij}$ and $\mathcal{X}_G(t_n - \Delta t) = \mathcal{X}_{G_{ij}} - 2\mathbf{d}_{ij}$. The displacement $\mathbf{d}_{ij}$ can be calculated by solving the implicit equation $\mathbf{d}_{ij} = \Delta t \mathbf{V}(\mathcal{X}_{G_{ij}} - \mathbf{d}_{ij}, t_n)$ (see p. 129 in [46]). This is done with a Taylor expansion which is equivalent at second order to a Newton algorithm.

3.1.3. GYSELA global algorithm

Concerning now the complete Boltzmann equation (5) the right hand side is also split to treat separately the collision operator, the Krook operator, the diffusion and source terms. Let $\tilde{\mathcal{X}}_G$ denote the shift operator in the poloidal cross-section (r, θ) associated to equation (54) over a time step Δt . Similarly, $\tilde{\varphi}$ and $\tilde{v}_{G\parallel}$ denote the shift operators respectively in the φ (equation (55)) and $v_{G\parallel}$ directions (equation (56)). As described in the previous paragraph, each of these three shift operators are based on a backward semi-Lagrangian scheme which means two steps for each mesh point: (i) first the computation of the characteristic feet and (ii) second an interpolation by cubic splines. Let us denote $\tilde{\mathcal{C}}$ the collision operator corresponding to solving $\partial_t \bar{F}_s = \mathcal{C}(\bar{F}_s)$ and $\tilde{\mathcal{D}}$ the operator associated to the radial diffusion $\partial_t \bar{F}_s = \mathcal{D}_r(\bar{F}_s)$. They are both solved by applying a Crank-Nicolson scheme (see Appendices AppendixB.2 and AppendixE). The Krook operator $\tilde{\mathcal{K}}$ corresponding to the Krook diffusion equation $\partial_t \bar{F}_s = \mathcal{K}(\bar{F}_s)$ is trivial while solving $\partial_t \bar{F}_s = \mathcal{S}$ associated to the source operator $\tilde{\mathcal{S}}$ is described in AppendixD.5. Then, using these notations the following sequence is used to solve the 5D Boltzmann equation (5)

$$\widetilde{\text{Boltz}} \equiv \left[\left(\frac{\tilde{\mathcal{K}}_r}{2}, \frac{\tilde{\mathcal{D}}_r}{2}, \frac{\tilde{\mathcal{C}}}{2}, \frac{\tilde{\mathcal{S}}}{2} \right), \left(\frac{\tilde{v}_{G\parallel}}{2}, \frac{\tilde{\varphi}}{2}, \tilde{\mathcal{X}}_G, \frac{\tilde{\varphi}}{2}, \frac{\tilde{v}_{G\parallel}}{2} \right), \left(\frac{\tilde{\mathcal{S}}}{2}, \frac{\tilde{\mathcal{C}}}{2}, \frac{\tilde{\mathcal{D}}_r}{2}, \frac{\tilde{\mathcal{K}}_r}{2} \right) \right] \quad (57)$$

where the factor 1/2 means that the operator is applied on half a time step. The choice of the sequence (57) is not unique but some constraints are imposed in the code: (i) the first one is to impose a symmetry to keep second order accuracy in the splitting (Strang splitting [93]), (ii) the second one is to fix the 2D operator $\tilde{\mathcal{X}}_G$ which is the most costly at the middle of the algorithm; (iii) finally the operators coupled to the right hand side, respectively to the Vlasov equation, are contiguous. Let us also define the operator $\tilde{Q}$ which denotes symbolically the four steps: (i) computation of right hand side of the quasi-neutrality equation (12) using expression (13), (ii) solving the QN equation (12) to deduce the electrostatic potential U , (iii) computation of the gyro-averaged electric potential $\bar{U} = \mathcal{J}_\mu U$ with a Padé approximation (18) or with an integration on the gyro-circles as described in section 2.5 and (iv) computation of the electric field as $\mathbf{E} = -\nabla \bar{U}$.

Finally, the global numerical algorithm of the GYSELA code can be summarized (see schematic view in Figure 2) as follows,

1. Initialization

Considering a prescribed magnetic field $B(r, \theta)$ (see section 2.2) and equilibrium profiles of density $n_0(r)$, ion temperatures $T_s(r)$ and safety factor $q(r)$ (see section 2.6), then

- (a) Computation of the equilibrium distribution function $\bar{F}_{s,\text{eq}}$ as a local or canonical Maxwellian by using (19) and (20).
- (b) Initialization of $\bar{F}_s(t=0)$ as $\bar{F}_s(t=0) = \bar{F}_{s,\text{eq}}(1 + \text{perturbation})$ as described in section 2.6.

2. For each time iteration,

Considering the distribution function $\bar{F}_s^n = \bar{F}_s(t = t_n)$ at time t_n known on the 5D mesh grid, then

the distribution function $\bar{F}_s^{n+1}$ at the next time $t_{n+1} = t_n + \Delta t$ on the same mesh grid is computed by using a predictor-algorithm as

- (a) Computation of the electric field $\mathbf{E}(t_n)$ by using the $\tilde{Q}$ sequence.
- (b) **Prediction on $\Delta t/2$:**
 - Computation of $\bar{F}_s(t = t_n + \Delta t/2)$ by solving $\widetilde{\text{Boltz}}/2$ sequence with $\mathbf{E}(t_n)$.
 - Computation of $\mathbf{E}(t_n + \Delta t/2)$ by solving $\tilde{Q}$.
- (c) **Correction on Δt :**
 - Starting from $\bar{F}_s(t = t_n)$ –given that this 5D distribution function has been stored before prediction–, computation of $\bar{F}_s(t = t_n + \Delta t)$ by solving $\widetilde{\text{Boltz}}$ on a time step Δt with the electric field $\mathbf{E}(t_n + \Delta t/2)$ at time $t_{n+1/2}$.

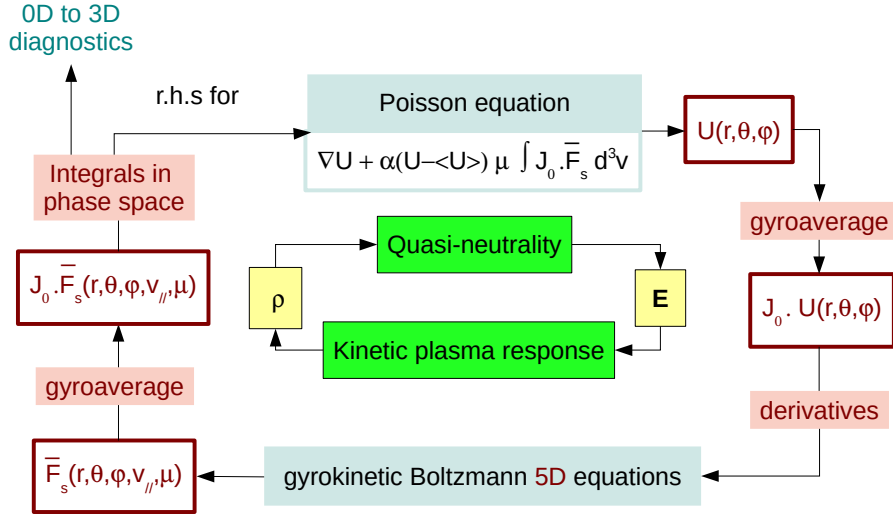


Figure 2: Schematic view of the coupling between the Poisson and Vlasov solvers in the GYSELA code.

3.2. An efficient hybrid OpenMP/MPI parallelization

The code is developed in Fortran 90 with some I/O routines in C (47k lines of Fortran 90 and 2.3k lines of C code) using GIT as version control system. Doxygen is used to generate documentation from the source code. The only external library dependence is the HDF5 library. HDF5 is the chosen format for all output saving, both 0D to 3D diagnostics and 5D restart files. Diagnostic analyses are performed with Python. The parallelization is based on a hybrid MPI/OpenMP paradigm. This hybrid approach is suitable for cluster of SMP (symmetric shared memory multiprocessor) nodes where MPI provides communication capability across nodes and OpenMP exploits loop level parallelism within a node. Let us denote by N_{species} the number of ion species and by N_{μ} the number of μ values. As the magnetic momentum μ is an adiabatic invariant it plays the role of a parameter. So for each species we have N_{μ} independent Boltzmann equations (5) to solve. Let N_r , N_{θ} , N_{φ} and $N_{v_{G\parallel}}$ be the number of points in each directions r , θ , φ and $v_{G\parallel}$. Large data structures are used in GYSELA: 5D data of size $N_r \times N_{\theta} \times N_{\varphi} \times N_{v_{G\parallel}} \times N_{\mu}$ for distribution functions and 3D data of size $N_r \times N_{\theta} \times N_{\varphi}$ for the electrostatic potential and its derivatives as for first moments of the distribution function (used for diagnostics). An MPI parallelization is mandatory to treat such large amount of data. Indeed, let us take the example of a typical 5D mesh used for our simulations, namely $(N_r \times N_{\theta} \times N_{\varphi} \times N_{v_{G\parallel}} \times N_{\mu}) = (256 \times 128 \times 128 \times 128 \times 16)$. It corresponds to a mesh of almost 8.6 billion of points. The size of one 5D array for the distribution function is of the order of 68 Gbytes, which is not

tractable on a single node. For information, the biggest simulation run so far with GYSELA was an ITER simulation [3] with 272 billion of points. Taking into account the fact that two distribution functions are necessary for the numerical integration over time due to predictor-corrector scheme, more than 1 Tbytes of data (just for 5D arrays) were manipulated. So, as described in the following, we use a domain decomposition so that a MPI process never contains the complete 5D distribution function.

Concerning the MPI parallelization, an MPI communicator is defined per species. Inside each one of the MPLSPECIES communicators, an MPI communicator is defined for each value of the magnetic moment μ . Within each MPLMU communicator a 2D domain decomposition allows us to assign to each MPI process a sub-domain in (r, θ) dimensions. Let us consider p_r (resp. p_θ) the number of sub-domains in r (resp. θ) direction. The number of MPI processes used during one run is equal to $N_{MPI} = p_r \times p_\theta \times N_\mu \times N_{\text{species}}$. This MPI decomposition of the default MPI.COMM.WORLD communicator is summarized in Figure 3. Thus, each MPI process is then responsible for the sub-domain of the distribution function $\bar{F}_s(r = [i_{\text{start}}, i_{\text{end}}], \theta = [j_{\text{start}}, j_{\text{end}}], \varphi = *, v_{G\parallel} = *, \mu = \mu_{\text{id}})$ with the integer $\mu_{\text{id}} \in [0, N_\mu - 1]$. The local values $i_{\text{start}}, i_{\text{end}}, j_{\text{start}}$ and j_{end} associated to the parallel decomposition are initially set by using a classical domain decomposition in $(p_r \times p_\theta)$ blocks. The OpenMP paradigm is used in addition to MPI. Let us denote by N_{thread} the number of threads in each MPI process, then the number of cores for a simulation corresponds to $N_{\text{thread}} \times N_{MPI}$.

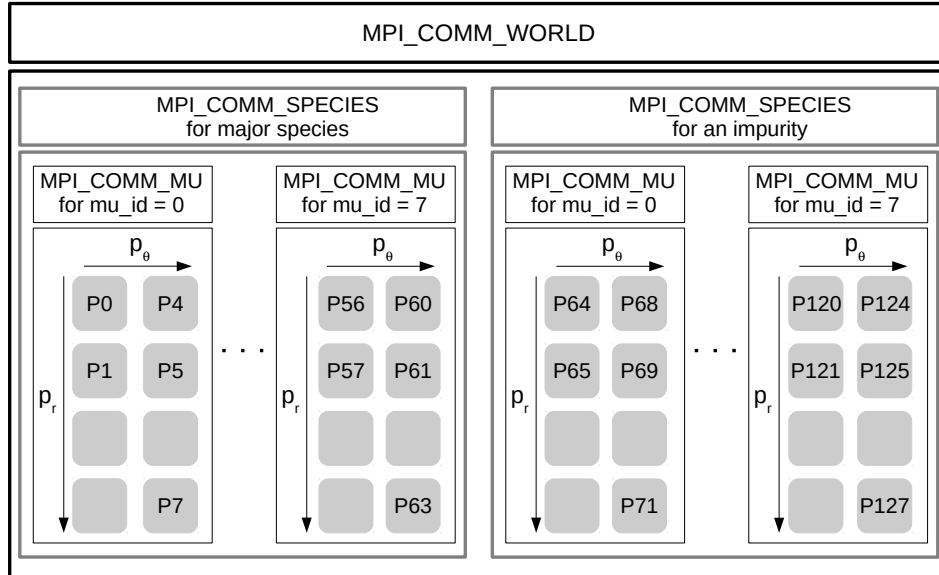


Figure 3: MPI.COMM.WORLD communicator decomposition for two species, 8 values of μ , $p_r = 4$ radial sub-domains and $p_\theta = 2$ sub-domains in the poloidal direction. In this case, the number of MPI processes is equal to 128.

3.2.1. Parallel Vlasov solver: How to treat non-local interpolation aspects ?

A difficulty when parallelizing the semi-Lagrangian Vlasov solver is due to the cubic spline interpolation. Cubic splines are a good compromise between simplicity and accuracy but a drawback is that they are non-local. Indeed, a lot of the values of the distribution function $\bar{F}_s$ are required to reconstruct the interpolated value of the function at any position in the domain. Two strategies are available in the code to overcome the problem: (i) local cubic spline interpolation or (ii) transposition. The Hermite spline interpolation on patches [24, 25, 65] has been specifically developed for dealing with 2D domains distributed on several MPI processes. The idea is to compute local cubic spline coefficients on each 2D (r, θ) sub-domains by solving reduced linear systems. Then one ensures a $\mathcal{C}^1$ global interpolator similar to the sequential one by imposing Hermite boundary conditions at the interface of each patch [24]. The first limitation of this technique is that a minimum of 32 points per directions is needed per MPI process [65] (*i.e.* $N_r/p_r \geq 32$ and $N_\theta/p_\theta \geq 32$) to provide good numerical stability and small communication overhead (in 2D, each processor has to exchange

derivatives with its 8 neighboring processors). The second constraint is that the shift at one point on the border of a sub-domain, which results from the motion along the trajectories in the 4D phase space, must not exceed the elementary cell width. This constraint is linked to the choice made to limit the size of the interface transferred between processors. This CFL condition can be extremely restrictive specially in the θ direction where large shifts can occur but also in the radial direction when a source is imposed in the case of flux-driven simulations. For these reasons the choice of a 4D data transposition is now often preferred. This transposition consists in modifying in each MPI_COMM_MU communicator the parallel decomposition of $\bar{F}_s$ such that each processor then contains only part of the data in φ and $v_{G\parallel}$ direction but all the information in the poloidal cross-section. Standard cubic spline interpolation in (r, θ) plane are then possible. Let us define the transposition operation T_F and its inverse T_F^{-1} as

$$\bar{F}_s(r_{\text{block}}, \theta_{\text{block}}, \varphi = *, v_{G\parallel} = *, \mu = \mu_{\text{id}}) \xrightleftharpoons[T_F^{-1}]{T_F} \bar{F}_s(r = *, \theta = *, \phi_{\text{block}}, v_{\parallel \text{block}}, \mu = \mu_{\text{id}})$$

then the sequence (57) described above for solving the Boltzmann equation is replaced by the following one

$$\widetilde{\text{Boltz}}_{\text{transp}} \equiv \left[\left(\frac{\tilde{\mathcal{K}}_r}{2}, \frac{\tilde{\mathcal{D}}_r}{2}, \frac{\tilde{\mathcal{C}}}{2}, \frac{\tilde{\mathcal{S}}}{2} \right), \left(\frac{\tilde{v}_{G\parallel}}{2}, \frac{\tilde{\varphi}}{2} \right), T_F \left(\tilde{\mathcal{X}}_g \right) T_F^{-1}, \left(\frac{\tilde{\varphi}}{2}, \frac{\tilde{v}_{G\parallel}}{2} \right), \left(\frac{\tilde{\mathcal{S}}}{2}, \frac{\tilde{\mathcal{C}}}{2}, \frac{\tilde{\mathcal{D}}_r}{2}, \frac{\tilde{\mathcal{K}}_r}{2} \right) \right]$$

3.2.2. A scalable quasi-neutral solver

The parallel quasi-neutrality algorithm presently used in the code is summarized in algorithm 1. For more details on the different improvements which have been performed to obtain this parallel solver see [68, 64]. The presence of the non-local term $\langle U \rangle_{\text{FS}}(r)$ couples the θ and φ directions and penalizes the parallelization but the most important cost is the communication induced by the computation of the r.h.s, namely task 2 in algorithm 1. Indeed this calculation requires a collective communication step that involves all MPI processes. Instead of broadcasting $U(r = *, \theta = *, \varphi = *)$ to all MPI processes, a refined strategy has been setup to reduce the large communication cost. It consists in sending to each process only a sub-domain in φ direction of U . With this decomposition the gyro-average computation of U as the partial derivatives in r and θ directions are straightforward. Afterwards, a transposition is performed to calculate $\partial_\varphi U$. For a complete performance analysis of the different steps of the algorithm see [69].

Algorithm 1: Quasi-neutrality algorithm in the GYSELA code

Input: $\bar{F}_s(r_{\text{block}}, \theta_{\text{block}}, \varphi = *, v_{G\parallel} = *, \mu = \mu_{\text{id}})$

1. $v_{G\parallel}$ integration of $\bar{F}_s$ to compute within each MPI.COMM.MU communicator

$$\text{intdvp_Fs}(r_{\text{block}}, \theta_{\text{block}}, \varphi = *, \mu = \mu_{\text{id}}) = \int J_v dv_{G\parallel} (\bar{F}_s - \bar{F}_{s,\text{eq}})$$

2. Remapping within each MPI.COMM.MU communicator of `intdvp_Fs` –because the gyroaverage operation requires to have all the data for each (r, θ) plane– as

$$\text{intdvp_Fs}(r_{\text{block}}, \theta_{\text{block}}, \varphi = *, \mu = \mu_{\text{id}}) \Rightarrow \text{intdvp_Fs}(r = *, \theta = *, \varphi_{\text{block}}, \mu_{\text{block}}).$$

3. For all φ , computation of the gyroaverage of `intdvp_Fs` and integration over μ to obtain

$$\rho_s(r = *, \theta = *, \varphi_{\text{block}}) = \int d\mu_s \mathcal{J}_\mu \cdot (\text{intdvp_Fs}(r = *, \theta = *, \varphi_{\text{block}}))$$

4. MPI reduction towards the N_{species} MPI.COMM.SPECIES communicators to finally compute $\rho = \sum_s Z_s \rho_s$, the right hand side of the quasi-neutrality equation (12).
5. Solving (14) and (15) to deduce $U(r = *, \theta = *, \varphi_{\text{block}})$ and broadcast to the $N_{\text{species}} \times N_\mu$ communicators.

Output: $U(r = *, \theta = *, \varphi_{\text{block}})$ on each MPI process.

3.2.3. Performing weak and strong scaling

The weak and strong scaling, presented in this section, have been performed on two different high performance computers. The strong scaling has been tested on the thin nodes of the Curie machine at TGCC⁴, Bruyères-le-Châtel, France which are based on bullx B510 architecture. Each node contains two 8 cores processors (INTEL Xeon E5-2680 Sandy-Bridge) running at 2.7 GHz with 64 GB of 1.6 GHz memory. The weak scaling has been performed on the IBM Blue Gene/Q machine JUQUEEN at JSC/IAS⁵, Juelich, Germany which is composed of 24 racks grouping 1024 nodes. Each node contains a single 17-cores processor running at 1.6 GHz with 16 GB of 1.33 GHz memory. These cores are twice slower compared to Sandy Bridge cores. The improvements of the code to adapt it efficiently to the BlueGene architecture are detailed in [9]. The scaling results are detailed in Figures 4 and 5 for the 4 main components of the code, namely (i) *Vlasov solver*, solving the Boltzmann equation as described in section 3.2.1; (ii) *Field solver*, solving the quasi-neutrality equation as summarized in algorithm 1; (iii) *Derivatives computation*, the computation of the first derivatives of the gyroaveraged electrostatic potential $\mathcal{J}_\mu U$ in the three r , θ and φ directions and (iv) *Diagnostics* corresponding to all physical quantities from 0D to 3D computed and saved in HDF5 format like densities, parallel and perpendicular temperatures, velocities, fluxes, energies et caetera. The strong scaling has been performed with the mesh size parameters $N_r = 512$, $N_\theta = 512$, $N_\varphi = 128$, $N_{v_{G\parallel}} = 128$ and $N_\mu = 32$. The number of threads was kept constant $N_{\text{thread}} = 8$ to assign two MPI process per node. The couple (p_r, p_θ) of processors in r and θ directions take the following values (2, 4), (8, 2), (8, 4), (16, 4), (16, 8) and (16, 16) so that the scaling spans from 2048 to 65536 cores. For the weak scaling the testbed case was composed from 64k to 458k cores. The considered meshes vary from 17.18 to 481 billion points defined with $N_r = 512$, $N_\theta = 1024$, $N_\varphi = 128$, $N_{v_{G\parallel}} = 128$ and 7 different values of $N_\mu = 2, 4, 8, 16, 32, 48$ and 56. The triplet $(p_r, p_\theta, N_{\text{thread}})$ is kept constant equal to (16, 32, 64). The number of threads is chosen equal to 64 such that a single MPI process is mapped per node and 4 threads are dedicated per core as determined

⁴<http://www-hpc.cea.fr/en/complexe/tgcc.htm>

⁵http://www.fz-juelich.de/ias/jsc/EN/Home/home_node.html

for BlueGene/Q optimization. The results of both scalings are summarized in Figures 4 and 5. Let us first remark that the weak scaling test (Fig. 5) exhibits an excellent scalability of the code with 90.9% of relative efficiency at 458752 cores –which corresponds to the totality of the JUQUEEN computer– compared to 16k. Concerning the strong scaling GYSELA globally scales with a relative efficiency of 89% at 16k cores and 60.6% on 65k cores compared to 2048 cores (Fig. 4). This is already a very good result for such a semi-lagrangian code. Looking into more detail, we see that the deterioration from 89% at 16k to 60.6% at 65k is mainly due to the diagnostics and field solver. Indeed, the Vlasov solver which represents 60% of the application at 2k cores and 48% at 65k cores exhibits a good efficiency of 74.6% at 65k processors. Conversely, the diagnostics which correspond to 37.5% of the total time for the reference case, end taking as much time as the Vlasov solver due to the decreased efficiency at 55.8%. A lot of work has already been done to improve the field solver [64, 68, 69] but this work must continue because an efficiency of 41% impacts the global scalability. Let us finally notice that even if the computation time of the derivatives remains negligible until 16k, a further effort is needed to prepare the code to future exascale machines.

The execution times are not comparable in the two scalings presented here because the considered meshes are not the same. Performance comparisons between BlueGene/Q and Bullx architectures are found in [9]. Production runs are commonly running on both architectures. In general, depending on simulation parameters, the code is a factor 2.5 to 3 times faster on bullx machines (as Curie or Helios at CSC, Rokkasho, Japan) than on BlueGene machines (as JUQUEEN or Turing at IDRIS, Orsay, France or Fermi at CINECA, Bologna, Italy). This behavior is consistent with that observed with many other codes.

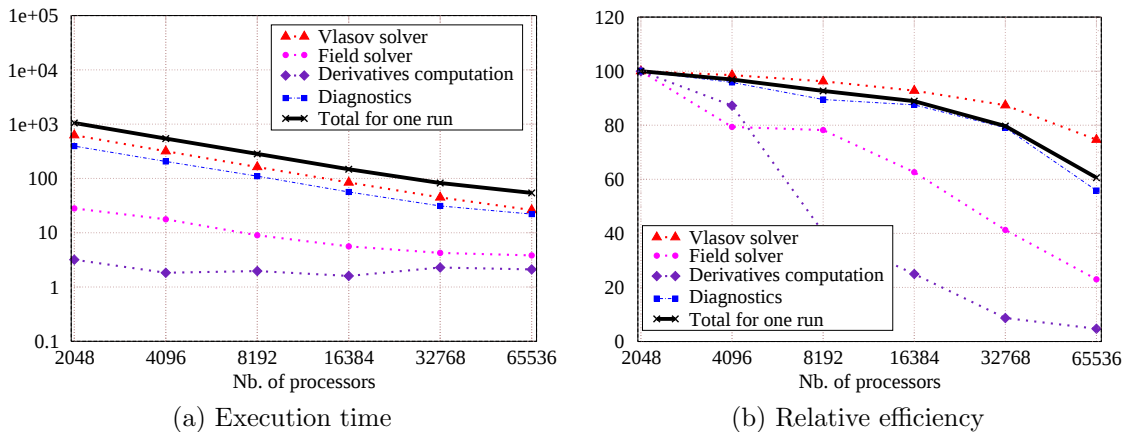


Figure 4: Strong scaling performed on the Curie machine from 2048 to 65536 cores: Execution time (a) and relative efficiency (b) for one GYSELA run of 4 iterations for a mesh $(N_r \times N_\theta \times N_\varphi \times N_{v_{G||}} \times N_\mu) = (512 \times 512 \times 128 \times 128 \times 32)$ with 8 threads and 32 values in μ direction.

3.2.4. Memory scalability

Due to the previous scalability results GYSELA uses frequently from 8k to 32k cores for one ion species with adiabatic electrons and the twice when an impurity is taken into account. Besides, a simulation of ten runs on several weeks. The annual time consumption on HPC machines is currently of 51 millions of core hours. So the code already benefits from petascale computational power of the current high performance computers. We also know that at short term when adding kinetic electrons, a simulation with ITER parameters on several confinement times will require exascale HPC capabilities. Among the Exascale challenges, the reduced memory per core has been identified as one of the most critical. This is particularly true for the GYSELA code due its global character which requires huge 5D meshes. The mesh discretization is already constrained by the memory required per node. The C/Fortran MTM (Modelization & Tracing Memory consumption) library has been developed [82] to investigate in detail the memory consumption of the code. This library provides an Application Programming Interface (API) which replaces the standard calls to allocation/de-allocation routines. This intrusive technique permits to retrieve precisely the peak of

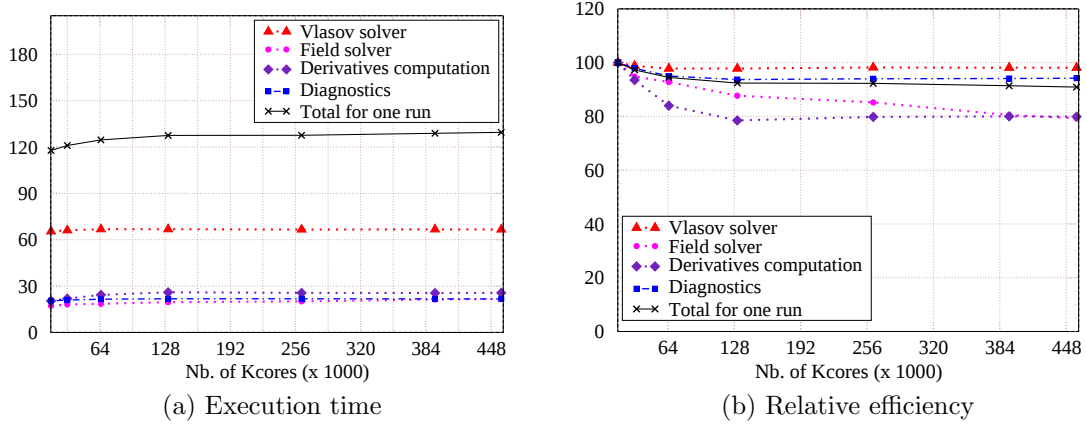


Figure 5: Weak scaling performed on the JUQUEEN machine from 64k to 458k cores: Execution time (a) and relative efficiency (b) for one GYSELA run of 4 iterations for a mesh fixed in 4D as $(N_r \times N_\theta \times N_\varphi \times N_{v_{G\parallel}}) = (512 \times 1024 \times 128 \times 128)$ but for 7 different values of $\mu = 2, 4, 8, 16, 32, 48, 56$. The number of parallel domains in r, θ directions and the number of threads are fixed such that $(p_r, p_\theta, N_{\text{thread}}) = (16, 32, 64)$. The number of cores varies as $p_r \times p_\theta \times N_\mu \times N_{\text{thread}}/4$ because 4 threads per core are used for BlueGene optimization.

memory consumption and all the arrays involved. External Python scripts have been designed to analyze these results and provide memory prediction. Continued efforts are made to reduce the memory footprint of the code and to improve its memory scalability (see [82] for details) but Table 1 shows that lots of work remains to be done to be able to run ITER simulations on machine with only 16GB per node. Indeed, for a minimum ITER mesh, namely $(N_r \times N_\theta \times N_\varphi \times N_{v_{G\parallel}} \times N_\mu) = (1024 \times 1024 \times 256 \times 128 \times 16)$, Table 1 shows that such simulations can run on Curie Machine (64 GB/node) with 8192 cores but would require 524288 cores on the BlueGene current architecture (16 GB/node)- knowing that the optimal number of threads is 64. This number of cores is still out of reach on European high performance computers. Currently, as

MPI procs	OpenMP threads			
	16	32	64	128
128	126.1	126.9	128.7	132.2
512	35.2	36.1	37.8	41.3
2048	16.4	16.4	16.5	23.2
8192	12.5	12.5	12.5	19.1

Table 1: Memory peak (in GB) depending on the number of MPI procs and of OpenMP threads for a 5D mesh $(N_r \times N_\theta \times N_\varphi \times N_{v_{G\parallel}} \times N_\mu) = (1024 \times 1024 \times 256 \times 128 \times 16)$.

shown in Table 2, the global memory scalability of the code is of 47.8%. This strong scaling was performed with a constant 5D mesh of $(1024 \times 4096 \times 1024 \times 128 \times 2)$ points using MTM prediction mode and varying the number of MPI processes from 128 to 2048. The 4D structures are very scalable (relative efficiency of 89.9%) compared to the others. Indeed, 3D structures with 23.8% represent 32% of the global cost for 32k cores against 16% for 2048 cores. The 3D structures are then no more negligible. The scalability of the 1D structures is with 19% of the order of the 3D arrays but the cost is still not significant. At the opposite, the incompressibility of 2D structures between 256 and 2k MPI processes is extremely penalizing leading to an increase of the cost percentage by a factor 10. Improvement of memory footprint of 3D and 2D structures will be pursued in the future.

Another bottleneck for Exascale applications will be the possible increased crash probability following that of the number of cores of future machines [20]. The Weibull law [88] gives an estimate of the time between two crashes which is of the order of the minute for a number on nodes larger than 10^5 . Two approaches

Number of cores Number of MPI processes	2k 128	4k 256	8k 512	16k 1024	32k 2048
4D structures	207.2 79.2%	104.4 71.5%	53.7 65.6%	27.3 52.2%	14.4 42.0%
3D structures	42.0 16.1%	31.1 21.3%	18.6 22.7%	15.9 30.4%	11.0 32.1%
2D structures	7.1 2.7%	7.1 4.9%	7.1 8.7%	7.1 13.6%	7.1 20.8%
1D structures	5.2 2.0%	3.3 2.3%	2.4 3.0%	2.0 3.8%	1.7 5.1%
Total per MPI process in GBytes	261.5	145.9	81.8	52.3	34.2

Table 2: Strong scaling: memory allocation size and percentage with respect to the total for each kind of data at the memory peak for a 5D mesh ($1024 \times 4096 \times 1024 \times 128 \times 2$). (Table from [82]).

have been explored until now to try to improve the fault tolerance of the code. The first one consists in employing an asynchronous method to increase the frequency of the restart file writing. A GYSELA simulation runs typically several days or even weeks. So the complete simulation is split into a series of jobs of approximately 10 hours with automatic re-submission. The restart files are not only saved at the end of each job but several times during the job in parallel to calculations (see [98] for complete description). Another checkpointing using the FTI library [6] (high performance Fault Tolerance Interface for hybrid systems) is under investigation. The main idea is to benefit from fast access local SSD disks available on some HPC architectures.

4. GYSELA verification

Since the first steps 15 years ago, the code has evolved a lot, including more and more physics with more and more numerical complexity and high level of parallelism. Portability of the code on several HPC architectures has been achieved. In such a complex code it becomes very difficult to track or to propagate changes by being sure to reproduce previous results and simulations. Over the past two years, a large effort has been made to improve the development process with due regard to modularity, reproducibility and efficiency. This approach is based on the joint use of a version control system (GIT⁶) together with that of a continuous integration platform such as JENKINS⁷. At each commit on the GIT reference branch, automatic compiling and executing are submitted. For more details on our strategy the reader can refer to [8]. A database of non-regression tests is also under construction with the objective to be run less frequently (every night or week) but to ensure that new changes have no impact on well-established results. All the tests proposed in this section, which were used for GYSELA verification⁸, are detailed with the objective to become part of the non-regression database. They all correspond to 5D gyrokinetic simulations for one species. Other verification tests can be founded on simpler 4D models in [45, 66].

4.1. Normalization

The GYSELA code is written in normalized units based on the following normalization choices. We use SI units and a thermal energy scale in electron volts ($1eV = 1.6022 \cdot 10^{-19} J$). The four fundamental dimensional normalizing quantities are: a reference ionic mass $m_0 = A_0 m_p$ (Kilogram), a reference ionic charge $q_0 = Z_0 e$ (Coulomb), a reference magnetic induction B_0 (Tesla) and a reference thermal energy T_0 (eV). Here, A_0 and Z_0 are the (dimensionless) mass number and charge state of the main ion species and e the modulus of the

⁶<http://git-scm.com/>

⁷<http://jenkins-ci.org>

⁸The numerical results presented in the following are based on the GIT release 17.0 of the code.

electron charge. These quantities are used to define the reference ion cyclotron frequency Ω_{c_0} , the reference thermal speed v_{T_0} and the reference Larmor-radius ρ_0 as $\Omega_{c_0} = \frac{Z_0 e B_0}{m_0}$, $v_{T_0} = \sqrt{\frac{T_0}{m_0}}$ and $\rho_0 = \frac{v_{T_0}}{\Omega_{c_0}} = \frac{\sqrt{T_0 m_0}}{Z_0 e B_0}$. Finally, we choose the equilibrium electron density at mid radius n_0 as reference density. Physical quantities (mass, length, time, charge and density) can be recovered from the normalized quantities used in the code (denoted with a hat symbol) by choosing values for $[A_0, Z_0, B_0, T_0, n_0]$ and applying $m_s = A_s m_p = m_0 \hat{A}_s$ with $A_s = A_0 \hat{A}_s$, $q_s = Z_0 e \hat{Z}_s$, $l = \rho_0 \hat{l}$, $t = \frac{\hat{t}}{\Omega_{c_0}}$ and $n_s = n_0 \hat{n}_s$. The velocities are normalized to the corresponding thermal velocities $v_{T_{s0}} = \sqrt{T_0/m_s}$, *i.e* $v = v_{T_{s0}} \hat{v}_s = \frac{v_{T_0}}{\sqrt{\hat{A}_s}} \hat{v}_s$. The main normalizations are thus $U = \frac{T_0}{Z_0 e} \hat{\phi}$, $B = B_0 \hat{B}$, $T_s = T_0 \hat{T}_s$ while $\mu_s = \frac{T_0}{B_0} \hat{\mu}_s$ with $\hat{\mu}_s = \frac{\hat{v}_{\perp s}^2}{2 \hat{B}}$ and $\mu_0 \mathbf{J} = \frac{B_0}{\rho_0} \mu_0 \hat{\mathbf{J}}$ with $\mu_0 \hat{\mathbf{J}} = \hat{\nabla} \times \hat{\mathbf{B}}$. By deduction, $\mathbf{E} = E_0 \hat{\mathbf{E}}$ with $E_0 = v_{T_0} B_0$ and the normalized distribution function $\hat{F}_s$, which evolves in the code, is defined as $\hat{F}_s = F_s v_{T_{s0}}^3 / n_0$. Finally, the energy is normalized to the reference thermal energy T_0 . The subsequent normalized equations used in the code are presented in Appendix [AppendixF](#). In what follows, all the quantities considered are normalized coordinates, but hat symbols are omitted for the sake of readability.

4.2. Invariance test

In the present work the source terms are not taken into account. Let us call **simu1** a first fixed gradient simulation (*i.e* no source term) for one unique species of charge Z_1 and mass A_1 and **simu2** a second one for a species of charge $Z_2 = \alpha_Z Z_1$ and mass $A_2 = \alpha_A A_1$. Then it is possible to define the other parameters of the second simulation, only depending on the multiplying factors $\alpha_Z \in \mathbb{N}$ and $\alpha_A \in \mathbb{N}$ such that both simulations **simu1** and **simu2** are identical. The idea is to define an invariance test which permits to verify that charge and mass are correctly taken into account in the code. Such a test can be defined as a non-regression test of any gyrokinetic code but it is not generic. It depends on the equations treated in each code. In the case of GYSELA, as shown in Appendix [AppendixG](#), 7 control parameters are required α_x , α_t , α_{T_e} , α_ϕ , α_{coll} , α_{diff} and α_{Krook} that respectively rescale space, time, electron temperature, electrostatic potential, collision operator, diffusion term and Krook operator. The invariance constraints (*i.e* the relationships between the various scaling factors which leave the Boltzmann equation invariant under these transformations), as summarized in Table 3, are $\alpha_x = \sqrt{\alpha_A}/\alpha_Z$, $\alpha_t = \alpha_A/\alpha_Z$, $\alpha_{T_e} = 1/\alpha_Z$, $\alpha_{\text{coll}} = 1$, $\alpha_{\text{diff}} = 1/\alpha_Z$ and $\alpha_{\text{Krook}} = \alpha_Z/\alpha_A$. Then, the relation between the electrostatic potential ϕ_1 solution of the first simulation and ϕ_2 the electrostatic potential of the second one is $\phi_2 = \alpha_\phi \phi_1$ with $\alpha_\phi = 1/\alpha_Z$.

	Simulation 1	Simulation 2
Charge	Z_1	$Z_2 = \alpha_Z Z_1$
Mass	A_1	$A_2 = \alpha_A A_1$
Mesh discretization	Δx_1	$\Delta x_2 = \frac{\sqrt{\alpha_A}}{\alpha_Z} \Delta x_1$
Time step Δt	Δt_1	$\Delta t_2 = \frac{\alpha_A}{\alpha_Z} \Delta t_1$
Temperature T_e	T_{e1}	$T_{e2} = \frac{1}{\alpha_Z} T_{e1}$
Collision coefficient	$K_{\text{coll}1}$	$K_{\text{coll}2} = K_{\text{coll}1}$
Krook coefficient	$K_{\text{krook}1}$	$K_{\text{krook}2} = \frac{\alpha_Z}{\alpha_A} K_{\text{krook}1}$
Diffusion coefficient	$K_{\text{diff}1}$	$K_{\text{diff}2} = \frac{1}{\alpha_Z} K_{\text{diff}1}$

Table 3: Parameter dependence between two equivalent simulations. α_Z and α_A are the multiplying factors between both cases respectively for charge and mass.

Three different simulations were run to test the invariance property of the code, considering first a reference plasma with Hydrogen (mass $A_s = 1$ and charge $Z_s = 1$), then Helium ($A_s = 4$ and $Z_s = 2$) and finally Tungsten ($A_s = 150$ and $Z_s = 50$). For the following, let us respectively call **simu1_A1Z1**, **simu2_A4Z2** and **simu3_A150Z50** these three simulations. These cases are not relevant in terms of physics

but have been designed for numerical tests. The idea was to define small cases tractable as non-regression tests. So the reference simulation is based on a small plasma ($\rho_* = 1/150$) for a radial domain between 0.15ρ and 0.85ρ (with $\rho = r/a$) and for half a torus. The parallel velocity space is defined as $-7v_{T_{s_0}} \leq v_{G\parallel} \leq 7v_{T_{s_0}}$ and the perpendicular direction is represented by 16 values of μ between 0 and 12. The radial profiles of density $n_{s_0}(r)$, temperature $T_{s_0}(r)$ and safety factor $q(r)$ are analytically prescribed as $d \log n_{s_0}(r)/dr = -2.2 \cosh^{-2}((r-0.5)/0.004)$, $d \log T_{s_0}(r)/dr = -8 \cosh^{-2}((r-0.5)/0.004)$ and $q(r) = 1.5 + 1.7 \exp(2.8 \log(r/a))$. A Krook operator (see section 2.10) of amplitude 0.01 and a diffusion (eq. (51)) of 0.015 are applied in a buffer region defined by Eq.(50) with a location $B_L = 0.06$ and a stiffness of 0.017635. The collision operator (Eq.(30)) is applied every iterations while $T_{s,\text{coll}}$ (Eq.(32)) and $V_{\parallel s,\text{coll}}$ (Eq.(31)) are refreshed every 10 iterations. All the numerical parameters of this **simu1_A1Z1** simulation are summarized in Tables 4 and 5. For the complete description of the case see the associated input data file of the code (Figure H.14) in Appendix AppendixH.

N_r	N_θ	N_φ	$N_{v_{G\parallel}}$	N_μ	$\text{nb}_{v_{th}0}$	$\mu_{\max}$	
256	256	128	128	16	7.	12.	
ρ_*	R_0/a	r_{int}/a	r_{ext}/a	Torus	Z_s	A_s	$\Omega_{c_0} \Delta t$
1/150	3.3	0.15	0.75	1/2	1.	1.	5.
q_1	q_2	q_3	$\kappa_{n_{s_0}}$	$\Delta_{n_{s_0}}$	$\kappa_{T_{s_0}}$	$\Delta_{T_{s_0}}$	T_i/T_e
1.5	1.7	2.8	2.2	0.04	8.	0.04	1.

Table 4: Main parameters for reference simulation **simu1_A1Z1**. The velocity phase space is defined by $-\text{nb}_{v_{th}0}v_{T_{s_0}} \leq v_{G\parallel} \leq \text{nb}_{v_{th}0}v_{T_{s_0}}$ and $0 \leq \mu \leq \mu_{\max}T_0/B_0$. Torus indicates the fraction of the torus simulated. The safety factor radial profile is defined as $q(r) = q_1 + q_2 \exp(q_3 \log(r/a))$. The radial density profile is defined by its gradient as $d \log n_{s_0}(r)/dr = -\kappa_{n_{s_0}} \cosh^{-2}((r-0.5)/\Delta r_{n_{s_0}})$. The same analytical expression is used for the temperature with $\kappa_{T_{s_0}}$ and $\Delta r_{T_{s_0}}$.

Buffer region					Collision operator		
B_L	B_S	D_{coeff}	$\text{Krook}_{\text{coeff}}$	$\nu_{*,\text{buff}}$	$\nu_{*,\text{coeff}}$	$\text{nbstep}_{\text{coll}}$	$\text{nbrefresh}_{\text{coll}}$
0.06	0.017635	0.015	0.01	20.	0.28	1	10

Table 5: Main parameters for **simu1_A1Z1** simulation concerning buffer region and collisions. B_L and B_S correspond respectively to the location and the stiffness of the buffer regions where both diffusion and Krook operator are applied with respective amplitude D_{coeff} and $\text{Krook}_{\text{coeff}}$. The collision operator is applied every $\text{nbstep}_{\text{coll}}$ iterations but refreshed every $\text{nbrefresh}_{\text{coll}}$ iterations.

So taking into account the equivalences defined in Table 3, the second simulation **simu2_A4Z2** is defined with the same parameters as **simu1_A1Z1** except the fact that Helium is considered instead of Hydrogen and (i) $\tau = 2.$, (ii) $\Omega_{c_0} \Delta t = 10.$, (iii) $D_{\text{coeff}} = 0.0075$, (iv) $\text{Krook}_{\text{coeff}} = 0.005$ while (v) the diagnostic time step $\Omega_{c_0} \Delta t_{\text{diag}} = 100$ instead of 50. The species concentration c_s has been divided by 2 to satisfy the constraint $\sum_s c_s Z_s = 1$. The third simulation **simu3_A150Z50** is performed for a Tungsten species $A_s = 150$ and $Z_s = 50$ and differs from **simu1_A1Z1** by the following parameters: (i) $1/\rho_* = 36.7423$ ($\rho_{*\text{simu1}} \times \sqrt{150/50}$), (ii) $c_s = 0.02$, (iii) $\tau = 50$, (iv) $\Omega_{c_0} \Delta t = 15$, (v) $\Omega_{c_0} \Delta t_{\text{diag}} = 150$, (vi) $D_{\text{coeff}} = 0.0003$ and (vii) $\text{Krook}_{\text{coeff}} = 0.01/3$. The differences between the three considered simulations are highlighted in Table 6. The time evolution of the zonal flow component ($m = 0, n = 0$), where m and n are respectively poloidal and toroidal mode number in the middle of the radial domain is presented in Figure 6 for the three simulations. It exhibits a relative error of 10^{-15} when Hydrogen case is compared to Helium case and of 10^{-7} when compared with Tungsten simulation. The loss of accuracy for Tungsten is probably due to the small size of the plasma considered ($a = 36.7423/\rho_0$). In this case boundary conditions can play a more important role. For each test the results are sufficiently accurate to consider that the charge and mass

factors have been correctly implemented in the code equations. The same order of magnitude is obtained when the distribution function values are compared.

	A_s	Z_s	$1/\rho_*$	τ	$\Omega_{c_0}\Delta t$	$\Omega_{c_0}\Delta t_{\text{diag}}$	D_{coeff}	Krook _{coeff}
simu1_A1Z1	1.	1.	150.	1.	5.	50.	0.015	0.01
simu2_A4Z2	4.	2.	150.	2.	10.	100.	0.0075	0.005
simu3_A150Z50	150.	50.	36.7423	50.	15.	150.	0.0003	0.00333333

Table 6: Modified parameters according to the scaling law for the three simulations used in invariance tests.

The same invariance exercise was performed for a smaller reference case: $\rho_* = 1/75$, $N_r = 128$, $N_\theta = 128$ and a full torus (all the other simulation parameters being identical to the ones described in Tables 4 and 5). The two others simulations defined for comparison follow the same rules as previously which corresponds to a ρ_* of $1/18.37$ for the Tungsten case. The relative error is also of the order of 10^{-15} when comparing Hydrogen and Helium but of 10^{-6} when considering the Tungsten. The two cases $\rho_* = 1/75$ for Hydrogen and Helium will be good candidates as non-regression tests. Regarding the Tungsten simulations more work could be done. The first idea would be to consider a $\rho_* = 1/75$ plasma for Tungsten, but this would imply a reference case with $\rho_* \sim 1/306.205$ for the Hydrogen (*i.e* a mesh of the order of 68 billion of points with $(N_r, N_\theta, N_\varphi, N_{v_{G\parallel}}, N_\mu) = (512, 512, 128, 128, 16)$). Such a simulation is feasible but extremely time consuming for a verification test (several hours on 8192 cores). It would not be possible on the development cluster Poincare (92 nodes of 8 Sandy Bridge E5-2670 bi-processors (2.60GHz)) at IDRIS french supercomputing center that has been used for all tests presented above (performed on 512 cores).

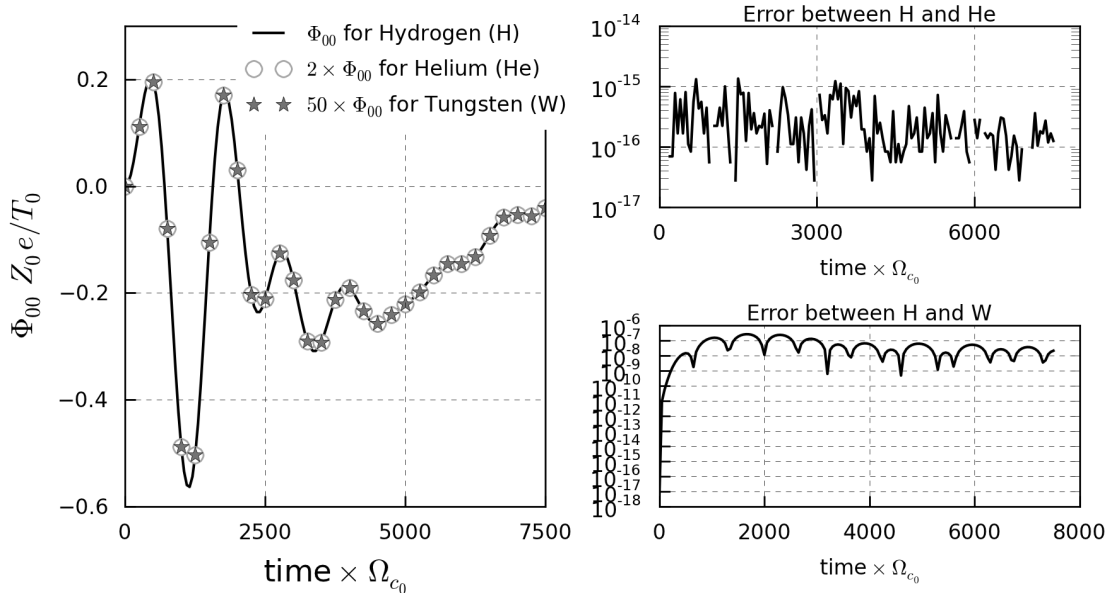


Figure 6: Invariance test between three simulations. Left Figure: Time evolution of Zonal Flows Φ_{00} at a fixed radial position 0.5ρ with $\rho = (r_{\text{max}} - r_{\text{min}})/a$. Φ_{00} for simu1_A1Z1 with Hydrogen must be compared to $2\Phi_{00}$ of simu2_A4Z2 for Helium and $50\Phi_{00}$ obtained with simu3_A150Z50 for Tungsten. Right Figures: Relative errors in function of time: (top) between simu1_A1Z1 and simu2_A4Z2, (bottom) between simu1_A1Z1 and simu3_A150Z50.

4.3. Usual linear and non-linear gyrokinetic benchmarks

Together with the verification invariance test described above, several other benchmarks have already been performed in the past to validate the code. The so-called “Rosenbluth-Hinton test” (RH) [75] –which became an essential test for gyrokinetic codes to check the validity of zonal flows and Geodesic-Acoustic-Modes (GAMs) treatment– was verified in 2008 [48]. Linear and non-linear benchmarks were successfully achieved [48] with the classical Cyclone DIII-D base case (CBC) [37]. Non-linear benchmarks were also performed [47] with the global PIC code ORB5 [60]. We also participated to the European turbulence code benchmarking effort [40]. Finally, flux-driven simulations have been compared [85] between GYSELA, ORB5 and XGC1 [21, 22]. In the following we present only the tests which have been recently investigated again with the new multi-species version of the code to be added as non-regression tests, namely the RH test and the linear CBC test. These tests have been re-designed with the objective of being sufficiently small to be automatically launched on the continuous integration platform (JENKINS) at a reasonable frequency with the aim of recovering the predicted results. Here are described the results of our investigation. This work is still in progress and we plan to continue decreasing the CPU time consumption of such tests in the future.

4.3.1. Rosenbluth-Hinton test as a non-regression test

This test consists in computing the linear evolution of the zonal flow component ϕ_{00} for an initial electrostatic perturbation. In practice it corresponds to initializing the code with the distribution function $\bar{F}_s = \bar{F}_{s,\text{eq}}(1 + \epsilon \sin(2\pi(r - r_{\min})/2L_r))$ with a perturbation amplitude ϵ equal to 10^{-3} . This initial state leads to the development of GAMs which correspond to the $(m, n) = (0, 0)$ mode coupled to sidebands $(m, n) = (\pm 1, 0)$ as a result of toroidal coupling. These GAMs are Landau-damped because of the finite poloidal wavenumber of the sideband while the zonal flows relax towards a residual value which has been analytically predicted in the case of large aspect ratio and small ρ_* [42] as $\phi_{00}(t_\infty) = \phi_{00}(t_0) \times A_r$ with $A_r = 1/(1 + 1.6q^2/\sqrt{r/R})$, r being the minor radius of the considered magnetic surface, R and q corresponding respectively to the major radius and the safety factor on this surface of interest. Eight simulations (detailed in Table 8) were performed varying both time and phase space discretizations. For all simulations (see Table 7 for common parameters), the safety factor profile is constant ($q(r) = 1.9 \forall r \in [r_{\min}, r_{\max}]$). The density and temperature profiles are quasi-constant with $\kappa_{Ts_0} = \kappa_{ns_0} = 10^{-7}$. For a complete description of the case see the associated input data file of the code (Figure H.15) in Appendix AppendixH. Time evolution of the $(0, 0)$ mode is plotted in Figure 7 for the most discretized simulation (simu 1 in Table 8) and shows that this residual value A_r is well recovered with the code. The numerical echo appearing at time $t = 15000\Omega_{C_0}^{-1}$ is due to finite discretization of the velocity phase space. The numerical damping rates γ_{GAM} and frequencies ω_{GAM} reported in Table 8 are all computed between $t \in [0, 15000\Omega_{C_0}^{-1}]$. We observe that all the values are equal to $\gamma_{\text{GAM}} = 2.46459 \cdot 10^{-4}\Omega_{C_0}$ and $\omega_{\text{GAM}} = 9.11061 \cdot 10^{-3}\Omega_{C_0}$ with an error smaller than 3% except for simulations number 4, 6 and 8 in Table 8. The error larger than 7% for both simulations 4 and 6 is due to the fact that the numerical echo appears at time $t = 9000\Omega_{C_0}^{-1}$ instead of $t = 15000\Omega_{C_0}^{-1}$, making the numerical residual estimation impossible. Error in simulation 4 is due to the fact that four values of μ are not sufficient for μ integral computation required for the r.h.s of the quasi-neutrality equation. Results for simulation 6 shows that 64 in parallel velocity direction are not sufficient. The CPU times reported as results in Table 8 correspond to mono-core hours on the Poincaré machine where all simulations were performed with 4 threads and $(2 \times N_\mu)$ MPI processes for a final time equal to $t = 20000\Omega_{C_0}^{-1}$. Simulation 7 corresponding to 30 minutes on 64 cores is as non-regression test for the code.

ρ_*	A_s	Z_s	R_0/a	r_{int}/a	r_{ext}/a
1/100	1.	1.	2.78	0.2	0.8
Torus	nb_{vth0}	μ_{max}	q_1	q_2	q_3
1.	7.	12.	1.9	0.	0.
r_{peak}/a	κ_{ns_0}	Δ_{ns_0}	κ_{Ts_0}	Δ_{Ts_0}	T_i/T_e
0.5	$1.e^{-7}$	0.2	$1.e^{-7}$	0.2	1.

Table 7: Common parameters for Rosenbluth-Hinton test. The velocity phase space is defined by $-\text{nb}_{vth0}v_{Ts_0} \leq v_{G\parallel} \leq \text{nb}_{vth0}v_{Ts_0}$ and $0 \leq \mu \leq \mu_{\text{max}}T_0/B_0$. Torus indicates the fraction of the torus simulated. The safety factor radial profile is defined as $q(r) = q_1 + q_2 \exp(q_3 \log(r/a))$. The radial density profile is defined by its gradient as $d \log n_{s_0}(r)/dr = -\kappa_{ns_0} \cosh^{-2}((r - r_{\text{peak}}/a)/\Delta_{ns_0})$. The same analytical expression is used for the temperature with κ_{Ts_0} and Δ_{Ts_0} .

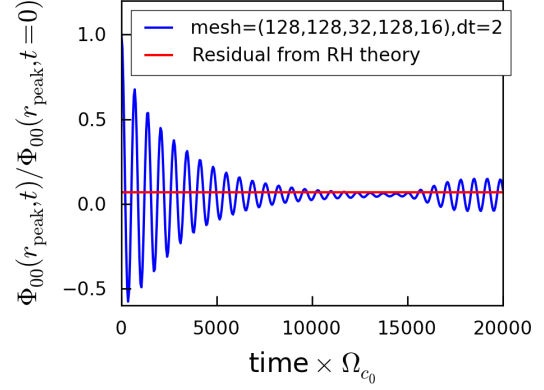


Figure 7: Comparison of the residual value of the (0,0) mode $\phi_{00}(r_{\text{peak}}, t)/\phi_{00}(r_{\text{peak}}, t=0) = A_r$ where $A_r = 1/(1 + 1.6q^2/\sqrt{r/R}) = 0.06825$ is given by Rosenbluth-Hinton theory. This result corresponds to simulation 1 described in Table 8.

simu.	N_r	N_θ	N_φ	$N_{v_{G\parallel}}$	N_μ	$\Omega_{C_0}\Delta t$	γ_{GAM}	ω_{GAM}	CPU time (monoproc)
1	128	128	32	128	16	2.	$2.46459 \cdot 10^{-4}$	$9.11061 \cdot 10^{-3}$	7219. h.
2	128	128	32	128	16	5.	$2.46459 \cdot 10^{-4}$	$9.11061 \cdot 10^{-3}$	2921. h.
3	64	64	16	128	8	5.	$2.46459 \cdot 10^{-4}$	$9.11061 \cdot 10^{-3}$	159. h.
4	64	64	16	128	4	5.	$2.63456 \cdot 10^{-4}$	$9.73893 \cdot 10^{-3}$	79. h.
5	64	64	16	128	8	10.	$2.46459 \cdot 10^{-4}$	$9.11061 \cdot 10^{-3}$	41. h.
6	64	64	16	64	8	10.	$2.54957 \cdot 10^{-4}$	$9.42477 \cdot 10^{-3}$	83. h.
7	64	64	16	128	8	25.	$2.46459 \cdot 10^{-4}$	$9.11061 \cdot 10^{-3}$	37. h.
8	64	64	16	128	8	50.	$2.54957 \cdot 10^{-4}$	$9.42477 \cdot 10^{-3}$	21. h.

Table 8: Rosenbluth-Hinton test: Values of numerical damping γ_{GAM} and frequency ω_{GAM} for 8 GYSELA simulations varying according to 5D mesh size ($N_r, N_\theta, N_\varphi, N_{v_{G\parallel}}, N_\mu$) and time discretization.

4.3.2. Linear benchmark based on Cyclone DIII-D case for global codes

For the present benchmark, the considered physical parameters are the same as the ones defined in Lapillone's paper [63] corresponding to the standard linear Cyclone base case (CBC) [37]. The circular concentric magnetic equilibrium is defined with an aspect ratio of $R/a = 2.78$ and a safety factor profile $q(r) = 0.86 - 0.16r/a + 2.52(r/a)^2$. This corresponds to a local safety factor $q(r_{\text{peak}}) = 1.4$ and a local magnetic shear $s(r_{\text{peak}}) = 0.8$ with $r_{\text{peak}} = r_{\text{min}} + 0.5(r_{\text{max}} - r_{\text{min}})$. The initial density and temperature profiles are defined with the radial form function $f(r) = \exp(-\kappa_x \Delta_{rx} \tanh((r - r_{\text{peak}})/\Delta_{rx}))$. The κ_x and Δ_{rx} parameters are chosen to obtain peaked profiles at the middle of the radial box r_{peak} with $\Delta_{ns_0} = \Delta_{Ts_0} = 0.04$, $\kappa_{ns_0} = 2.2$ and $\kappa_{Ts_0} = 7.1$ (see Figure 1 in [63]). To be consistent with the DIII-D shot ρ_* is chosen equal to 1/180. For a complete description of the case see the associated input data file of the code (Figure H.16) in Appendix AppendixH.

The difficulty for GYSELA is that the CBC benchmark was historically designed for local δf codes. Indeed,

ρ_*	R_0/a	r_{int}/a	r_{ext}/a	nb_{vth0}	μ_{max}	Z_s	A_s
1/180	2.78	0.1	0.9	6.	7.	1.	1.
q_1	q_2	q_3	κ_{ns_0}	Δ_{ns_0}	κ_{Ts_0}	Δ_{Ts_0}	T_i/T_e
0.86	-0.16	2.52	2.2	0.04	7.1	0.04	1.

Table 9: Main parameters of the $\rho_* = 1/180$ linear CBC simulation. The velocity phase space is defined by $-\text{nb}_{vth0}v_{Ts_0} \leq v_{G\parallel} \leq \text{nb}_{vth0}v_{Ts_0}$ and $0 \leq \mu \leq \mu_{\text{max}}T_0/B_0$. The safety factor radial profile is defined as $q(r) = q_1 + q_2(r/a) + q_3(r/a)^2$. The radial density profile is defined by its gradient as $d \log n_{s_0}(r)/dr = -\kappa_{ns_0} \cosh^{-2}((r-0.5)/\Delta r_{ns_0})$. The same analytical expression is used for the temperature with κ_{Ts_0} and Δr_{Ts_0} .

the first difficulty with a full- f code is that the separation between linear and non-linear terms is not possible. So non-linear mode coupling cannot be avoided, which implies that the time window for linear growth rate estimation is limited. To prevent this coupling all the toroidal mode numbers n are filtered except the one n initialized as perturbation in the initial distribution function $\bar{F}_s(t=0) = \bar{F}_{s,\text{eq}}(1 + \epsilon \sum_{m=1}^{m_{\text{max}}} \cos(m\theta + n\varphi + \delta_{mn}))$ with a perturbation amplitude $\epsilon = 10^{-6}$. This filtering is performed by applying, after the quasi-neutrality equation solving, the condition $\mathcal{F}(\phi)_{mn} = 0$ for all $1 \leq m \leq m_{\text{max}}$ where $\mathcal{F}$ denotes the 2D Fourier transform in (θ, φ) . Besides, the global aspect of the code implies that large toroidal mode numbers n are hardly accessible because a large mesh discretization is then necessary. Indeed, for the 6 cases which were run for $|n|$ ranging from 5 to 30 (see Table 10) the discretization of the velocity space was kept constant $(N_{v_{G\parallel}}, N_\mu) = (64, 16)$ but increased in real space for $|n| > 15$. The number of radial points is kept constant ($N_r = 128$) but doubled in both poloidal and toroidal directions for $|n| \geq 25$. This corresponds to a mesh of 8 billions of points for the smallest poloidal mode numbers but of 32 billions of points for the largest. The time step Δt is equal to $40./\Omega_{C_0}$ for $|n| \geq 15$ and divided by a factor 4. The linear growth rate is estimated by a linear fit of the exponential growth of $\int \phi^2 d^3\mathbf{x} d^3\mathbf{v}$ during the linear phase. An example of this exponential growth is plotted at the top of Figure 9 for the smallest $k_\theta \rho_s = |n|q(r_{\text{peak}})\rho_*/r_{\text{peak}} = 0.078$. So the linear fit is performed on a time interval $\Omega_{C_0}[t_{\text{init}}, t_{\text{end}}]$ depending on the duration of this linear phase. This interval is taken equal to $[5000, 50000]$ for $k_\theta \rho_s = 0.078$ and $[4000, 18000]$ for the others. The frequency is estimated on the same time intervals by spectra analysis. Both have been compared with GENE results [63] and show a very good agreement (see Figure 8). The smallest runs were performed on Poincare machine on 512 cores (with $(p_r, p_\theta, p_\mu, N_{\text{thread}}) = (2, 4, 16, 4)$) with a CPU time cost of 9k hours/monoprocessor (for 1300 iterations). The two largest cases ran on Turing (IDRIS-France) BlueGene machine using 32768 processors $((p_r, p_\theta, p_\mu, N_{\text{thread}}) = (8, 4, 16, 64))$ with an expensive CPU time cost of 1.7 millions of monoprocessor hours for 2000 iterations (equivalent to $1.3 \cdot 10^6$ h. on an INTEL machine as Curie (CCRT-France)). Such simulations are definitively more expensive for a global full- f code than for δf flux-tube codes. This explains why the two last points performed by GENE for $k_\theta \rho_* > 0.5$ (Fig. 8) have not been simulated with GYSELA. Nevertheless, the global aspect can give access to useful information like the time evolution of the radial structure of poloidal modes (for toroidal mode $|n|=5$). This radial structure is plotted at 4 different times on Figure 9, going from initialization to the end of the linear phase. It shows that the linear phase, *i.e* the exponential growth, starts as soon as the global eigenmode –characterized by a single n and several m mode numbers– acquires its radial structure. The process of mode reorganization and the main physical parameters involved (as q profile, $1/L_T$ shape or ...) would be interesting to investigate in more detail.

5. Local momentum conservation laws

Verifying adequate conservation laws is an essential step in providing a correct description of mean flows. Since the controversy on the suitability of gyrokinetic codes for describing toroidal momentum transport [79, 80] a specific effort has been devoted to addressed conservation equations both in the context of a reduced gyrofluid model [15] or gyrokinetic field theory [89, 16]. As shown in this section, it is also possible [2] to derive local conservation equations for density, energy and toroidal momentum from the gyrokinetic

n	$m_{\max}$	N_θ	N_φ	$k_\theta \rho_s$	$\Omega_{C_0} \Delta t$	$\Omega_{C_0} [t_{\text{init}}, t_{\text{end}}]$	γ	err_γ	ω
-5	7	256	128	0.078	40.	[5000, 50000]	0.045	0.00258	0.106
-10	14	256	128	0.156	40.	[4000, 18000]	0.158	0.0069	0.238
-15	21	256	128	0.233	40.	[4000, 18000]	0.217	0.00123	0.476
-20	25	256	128	0.311	10.	[4000, 18000]	0.257	0.00127	0.636
-25	35	512	256	0.389	10.	[4000, 18000]	0.271	0.00468	0.827
-30	42	512	256	0.467	10.	[4000, 18000]	0.253	0.002	0.978

Table 10: CBC results: Linear growth rate γ and frequencies ω estimated in the time interval $\Omega_{C_0} [t_{\text{init}}, t_{\text{end}}]$ for 6 different toroidal mode numbers n . The poloidal wave number k_θ is computed as $k_\theta \rho_s = |n|q(r_{\text{peak}})\rho_*/r_{\text{peak}} = |n| \times 1.4/(2 \times 180)$, r_{peak} being the middle of the radial box. These results are compared to GENE results in Figure 8

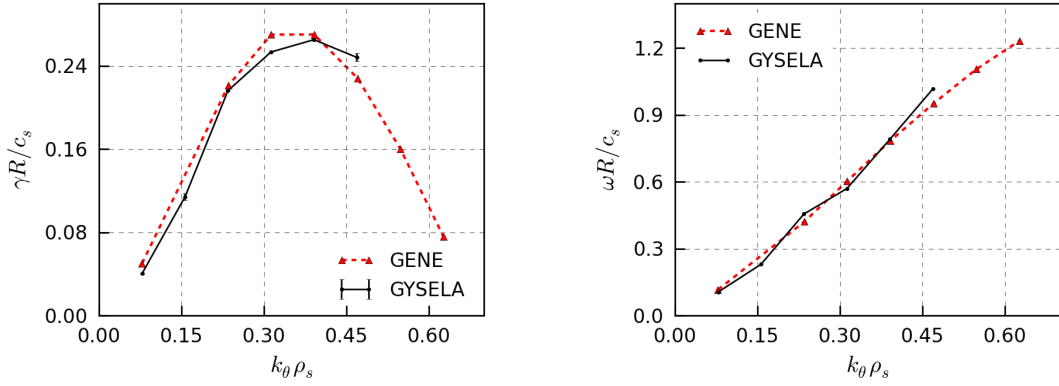


Figure 8: Benchmark between GYSELA and GENE codes for the Cyclone base Case for 6 different $k_\theta \rho_s$ values: (Left) Linear growth rate (plotted with the error bar defined in Table 10), (Right) Linear frequency.

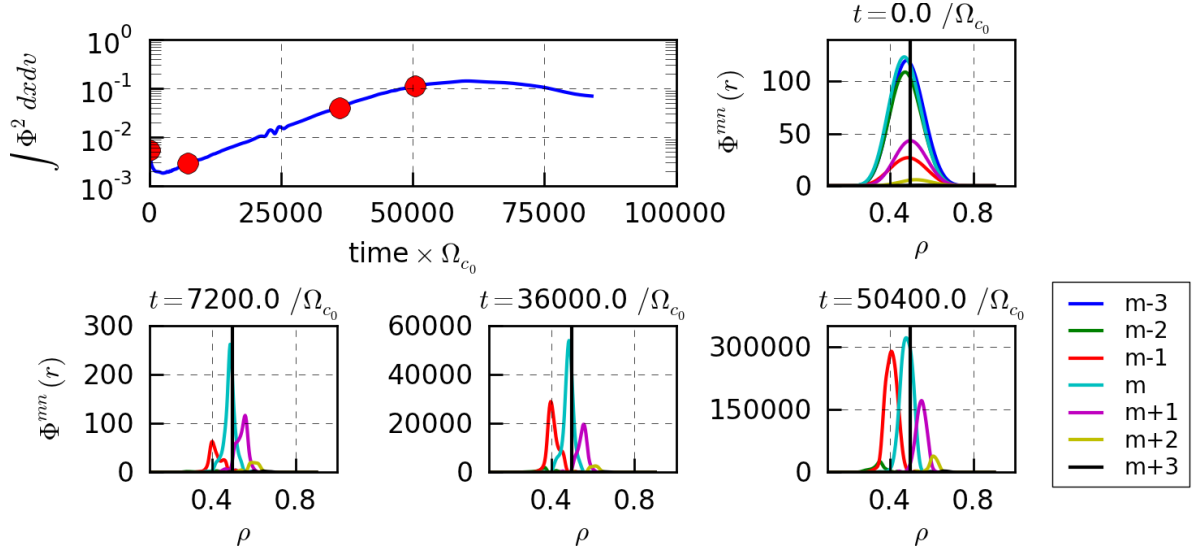


Figure 9: Radial structure of the electrostatic potential for $(m, n) = (7, 5)$ and a bandwidth of $m \pm 3$ modes, at different times: (i) initial time, (ii) beginning of the linear phase $t = 7200\Omega_{C_0}^{-1}$, (iii) $t = 36000\Omega_{C_0}^{-1}$ and (iv) end of the linear phase $t = 50400\Omega_{C_0}^{-1}$.

electrostatic model (5)-(12) implemented in the GYSELA code (more detailed calculations can be found in Abiteboul's PhD [1]). The radial force balance is presented in section 5.1 while the conservation laws are detailed for: (i) charge density in section 5.2, (ii) energy in section 5.3 and toroidal momentum in section 5.4. As explained in [2], the force balance equation added to the energy and toroidal momentum conservation equations ensures a self-consistent treatment of the radial electric field and flows. We present for each property an example of numerical results, see Figures 10 to 13) for a simulation close to the non-linear CBC simulation proposed in [63] for $\rho_* = 1/180$. It corresponds to a collisionless simulation, with Dirichlet boundary conditions and without diffusion or Krook operators. Contrary to the linear case, there is no filtering on the toroidal modes. The initial perturbation is defined as $p(r, \theta, v_{G\parallel}) = \epsilon \sum_{m=1}^{m_{\max}} \sum_{n=1}^{n_{\max}} \cos(m\theta + n\varphi + \delta_{mn}) f_{\text{eq}}(r, \theta, v_{G\parallel}) / (m_{\max} n_{\max})$ for $m_{\max} = 28$ and $n_{\max} = 20$ and random values for δ_{mn} phases. The perturbation amplitude ϵ is equal to 10^{-6} . The equilibrium distribution function f_{eq} is a local Maxwellian function given by Eq.(19). The other numerical parameters of this simulation are summarized in Table 11.

The conservation equations are derived at second order in ρ_* . Accurate results (error of 2%) were already

N_r	N_θ	N_φ	$N_{v_{G\parallel}}$	N_μ	nb _{vth0}	$\mu_{\max}$	
256	256	128	64	16	7.	12.	
ρ_*	R_0/a	r_{int}/a	r_{ext}/a	Torus	Z_s	A_s	$\Omega_{c_0} \Delta t$
1/180	2.78	0.1	0.9	1/2	1.	1.	10.
q_1	q_2	q_3	κ_{ns_0}	Δ_{ns_0}	κ_{Ts_0}	Δ_{Ts_0}	T_i/T_e
0.86	-0.16	2.52	2.2	0.04	7.1	0.04	1.

Table 11: Main parameters of the $\rho_* = 1/180$ simulation used to check conservation law properties. The velocity phase space is defined by $-\text{nb}_{vth0} v_{Ts_0} \leq v_{G\parallel} \leq \text{nb}_{vth0} v_{Ts_0}$ and $0 \leq \mu \leq \mu_{\max} T_0/B_0$. Torus indicates the fraction of the torus simulated. The safety factor radial profile is defined as $q(r) = q_1 + q_2(r/a) + q_3(r/a)^2$. The radial density profile is defined by its gradient as $d \log n_{s_0}(r)/dr = -\kappa_{ns_0} \cosh^{-2}((r-0.5)/\Delta r_{ns_0})$. The same analytical expression is used for the temperature with κ_{Ts_0} and Δr_{Ts_0} .

shown with the code for force balance and toroidal momentum for a simulation with $\rho_* = 1/512$ (see Figure 2 in [2]). For a larger $\rho_* = 1/180$ the results are still accurate with a relative error of 5 to 10%. As expected, boundary conditions play a more important role leading to a degradation of the accuracy outside an internal region $0.4 \leq r/a \leq 0.7$. Comparing to relative errors for all conservation properties (Figures 10 to 13), the local energy conservation (Fig. 12) is the least accurate. It was not possible to analyze the energy behaviour for smaller ρ_* (as 1/300 and 1/512) because such simulations are extremely expensive (several millions of mono-processor hours) and the diagnostic was not appropriate in large simulations until present. This will be investigated in more details on the upcoming big simulations. Complete flux driven simulations with source terms and collisions have also been analyzed (but not presented here). Even for large $\rho_* = 1/150$, the force balance equation, local charge density and toroidal momentum are conserved with less than 10% error even at times when turbulence is fully developed. The fact that this requirement is not met for energy is still under investigation.

5.1. Radial force balance equation

In the fluid description, the radial electric field and the flows are related via the force balance equation. It was verified that this relation holds also in gyrokinetics [32, 43]. Indeed the radial force balance can be recovered analytically from the conventional first order gyrokinetic equations (see Appendix E in [2]), yielding the standard fluid expression

$$\partial_\chi \phi_{eq} + \frac{\partial_\chi P_{eq}}{n_{eq} e} + \frac{B}{I} V_{\parallel eq} = q \frac{B^2 R^2}{I^2} (\mathbf{V} \cdot \nabla \theta) \quad (58)$$

where the equilibrium density n_{eq} , pressure P_{eq} , velocity $V_{\parallel eq}$ and potential ϕ_{eq} are functions of $-P_\varphi/e$ (which is approximately equal to χ at first order in ρ_s). The label of flux surfaces χ is chosen as the opposite of the poloidal flux of the magnetic field, *i.e* $\chi \equiv -\psi_{\text{pol}}$ with $\psi_{\text{pol}} = -(2\pi)^{-1} \int_{S_\theta} dS \mathbf{B} \cdot \nabla \theta / |\nabla \theta|$. Considering the form of the magnetic field used presently in the code $\mathbf{B} = (B_0 R_0 / R(r, \theta)) [\zeta(r) \mathbf{e}_\theta + \mathbf{e}_\varphi]$ with $\zeta(r) = r/(qR_0)$, the label χ is determined by $\chi = B_0 \int_0^r r'/q(r') dr'$ which implies $d\chi/dr = B_0 r/q$.

In order to check that the force balance equation is verified numerically, we compare v_θ^{GYS} the poloidal velocity directly computed within GYSELA—from the distribution function as the sum of the poloidal ExB, curvature, grad- B and magnetization flows—with the expected velocity from the force balance equation, namely v_θ^{FB} corresponding to the left-hand side of Eq. (58) (see Figure 10 (left)). This agreement was robustly verified in GYSELA simulations for a wide range of parameters such as temperature gradient, collisionality and normalized gyroradius ρ_s [33, 31]. A precision of 2% was obtained for a ITER parameter case with $\rho_* = 1/512$ [2]. In Figure 10 (right), we recover this good agreement for the case $\rho_* = 1/180$ described in Table 11. Indeed, even for a larger $\rho_* = 1/180$, the relative error—defined here as the difference normalized to the quadratic mean—is smaller than 0.08 at time $t = 20040\Omega_{C_0}^{-1}$ which corresponds to the beginning of the non-linear phase.

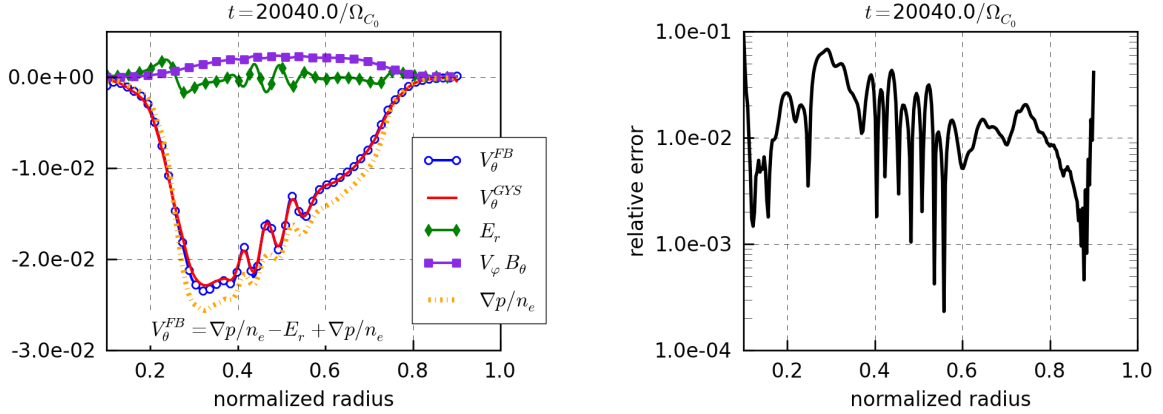


Figure 10: Left: Numerical test of the radial force balance equation (58) at time $t = 20040\Omega_{C_0}^{-1}$, comparing the poloidal velocity v_θ^{GYS} directly computed in the code and v_θ^{FB} the sum of the three contributions $-E_r$, $\nabla p/n_e$ and $v_\varphi B_\theta$. Right: Relative error between both. The parameters used for this simulation are summarized in Table 11.

5.2. Local charge density conservation

To obtain local conservation equations, we perform integrations of the conservative form of the gyrokinetic equation (5) over the velocity space and over the magnetic flux surfaces, *i.e* over $d\tau^* \equiv J_x d\theta d\varphi d^3\mathbf{v}$ with $d^3\mathbf{v} = J_v dv_{G\parallel} d\mu_s$. We recall that $J_x = 1/(\mathbf{B} \cdot \nabla \theta)$ is the Jacobian in real space and $J_v = 2\pi B_{\parallel s}^*/m_s$ is the Jacobian in gyro-center velocity-space.

We consider a radial region outside the buffer region and without source terms, *i.e* we consider the following simplified version of the conservative Boltzmann equation (5)

$$\frac{d\bar{F}_s}{dt} = \frac{\partial \bar{F}_s}{\partial t} + \frac{1}{B_{\parallel s}^*} \nabla_{\mathbf{z}} \cdot (\dot{\mathbf{z}} B_{\parallel s}^* \bar{F}_s) = \mathcal{C}(\bar{F}_s) \quad (59)$$

where $\mathbf{z} = (\chi, \theta, \varphi, v_{G\parallel}, \mu_s)$ and $\dot{\mathbf{z}} = d_t \mathbf{z}$. We integrate Eq. (59) in a small phase-space volume between two surfaces χ and $\chi + \delta\chi$ and apply a divergence theorem. Using the collision operator conservation property $\int \mathcal{C}(\bar{F}_s) d\tau^* = 0$ and summing over all species, this leads directly to the usual expression for local transport of charge density

$$\partial_t \bar{\rho} + \partial_\chi J^\chi = 0 \quad (60)$$

where $\bar{\rho}$ is the charge density and J^χ is the radial current of gyrocenters:

$$\bar{\rho} = \sum_{species} q_s \int d\tau^* \bar{F}_s \quad \text{and} \quad J^\chi = \sum_{species} q_s \int d\tau^* (\dot{\mathbf{z}} \cdot \nabla \chi) \bar{F}_s \quad (61)$$

Figure 11 shows that equation (61) is numerically satisfied with a very good accuracy. Indeed, even for a large $\rho_* = 1/180$, the relative error between $\partial_t \bar{\rho}$ and $-\partial_\chi J^\chi$ is smaller than 1%.

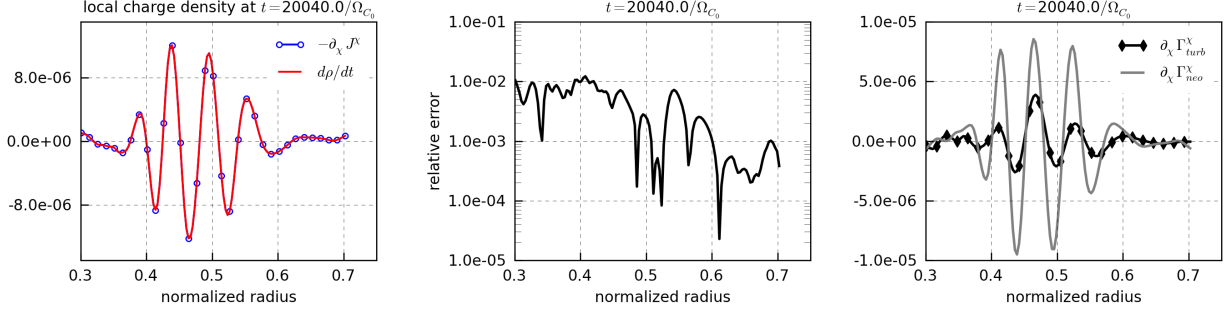


Figure 11: Numerical test of the charge density conservation for simulation defined in Table 11. All the quantities are flux-surface averaged. Left: Comparison of the two terms $-\partial_\chi J^\chi$ and $d\rho/dt$ which must be equivalent according to equation Eq. 60. Middle: Relative error (defined as the difference normalized to the quadratic mean). Right: Contribution of the neoclassical and turbulent parts.

5.3. Local energy conservation

Similarly, a conservation equation for the total energy can be derived by multiplying the gyrokinetic equation (59) by the gyrocenter Hamiltonian, which reads

$$\bar{H}_s = \frac{1}{2} m_s v_{G\parallel}^2 + \mu_s B + q_s \bar{U} \quad \text{with} \quad \bar{U} = J_0 \cdot U \quad (62)$$

Using the fact that our *collision operator* is constructed to *conserve the total Hamiltonian* and integrating over $d\tau^*$ yields

$$\partial_t \int d\tau^* \left(\frac{1}{2} m_s v_{G\parallel}^2 + \mu_s B \right) \bar{F}_s + \int d\tau^* q_s \bar{U} \partial_t \bar{F}_s + \int d\tau^* \bar{H}_s \frac{1}{B_{\parallel s}^*} \nabla_{\mathbf{z}} \cdot (\dot{\mathbf{z}} B_{\parallel s}^* \bar{F}_s) = 0 \quad (63)$$

Then for each species s , equation (63) can be rewritten as an equation for the radial energy transport as

$$\partial_t E_{K_s} + \partial_\chi Q_s = W_s \quad (64)$$

where E_{K_s} corresponds to the kinetic energy of the gyrocenters for the species; the term Q_s is equivalent to a radial flux of energy and the right-hand side W_s appears as a kinetic energy source

$$E_{K_s} \equiv \int d\tau^* \left(\frac{1}{2} m_s v_{G\parallel}^2 + \mu_s B \right) \bar{F}_s \quad ; \quad Q_s = \int d\tau^* \bar{H}_s (\dot{\mathbf{z}} \cdot \nabla \chi) \bar{F}_s \quad ; \quad W_s = -q_s \int d\tau^* \bar{U} \partial_t \bar{F}_s$$

The term W_s corresponds to an exchange of energy between a given species and the turbulence, and is generally referred to as turbulent heating [73, 101, 54, 100, 44]. It corresponds to a transfer of energy between particles and the electromagnetic field. The numerical computation of this term is expensive because it requires saving the 5D distribution function of each species at two successive time steps which

corresponds to a large amount of memory. So to obtain a local conservation equation with no source term, we consider the total energy by summing (64) over all species. Then,

$$\partial_t E_K + \partial_\chi Q = W \quad \text{with} \quad E_K = \sum_s E_{K_s} \quad , \quad Q = \sum_s Q_s \quad \text{and} \quad W = \sum_s W_s \quad (65)$$

The term W is decomposed into two parts as

$$\sum_s q_s \int d\tau^* \bar{U} \partial_t \bar{F}_s = \sum_s q_s \int d\tau^* U (J_0 \cdot \partial_t \bar{F}_s) + \sum_s q_s \int d\tau^* \{ (J_0 \cdot U) \partial_t \bar{F}_s - U (J_0 \cdot \partial_t \bar{F}_s) \} \quad (66)$$

As detailed in Appendix [Appendix I.2](#) the first term in Eq. (I.5) can be expressed in function of the potential energy E_p as

$$\sum_{species} q_s \int d\tau^* U (J_0 \cdot \partial_t \bar{F}_s) \equiv \partial_t E_p \quad \text{with} \quad E_p \equiv \frac{1}{2} \sum_{species} q_s \int d\tau^* U (J_0 \cdot \bar{F}_s)$$

Besides, the second term $\mathcal{I}_E = \sum_s q_s \int d\tau^* \{ (J_0 \cdot U) \partial_t \bar{F}_s - U (J_0 \cdot \partial_t \bar{F}_s) \}$ in Eq. (I.5) corresponds to a polarization term, due to the difference between particles and gyro-center densities. As a remark, considering that the gyroaverage operator J_0 is a self-adjoint operator, this term vanishes when integrating over the whole phase-space volume. It is indeed the divergence of a flux in the local conservation equation. To express this term explicitly as a flux contribution, let us consider the low wavenumber approximation of the gyroaverage operator used in the code, *i.e.* the Padé approximation $J_0 \simeq 1 + \frac{1}{2} \nabla \cdot \left(\frac{m_s \mu_s}{q_s^2 B} \nabla_\perp \right)$ (see section [2.5.1](#) for more details). Using this approximation, $\mathcal{I}_E$ can be expressed for each species as

$$\mathcal{I}_E = \frac{m_s}{2q_s} \partial_\chi \left[\int d\tau^* \partial_t \bar{F}_s \frac{\mu}{B} \nabla \chi \cdot \nabla_\perp U - \int d\tau^* U \frac{\mu}{B} \nabla \chi \cdot \nabla_\perp (\partial_t \bar{F}_s) \right]$$

It can also be expressed in a more compact form using the gyrocenter perpendicular stress $P_{s,\perp} = \int d^3v \bar{F}_s \mu_s B$. This leads to the following conservation equation summed over all species

$$\partial_t (E_K + E_p) + \partial_\chi (Q + Q_{\text{pot}} + Q_{\text{pol}}) = 0 \quad (67)$$

where $E_K = \sum_s E_{K_s}$, $E_p = \sum_s E_{p_s}$, $Q = \sum_s Q_s$, $Q_{\text{pot}} = \sum_s Q_{s,\text{pot}}$ and $Q_{\text{pol}} = \sum_s Q_{s,\text{pol}}$ with

$$E_{K_s} = \int d\tau^* \mathcal{E}_s \bar{F}_s \quad \text{with} \quad \mathcal{E}_s = \frac{1}{2} m_s v_{G\parallel}^2 + \mu_s B \quad (68)$$

$$E_{p_s} = \frac{q_s}{2} \int d\tau^* U (J_0 \cdot \bar{F}_s) \quad (69)$$

$$Q_s = \int d\tau^* \mathcal{E}_s (\dot{\mathbf{z}} \cdot \nabla \chi) \bar{F}_s \quad (70)$$

$$Q_{s,\text{pot}} = \int d\tau^* (J_0 \cdot U) (\dot{\mathbf{z}} \cdot \nabla \chi) \bar{F}_s \quad (71)$$

$$Q_{s,\text{pol}} = \frac{m_s}{2q_s} \int J_x d\theta d\varphi \frac{1}{B^2} \left\{ \partial_t P_{s,\perp} \nabla U \cdot \nabla \chi - U \nabla (\partial_t P_{s,\perp}) \cdot \nabla \chi \right\} \quad (72)$$

We recall that, although the polarization term is necessarily the divergence of a flux term, the expression for Q_{pol} given here is not exact as it relies on an approximation of the gyroaverage operator. Q_s corresponds to the energy flux of species s while $Q_{s,\text{pot}}$ corresponds to the flux due to the electric potential. For the numerical results presented in Figure [12](#) the term $Q_{s,\text{pol}}$ is not taken into account. Indeed, it requires 3D values of U and $P_{s,\perp}$ which were not saved for this simulation because this possibility has been more recently implemented in the code. Analyzing recent non-linear flux-driven simulations where this computation is

available shows that the assumption $Q_{s,\text{pot}} \approx 0$ is justified. Even in regimes where turbulence is well developed the term $Q_{s,\text{pot}}$ stays sufficiently small to have no impact on the local energy conservation law.

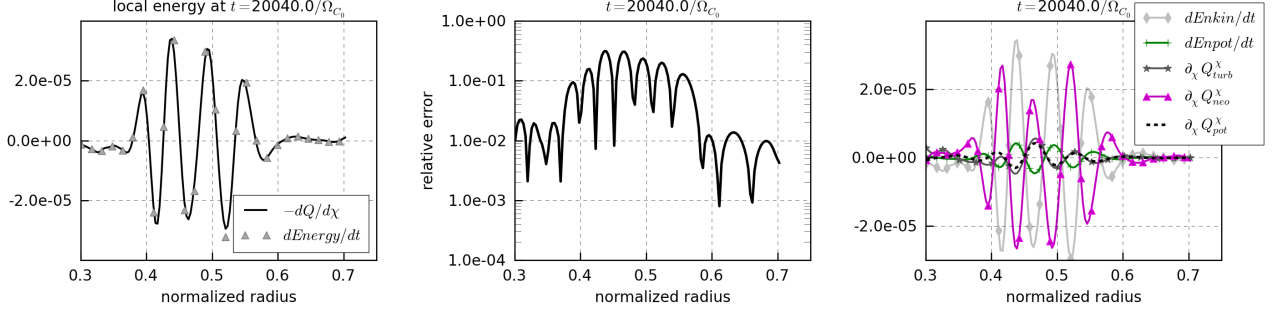


Figure 12: Numerical test for local energy conservation for simulation defined in Table 11. Left: Comparison between $-\langle \partial_\chi Q \rangle_{\text{FS}}$ and $d\langle E \rangle_{\text{FS}}/dt$ with $Q = Q_s + Q_{s,\text{pot}} + Q_{s,\text{pol}}$ and $E = E_{K_s} + E_{p_s}$ defined by Eqs. (67)-(72). Middle: Relative error (defined here as the difference normalized to the quadratic mean). Right: Detailed contribution of each terms.

The numerical validation of equation (67) is performed on flux surface average quantities. The comparison between $\langle \partial_\chi (Q + Q_{\text{pot}} + Q_{\text{pol}}) \rangle_{\text{FS}}$ and $\langle (E_K + E_p) \rangle_{\text{FS}}$ seen in Figure 12 (left) shows an agreement better than 15% at time $t = 20040\Omega_{C_0}^{-1}$. The separate contribution of each terms given by equations (68)-(72) are plotted at Figure 12 (right) showing that this energy conservation derives from the compensation of different complex radial profiles.

5.4. Local toroidal momentum conservation

Formally, the derivation of a conservation law for toroidal angular momentum is very similar to that for energy in the previous section. The general idea is to multiply the conservative gyrokinetic equation (59) by an invariant of motion. For the energy, this invariant was the gyro-center Hamiltonian $\bar{H}_s$ given by Eq. (62). For this conservation law, let us consider the gyrocenter toroidal canonical momentum P_φ defined as

$$P_\varphi = m_s u_\varphi - q_s \chi \quad (73)$$

with the definition $u_\varphi = (I/B)v_{G\parallel} = R^2 \mathbf{b} \cdot \nabla \varphi v_{G\parallel}$. P_φ is an exact invariant of the *unperturbed gyrocenter motion* described by the Hamiltonian $\bar{H}_{s,\text{eq}} = \frac{1}{2} m_s v_{G\parallel}^2 + \mu_s B + \bar{U}_{\text{eq}}$, which corresponds to collisionless motion in a fully axisymmetric system. $\bar{U}_{\text{eq}}$ is the gyroaverage of the equilibrium electric potential U_{eq} independent on the toroidal angle. When axisymmetry is broken, which can occur for instance due to turbulence or magnetic field ripple, P_φ is no longer a motion invariant. In particular, when the electric potential becomes non-axisymmetric, the evolution of P_φ is governed by the equation $d_t P_\varphi = -q_s \partial \varphi \bar{U}$. This result can be obtained by using the expression of the gyrokinetic Poisson brackets in the gyro-center coordinates $[\cdot, \cdot]_{\text{GC}}$ (see [14], equation 150)

$$[X, Y]_{\text{GC}} = \frac{\mathbf{B}_s^*}{m_s B_{\parallel s}^*} \cdot (\nabla X \partial_{v_{G\parallel}} Y - \partial_{v_{G\parallel}} X \nabla Y) - \frac{\mathbf{b}}{q_s B_{\parallel s}^*} \cdot (\nabla X \times \nabla Y) \quad (74)$$

Details of the calculation are presented in Appendix A.3. From the definition of P_φ given by Eq. (73), we define the local toroidal angular momentum as

$$\mathcal{L}_\varphi = \sum_{\text{species}} m_s \int d\tau^* u_\varphi \bar{F}_s \quad (75)$$

Note that $\mathcal{L}_\varphi$ is the momentum for *gyrocenters*, which differs from the particle momentum by terms of order $\mathcal{O}(\rho_s^2)$. As said before, in order to derive a local conservation equation for $\mathcal{L}_\varphi$, we multiply the conservative gyrokinetic equation (59) by P_φ and integrate over all variables other than χ , leading to

$$\int d\tau^* P_\varphi \frac{\partial \bar{F}_s}{\partial t} + \int d\tau^* P_\varphi \frac{1}{B_{\parallel s}^*} \nabla_{\mathbf{z}} \cdot (\dot{\mathbf{z}} B_{\parallel s}^* \bar{F}_s) = \int d\tau^* P_\varphi \mathcal{C}(\bar{F}_s)$$

Using the conservation properties of the collision operator $\int d\tau^* P_\varphi \mathcal{C}(\bar{F}_s) = 0$ and integrating by parts the second term, then

$$\partial_t \left(\int d\tau^* P_\varphi \bar{F}_s \right) - \int d\tau^* \bar{F}_s \frac{\partial P_\varphi}{\partial t} - \int d\tau^* \bar{F}_s \frac{1}{B_{\parallel s}^*} \nabla_{\mathbf{z}} \cdot (\dot{\mathbf{z}} B_{\parallel s}^* P_\varphi) + \partial_\chi \int d\tau^* (\dot{\mathbf{z}} \cdot \nabla \chi) \bar{F}_s P_\varphi = 0$$

Finally, using the fact that $d_t P_\varphi = -q_s \partial_\varphi \bar{U}$, we obtain for each species s ,

$$m_s \partial_t \int d\tau^* u_\varphi \bar{F}_s - q_s \int d\tau^* \chi \partial_t \bar{F}_s + q_s \int d\tau^* \bar{F}_s \partial_\varphi \bar{U} + \partial_\chi \int d\tau^* (\dot{\mathbf{z}} \cdot \nabla \chi) \bar{F}_s P_\varphi = 0 \quad (76)$$

Summing over all species, this leads to

$$\begin{aligned} & \partial_t \left(\sum_s m_s \int d\tau^* u_\varphi \bar{F}_s \right) - \chi \sum_s q_s \int d\tau^* \partial_t \bar{F}_s + \sum_s q_s \int d\tau^* \bar{F}_s \partial_\varphi \bar{U} + \\ & \partial_\chi \left(\sum_s m_s \int d\tau^* (\dot{\mathbf{z}} \cdot \nabla \chi) \bar{F}_s u_\varphi \right) - \partial_\chi \left(q_s \sum_s \chi \int d\tau^* (\dot{\mathbf{z}} \cdot \nabla \chi) \bar{F}_s \right) = 0 \end{aligned}$$

Using the local particle conservation Eq. (60), the second term can be written as $\chi \partial_\chi J^\chi$. We also identify the last term as $-\partial_\chi (\chi J^\chi)$. Then the conservation equation of the local toroidal momentum $\mathcal{L}_\varphi = \sum_s m_s \int d\tau^* u_\varphi \bar{F}_s$ reads

$$\partial_t \mathcal{L}_\varphi + \partial_\chi \Pi_\varphi^\chi + T_{\text{pol}}^\chi = J^\chi \quad (77)$$

where

$$\Pi_\varphi^\chi = \sum_{\text{species}} m_s \int d\tau^* \bar{F}_s u_\varphi v_G^\chi \quad ; \quad T_{\text{pol}}^\chi = \sum_{\text{species}} q_s \int d\tau^* \bar{F}_s \partial_\varphi \bar{U} \quad ; \quad J^\chi = \sum_{\text{species}} q_s \int d\tau^* v_G^\chi \bar{F}_s$$

Equation (77) is an *exact* equation for the transport of gyrocenter toroidal momentum, in the sense that it was obtained directly from the gyrokinetic model, with no specific assumptions regarding the ordering. The term T_{pol}^χ corresponds to a polarization flux of momentum term.

The numerical results obtained with the code are presented in Figure (13) with the contribution of the different terms of equation (77) (at the right) averaged on the flux surface. The relative error plotted at the middle shows an accuracy better than 10%.

6. Conclusion

A complete description of the multi-ion species version of the 5D non-linear gyrokinetic code GYSELA has been presented. Adding to its global full- f character its original semi-lagrangian scheme, the GYSELA code is unique. It can tackle the complex problem of turbulence self-organization, a special care having been taken to fully consistently model the interplay between turbulence and collisions as well as the interplay between all possible scales in the problem, with no assumption of scale separation. Code verification has been successfully performed with a permanent concern for improving its reliability, namely (i) comparison with

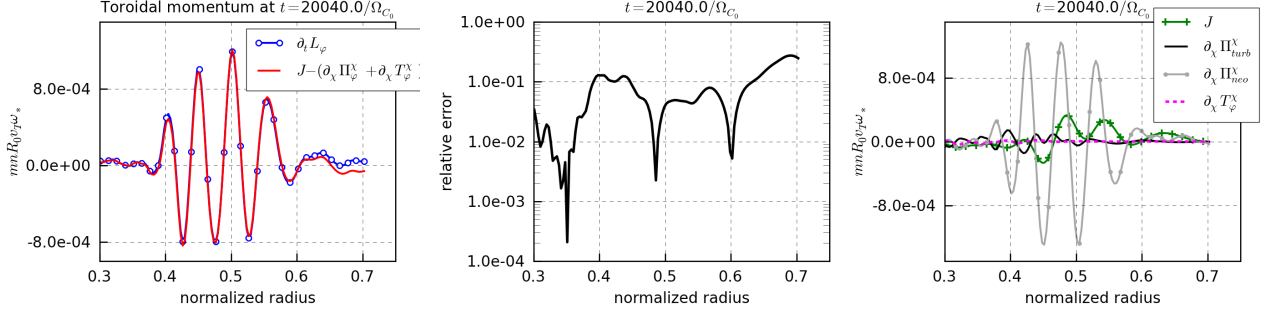


Figure 13: Numerical test for local toroidal momentum conservation for simulation detailed in Table 11. All the quantities are flux-surface averaged. Left: Comparison between $\partial_t L_\varphi$ and $J^X - \partial_\chi \Pi_\varphi^X - T_{\text{pol}}^X$ which must be equal according to Eq. (77). Middle: Relative error (defined here as the difference normalized to the quadratic mean). Right: Detailed contributions of neoclassical and turbulent parts.

analytical theory; (ii) benchmarks with other gyrokinetic codes; (iii) invariance tests to check the correct implementation of mass and charge parameters; and (iv) conservation law properties of the semi-lagrangian method. Particular attention was recently paid to ensure the traceability of the code for the two main developments under progress; the implementation of kinetic electrons and taking into account of a more realistic magnetic configuration. The code is scalable up to actual parallel machine sizes.

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Appendix A. Quasi-neutrality solver

Appendix A.1. How to overcome the difficulty due to $\langle \hat{\phi} \rangle_{\text{FS}}$ term ?

In this appendix, the numerical solving of the quasi-neutrality equation is described. This description is based on the normalized quasi-neutrality equation implemented in the code

$$-\frac{1}{\hat{n}_{e0}} \sum_s \hat{A}_s \hat{\nabla}_\perp \cdot (\hat{n}_{s0} \hat{\nabla}_\perp \hat{\phi}) + \frac{1}{Z_0^2 \hat{T}_e} [\hat{\phi} - \lambda \langle \hat{\phi} \rangle_{\text{FS}}] = \frac{1}{\hat{n}_{e0}} \sum_s \hat{Z}_s [\hat{n}_{G_s} - \hat{n}_{G_s, \text{eq}}] \quad (\text{A.1})$$

It explains the particular treatment performed to overcome the problem of the term $\langle \cdot \rangle_{\text{FS}}$ in Fourier space. The previous equation (F.11) can be written as

$$\mathcal{L} \hat{\phi} + \frac{1}{Z_0^2 \hat{T}_e(\hat{r})} [\hat{\phi} - \lambda \langle \hat{\phi} \rangle_{\text{Flux Surf}}] = \rho(\hat{r}, \theta, \varphi) \quad (\text{A.2})$$

where the differential operator of second order $\mathcal{L}$ is defined as

$$\mathcal{L} = -\frac{1}{\hat{n}_{e0}(\hat{r})} \sum_s \hat{A}_s \hat{n}_{s0}(\hat{r}) \left\{ \frac{\partial^2}{\partial \hat{r}^2} + \left[\frac{1}{\hat{r}} + \frac{1}{\hat{n}_{s0}(\hat{r})} \frac{d\hat{n}_{s0}(\hat{r})}{d\hat{r}} \right] \frac{\partial}{\partial \hat{r}} + \frac{1}{\hat{r}^2} \frac{\partial^2}{\partial \theta^2} \right\}$$

and the right hand side reads

$$\rho(\hat{r}, \theta, \varphi) = \frac{1}{\hat{n}_{e0}(\hat{r})} \sum_s Z_s [\hat{n}_{G_s}(\hat{r}, \theta, \varphi) - \hat{n}_{G_s, \text{eq}}(\hat{r}, \theta)]$$

The constant λ can be chosen equal to 1 (by default) or equal to 0. Let for all function g , $\langle g \rangle_{\theta, \varphi}$ being the radial function equal to

$$\langle g(\hat{r}) \rangle_{\theta, \varphi} = \frac{1}{L_\theta L_\varphi} \int \int g(\hat{r}, \theta, \varphi) d\theta d\varphi$$

By applying the integration $\frac{1}{L_\theta L_\varphi} \int \int \cdot d\theta d\varphi$ to the previous equation (A.2) and by using the fact that $\langle \langle \hat{\phi} \rangle_{\theta, \varphi} \rangle_{\text{Flux Surf}} = \langle \hat{\phi} \rangle_{\text{Flux Surf}}$ then:

$$\mathcal{L} \langle \hat{\phi} \rangle_{\theta, \varphi} + \frac{1}{Z_0^2 \hat{T}_e} \left[\langle \hat{\phi} \rangle_{\theta, \varphi} - \lambda \langle \hat{\phi} \rangle_{\text{Flux Surf}} \right] = \langle \rho \rangle_{\theta, \varphi} \quad (\text{A.3})$$

Let $\hat{\Phi}$ being $\hat{\Phi} = \hat{\phi} - \langle \hat{\phi} \rangle_{\theta, \varphi}$ then, by subtracting (A.3) to (A.2), and by using Dirichlet boundary conditions we obtain $\forall \lambda \in \mathbb{R}, \forall \theta \in [0, L_\theta]$ and $\forall \varphi \in [0, L_\varphi]$:

$$\begin{cases} \left(\mathcal{L} + \frac{1}{Z_0^2 \hat{T}_e} \right) \hat{\Phi}(\hat{r}, \theta, \varphi) = \varrho(\hat{r}, \theta, \varphi) & \text{with } \varrho = \rho - \langle \rho \rangle_{\theta, \varphi} \quad \forall \hat{r} \in [\hat{r}_{\min}, \hat{r}_{\max}] \\ \hat{\Phi}(\hat{r}_{\min}, \theta, \varphi) = \hat{\Phi}(\hat{r}_{\max}, \theta, \varphi) = 0 \end{cases} \quad (\text{A.4})$$

Besides, (A.3) can be written as

$$\begin{aligned} & \mathcal{L} \left(\langle \hat{\phi} \rangle_{\theta, \varphi} - \langle \hat{\phi} \rangle_{\text{Flux Surf}} \right) + \mathcal{L} \langle \hat{\phi} \rangle_{\text{Flux Surf}} \\ & + \frac{1}{Z_0^2 \hat{T}_e} \left[\left(\langle \hat{\phi} \rangle_{\theta, \varphi} - \langle \hat{\phi} \rangle_{\text{Flux Surf}} \right) + \langle \hat{\phi} \rangle_{\text{Flux Surf}} - \lambda \langle \hat{\phi} \rangle_{\text{Flux Surf}} \right] = \langle \rho \rangle_{\theta, \varphi} \end{aligned}$$

Then using the fact that $\langle \hat{\Phi} \rangle_{\text{Flux Surf}} = \langle \hat{\phi} \rangle_{\text{Flux Surf}} - \langle \langle \hat{\phi} \rangle_{\theta, \varphi} \rangle_{\text{Flux Surf}}$ and using Dirichlet boundary conditions, the previous equation leads to the following system:

$$\begin{cases} \left(\mathcal{L} + (1 - \lambda) \frac{1}{Z_0^2 \hat{T}_e} \right) \langle \hat{\phi} \rangle_{\text{Flux Surf}} = \langle \rho \rangle_{\theta, \varphi} + \left(\mathcal{L} + \frac{1}{Z_0^2 \hat{T}_e} \right) \langle \hat{\Phi} \rangle_{\text{Flux Surf}} \\ \langle \hat{\phi} \rangle_{\text{Flux Surf}}(\hat{r}_{\min}) = \langle \hat{\phi} \rangle_{\text{Flux Surf}}(\hat{r}_{\max}) = 0 \end{cases} \quad (\text{A.5})$$

Using the definition of $\hat{\Phi}$, then

$$\begin{aligned} \langle \hat{\Phi} \rangle_{\text{Flux Surf}} &= \langle \hat{\phi} \rangle_{\text{Flux Surf}} - \langle \langle \hat{\phi} \rangle_{\theta, \varphi} \rangle_{\text{Flux Surf}} \\ &= \langle \hat{\phi} \rangle_{\text{Flux Surf}} - \langle \hat{\phi} \rangle_{\theta, \varphi} \quad (\text{because } \langle \langle \hat{\phi} \rangle_{\theta, \varphi} \rangle_{\text{Flux Surf}} = \langle \hat{\phi} \rangle_{\theta, \varphi}) \\ \text{so } \langle \hat{\phi} \rangle_{\theta, \varphi} &= \langle \hat{\phi} \rangle_{\text{Flux Surf}} - \langle \hat{\Phi} \rangle_{\text{Flux Surf}} \end{aligned}$$

and using the fact that $\hat{\phi} = \hat{\Phi} + \langle \hat{\phi} \rangle_{\theta, \varphi}$, we obtain the expression of the electric potential $\hat{\phi}$ as:

$$\hat{\phi}(\hat{r}, \theta, \varphi) = \hat{\Phi}(\hat{r}, \theta, \varphi) - \langle \hat{\Phi} \rangle_{\text{Flux Surf}}(\hat{r}) + \langle \hat{\phi} \rangle_{\text{Flux Surf}}(\hat{r}) \quad (\text{A.6})$$

To summarize, the solving of the equation (F.11) can be replaced by the solving of two simpler equations (A.4) and (A.5). Indeed, the equation (A.5) is a differential equation only depending on the radial direction. Besides, in (A.4) the variable φ plays the role of a parameter, then the discretization of the equation can be performed by projecting in Fourier space in θ direction and by using finite differences in the radial direction

as described in the following paragraph. However, it is important to realize that the boundary conditions are not directly applied on $\hat{\phi}$ but on $\hat{\Phi} = \hat{\phi} - \langle \hat{\phi} \rangle_{\theta, \varphi}$. So the fact to impose $\hat{\Phi}(\hat{r}_{\min}) = 0$ does not imply $\hat{\phi}(\hat{r}_{\min}) = 0$ but $\hat{\phi}(\hat{r}_{\min}) = \langle \hat{\phi} \rangle_{\theta, \varphi}(\hat{r}_{\min})$ (same remark can be done at $\hat{r} = \hat{r}_{\max}$). Another treatment is available when $\hat{r}_{\min}$ is sufficiently close to 0 ($\hat{r}_{\min} < 10^{-2}$ in the code). Indeed, let us assume that in this case $J_x(\hat{r}_{\min}, \theta)$ is equal to a constant. Then, for all function g ,

$$\langle g \rangle_{\text{Flux Surf}}(\hat{r}_{\min}) = \frac{\int g J_x(\hat{r}_{\min}, \theta) d\theta d\varphi}{\int J_x(\hat{r}_{\min}, \theta) d\theta d\varphi} = \frac{1}{4\pi} \langle g \rangle_{\theta, \varphi} \quad \forall |\hat{r}_{\min}| \ll 1$$

In this case applying a Neumann boundary condition on the (0,0) mode at the axis (*i.e* $\partial_r \langle \hat{\phi} \rangle_{\theta, \varphi}(\hat{r}_{\min}) = 0$) is equivalent to apply $\partial_r \langle \hat{\phi} \rangle_{\text{FS}}(\hat{r}_{\min}) = 0$ in the matrix system (A.5).

Appendix A.2. Finite differences in radial direction and Fourier projections in θ and ϕ

Appendix A.2.1. Solving of the equation system (A.4)

Let $\hat{\Phi}$ and ϱ be represented in terms of the Fourier expansion as $\hat{\Phi}(\hat{r}, \theta, \varphi) = \sum_m \hat{\Phi}^m(\hat{r}, \varphi) \exp(im\theta)$ and $\varrho(\hat{r}, \theta, \varphi) = \sum_m \varrho^m(\hat{r}, \varphi) \exp(im\theta)$ then the equation (A.4) can be rewritten in the wave number representation, for each poloidal mode m and for each independent value of φ , as the following differential equation:

$$\left(\mathcal{L}^m + \frac{1}{Z_0^2 \hat{T}_e(\hat{r})} \right) \hat{\Phi}^m(\hat{r}, \varphi) = \varrho^m(\hat{r}, \varphi) \quad (\text{A.7})$$

with the operator $\mathcal{L}^m$ defined as

$$\mathcal{L}^m = -\frac{1}{\hat{n}_{e0}(\hat{r})} \sum_s \hat{A}_s \hat{n}_{s0}(\hat{r}) \left\{ \frac{\partial^2}{\partial \hat{r}^2} + \left[\frac{1}{\hat{r}} + \frac{1}{\hat{n}_{s0}(\hat{r})} \frac{d\hat{n}_{s0}(\hat{r})}{d\hat{r}} \right] \frac{\partial}{\partial \hat{r}} - \frac{m^2}{\hat{r}^2} \right\} \quad (\text{A.8})$$

and where $\hat{\Phi}^m$ (resp. ϱ^m) is the Fourier transform in θ of $\hat{\Phi}$ (resp. ϱ). Let N_r be the number of radial points and let assumes that the radial domain is defined inside $[\hat{r}_1, \hat{r}_{N_r}]$ (*i.e* $\hat{r}_{\min} = \hat{r}_1$ and $\hat{r}_{\max} = \hat{r}_{N_r}$), then up to second order in $\Delta\hat{r}$, the system of equations (A.4) leads to the tridiagonal $(N_r - 2) \times (N_r - 2)$ system:

$$\begin{pmatrix} d_{\hat{r}_2}^m & u_{\hat{r}_2} & 0 & 0 \\ l_{\hat{r}_3} & d_{\hat{r}_3}^m & u_{\hat{r}_3} & 0 \\ & \ddots & \ddots & \ddots \\ 0 & l_{\hat{r}_{N_r-2}} & d_{\hat{r}_{N_r-2}}^m & u_{\hat{r}_{N_r-2}} \\ & 0 & l_{\hat{r}_{N_r-1}} & d_{\hat{r}_{N_r-1}}^m \end{pmatrix} \begin{pmatrix} \hat{\Phi}_2^m \\ \hat{\Phi}_3^m \\ \vdots \\ \hat{\Phi}_{N_r-2}^m \\ \hat{\Phi}_{N_r-1}^m \end{pmatrix} = \begin{pmatrix} \varrho_2^m \\ \varrho_3^m \\ \vdots \\ \varrho_{N_r-2}^m \\ \varrho_{N_r-1}^m \end{pmatrix} \quad (\text{A.9})$$

with for each $\hat{r}_i \in [\hat{r}_1, \hat{r}_{N_r}]$

$$\begin{cases} l_{\hat{r}_i} = -\left(\frac{\beta(\hat{r}_i)}{\Delta\hat{r}^2} - \frac{\alpha(\hat{r}_i)}{2\Delta\hat{r}} \right) \\ d_{\hat{r}_i}^m = \beta(\hat{r}_i) \left(\frac{2}{\Delta\hat{r}^2} + \frac{m^2}{\hat{r}_i^2} \right) + \frac{1}{Z_0^2 \hat{T}_e(\hat{r}_i)} \\ u_{\hat{r}_i} = -\left(\frac{\beta(\hat{r}_i)}{\Delta\hat{r}^2} + \frac{\alpha(\hat{r}_i)}{2\Delta\hat{r}} \right) \\ \varrho^m = \varrho^m(\hat{r}_i) \end{cases} \quad \text{with} \quad \begin{cases} \alpha(\hat{r}_i) = \sum_s \hat{A}_s \frac{\hat{n}_{s0}(\hat{r}_i)}{\hat{n}_{e0}(\hat{r}_i)} \left(\frac{1}{\hat{r}_i} + \frac{1}{\hat{n}_{s0}(\hat{r}_i)} \frac{d\hat{n}_{s0}(\hat{r}_i)}{d\hat{r}} \right) \\ \beta(\hat{r}_i) = \sum_s \hat{A}_s \frac{\hat{n}_{s0}(\hat{r}_i)}{\hat{n}_{e0}(\hat{r}_i)} \end{cases} \quad (\text{A.10})$$

and where $\hat{\Phi}_1^m = \hat{\Phi}_{N_r}^m = 0$.

Solving the previous matrix system (A.9) is equivalent to solve a matrix system of the form $A\mathbf{x} = \mathbf{b}$ where A is tridiagonal, with $[d_1, \dots, d_N] = [d_{\hat{r}_2}, d_{\hat{r}_3}, \dots, d_{\hat{r}_{N_r-1}}]$ the diagonal, $[l_1, \dots, l_N] = [0, l_{\hat{r}_3}, \dots, l_{\hat{r}_{N_r-1}}]$ the lower diagonal and $[u_1, \dots, u_N] = [u_{\hat{r}_2}, \dots, u_{\hat{r}_{N_r-2}}, 0]$ the upper one while the right side vector $\mathbf{b}$ corresponds to $[b_1, \dots, b_N] = [\varrho^m(\hat{r}_2), \dots, \varrho^m(\hat{r}_{N_r-1})]$. The result vector $\hat{\Phi}^m$ is given by $[\hat{\Phi}_1^m, \dots, \hat{\Phi}_{N_r}^m] = [0, x_1, x_2, \dots, x_N, 0]$.

Appendix A.2.2. Solving of the equation system (A.5)

In the following $\langle \cdot \rangle_{\text{Flux Surf}}$ is replaced by $\langle \cdot \rangle_{\text{FS}}$ for more readability. The system (A.5) can be rewritten as:

$$\begin{cases} \left(\mathcal{L} + (1 - \lambda) \frac{1}{Z_0^2 T_e(\hat{r}_i)} \right) \langle \hat{\phi} \rangle_{\text{FS}}(\hat{r}_i) = \Gamma(\hat{r}_i) & \text{for each } \hat{r}_i \in [\hat{r}_1, \hat{r}_{N_r}] \\ \langle \hat{\phi} \rangle_{\text{FS}}(\hat{r}_{\min}) = \langle \hat{\phi} \rangle_{\text{FS}}(\hat{r}_{\max}) = 0 \end{cases}$$

with $\Gamma(\hat{r}_i) = \langle \rho \rangle_{\theta, \varphi}(\hat{r}_i) + \left(\mathcal{L} + \frac{1}{Z_0^2 T_e(\hat{r}_i)} \right) \langle \hat{\Phi} \rangle_{\text{FS}}$ where $\langle \rho \rangle_{\theta, \varphi}(\hat{r}_i) = \frac{1}{L_\theta L_\varphi} \int \int \rho(\hat{r}_i, \theta, \varphi) d\theta d\varphi$ which is equivalent (by using the same notation than for the previous matrix system (A.9)) to:

$$\begin{pmatrix} d_{\hat{r}_2} & u_{\hat{r}_2} & 0 & & \\ l_{\hat{r}_3} & d_{\hat{r}_3} & u_{\hat{r}_3} & 0 & \\ & \ddots & \ddots & \ddots & \\ & 0 & l_{\hat{r}_{N_r-2}} & d_{\hat{r}_{N_r-2}} & u_{\hat{r}_{N_r-2}} \\ & & 0 & l_{\hat{r}_{N_r-1}} & d_{\hat{r}_{N_r-1}} \end{pmatrix} \begin{pmatrix} \langle \hat{\phi} \rangle_{\text{FS}}(\hat{r}_2) \\ \langle \hat{\phi} \rangle_{\text{FS}}(\hat{r}_3) \\ \vdots \\ \langle \hat{\phi} \rangle_{\text{FS}}(\hat{r}_{N_r-2}) \\ \langle \hat{\phi} \rangle_{\text{FS}}(\hat{r}_{N_r-1}) \end{pmatrix} = \begin{pmatrix} \Gamma(\hat{r}_2) \\ \Gamma(\hat{r}_3) \\ \vdots \\ \Gamma(\hat{r}_{N_r-2}) \\ \Gamma(\hat{r}_{N_r-1}) \end{pmatrix} \quad (\text{A.11})$$

with for each $\hat{r}_i \in [\hat{r}_2, \hat{r}_{N_r-1}]$

$$\begin{cases} l_{\hat{r}_i} = - \left(\frac{\beta(\hat{r}_i)}{\Delta \hat{r}^2} - \frac{\alpha(\hat{r}_i)}{2\Delta \hat{r}} \right) \\ d_{\hat{r}_i} = \frac{2}{\Delta \hat{r}^2} \beta(\hat{r}_i) + (1 - \lambda) \frac{1}{Z_0^2 T_e(\hat{r}_i)} \\ u_{\hat{r}_i} = - \left(\frac{\beta(\hat{r}_i)}{\Delta \hat{r}^2} + \frac{\alpha(\hat{r}_i)}{2\Delta \hat{r}} \right) \\ \Gamma(\hat{r}_i) = \langle \rho \rangle_{\theta, \varphi}(\hat{r}_i) + \left(\mathcal{L} + \frac{1}{Z_0^2 T_e(\hat{r}_i)} \right) \end{cases}$$

where $\alpha(\hat{r}_i)$ and $\beta(\hat{r}_i)$ are defined by (A.10). Let us notice that the super-diagonal $(u_{\hat{r}_i})_{i=2, \dots, N_r-2}$ and the lower-diagonal $(l_{\hat{r}_i})_{i=3, \dots, N_r-1}$ are the same than the ones in the previous matrix system (A.9), while the diagonal can be deduced from the previous diagonal of the poloidal mode $m = 0$ (i.e. $(d_{\hat{r}_i}^m)_{i=2, \dots, N_r-1}$ for $m = 0$) by the relation $d_{\hat{r}_i} = d_{\hat{r}_i}^0 - \lambda / (Z_0^2 T_e(\hat{r}_i))$.

Appendix A.2.3. Global algorithm for the quasi-neutrality solver

Then the different steps for solving (A.2) and obtaining $\hat{\phi}$ are the following:

- (i) Compute and save ρ ,
- (ii) Solve (A.4) to obtain $\hat{\Phi}$ and save the 3D array $\hat{\Phi}$,
- (iii) Compute $\langle \hat{\Phi} \rangle_{\text{Flux Surf}}$ and save this 1D array,
- (iv) Compute the RHS of (A.5), i.e. $\langle \rho \rangle_{\theta, \varphi} + \left(\mathcal{L} + \frac{1}{Z_0^2 T_e} \right) \langle \hat{\Phi} \rangle_{\text{Flux Surf}}$ and save this 1D array,
- (v) Solve (A.5) to obtain $\langle \hat{\phi} \rangle_{\text{Flux Surf}}$ and store it, and finally
- (vi) Compute $\hat{\phi}(\hat{r}, \theta, \varphi)$ by using (A.6).

Appendix B. Numerical implementation of the collision operator in GYSELA

In this appendix, the simplified expression of the Lorentz-type operator which is used in GYSELA is detailed. The expression of this collision operator (including the perpendicular direction, which is not yet implemented in GYSELA) is

$$\mathcal{C}(\bar{F}_s) = \frac{1}{B_{\parallel s}^*} \partial_{v_{G\parallel}} \left\{ B_{\parallel s}^* \mathcal{D}_{\parallel s} \bar{F}_{M_s} \partial_{v_{G\parallel}} \left(\frac{\bar{F}_s}{\bar{F}_{M_s}} \right) \right\} + \frac{1}{B_{\parallel s}^*} \partial_\mu \left\{ B_{\parallel s}^* \mathcal{D}_{\perp s} \bar{F}_{M_s} \frac{1}{B^2} \partial_\mu \left(\frac{\bar{F}_s}{\bar{F}_{M_s}} \right) \right\} \quad (\text{B.1})$$

where $B_{\parallel s}^*(r, \theta, v_{G\parallel}) = B(r, \theta) + v_{G\parallel} \mathbf{b} \cdot \nabla \times \mathbf{b}$ is the Jacobian of the guiding-center coordinates, and $\bar{F}_{M_s}$ is the following shifted Maxwellian distribution

$$\bar{F}_{M_s} = \frac{n_{s0}}{(2\pi T_{s,\text{coll}}/m_s)^{3/2}} \exp\left(-\frac{m_s (v_{G\parallel} - V_{\parallel s,\text{coll}})^2}{2T_{s,\text{coll}}} - \frac{\mu_s B}{T_{s,\text{coll}}}\right) \quad (\text{B.2})$$

with the mean temperature $T_{s,\text{coll}} = T_{s,\text{coll}}(r, \theta, \varphi)$ and the mean velocity $V_{\parallel s,\text{coll}} = V_{\parallel s,\text{coll}}(r, \theta, \varphi)$. $\bar{F}_{M_s}$ is such that $\mathcal{C}(\bar{F}_s) = 0$. The collision term $\mathcal{D}_{\parallel s}$ is defined by equation (25). The expression of $T_{s,\text{coll}}$ and $V_{\parallel s,\text{coll}}$ are constrained such that the collision operator is momentum and energy preserving. These calculations and expressions are detailed in a first paragraph [AppendixB.1](#) while the numerical implementation based on a semi-implicit second order Crank-Nicolson scheme is described in the second one [AppendixB.2](#).

AppendixB.1. Expressions of the mean temperature $T_{s,\text{coll}}$ and mean velocity $V_{\parallel s,\text{coll}}$ for the collision operator

While this simplified collision operator obviously conserves the number of particles, the profiles $T_{s,\text{coll}}(r, \theta, \varphi)$ and $V_{\parallel s,\text{coll}}(r, \theta, \varphi)$ must be chosen so that the operator is also compatible with the conservations of momentum and energy

$$\int d^3\mathbf{v} m_s v_{G\parallel} \mathcal{C}(\bar{F}_s) = 0 \quad (\text{B.3})$$

$$\int d^3\mathbf{v} \left(\mu_s B + \frac{1}{2} m_s v_{G\parallel}^2 \right) \mathcal{C}(\bar{F}_s) = 0 \quad (\text{B.4})$$

where $J_v = 2\pi B_{\parallel s}^*/m_s$ is the jacobian in velocity space and $d^3\mathbf{v} = J_v d\mu_s dv_{G\parallel}$ is the space velocity element. Let us first consider only the contribution of collisions in the parallel direction. We use the expression of the collision operator (B.1) and integrate Eq. (B.3) by parts twice in the variable $v_{G\parallel}$

$$(\text{B.3})_{\parallel} = 2\pi \int d\mu_s dv_{G\parallel} \left\{ \frac{\bar{F}_s}{\bar{F}_{M_s}} B_{\parallel s}^* \mathcal{D}_{\parallel s} \partial_{v_{G\parallel}} \bar{F}_{M_s} + \bar{F}_s \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s}) \right\}$$

Let us define the average $\langle \cdot \rangle$ as $\langle \cdot \rangle = \int d^3\mathbf{v} \cdot \bar{F}_s$ and use the fact that $\partial_{v_{G\parallel}} \bar{F}_{M_s} = -\bar{F}_{M_s} \frac{m_s (v_{G\parallel} - V_{\parallel s,\text{coll}})}{T_{s,\text{coll}}}$, then

$$(\text{B.3})_{\parallel} = \frac{m_s}{T_{s,\text{coll}}} \left\{ V_{\parallel s,\text{coll}} \langle m_s \mathcal{D}_{\parallel s} \rangle - \langle m_s \mathcal{D}_{\parallel s} v_{G\parallel} \rangle + T_{s,\text{coll}} \left\langle \frac{1}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s}) \right\rangle \right\} \quad (\text{B.5})$$

We perform similar operations on the contribution to Eq. (B.4) of collisions in the parallel direction:

$$(\text{B.4})_{\parallel} = \frac{m_s}{T_{s,\text{coll}}} \left\{ V_{\parallel s,\text{coll}} \langle m_s \mathcal{D}_{\parallel s} v_{G\parallel} \rangle - \langle m_s \mathcal{D}_{\parallel s} v_{G\parallel}^2 \rangle + T_{s,\text{coll}} \left\langle \frac{1}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s} v_{G\parallel}) \right\rangle \right\} \quad (\text{B.6})$$

For collisions in the perpendicular direction, Eq. (B.3) is trivially verified. We perform two integrations by parts in the variable μ_s for Eq. (B.4),

$$(\text{B.4})_{\perp} = \frac{2\pi}{m_s} \int dv_{G\parallel} d\mu_s \frac{\bar{F}_s}{\bar{F}_{M_s}} \frac{1}{B} \partial_{\mu} (B_{\parallel s}^* \mathcal{D}_{\perp s} \bar{F}_{M_s})$$

Then using the fact that $\partial_{\mu} \bar{F}_{M_s} = -\bar{F}_{M_s} \frac{B}{T_{s,\text{coll}}}$,

$$(\text{B.4})_{\perp} = \frac{1}{T_{s,\text{coll}}} \left\{ \left\langle \frac{1}{B B_{\parallel s}^*} \partial_{\mu} (B_{\parallel s}^* \mathcal{D}_{\perp s}) \right\rangle T_{s,\text{coll}} - \langle \mathcal{D}_{\perp s} \rangle \right\} \quad (\text{B.7})$$

Using equations (B.5), (B.6) and (B.7), the conservation equations (B.3) and (B.4) form a linear system

in $T_{s,\text{coll}}(r, \theta, \varphi)$ and $V_{\parallel s,\text{coll}}(r, \theta, \varphi)$ as follows

$$\begin{aligned} V_{\parallel s,\text{coll}} \langle m_s \mathcal{D}_{\parallel s} \rangle + T_{s,\text{coll}} \left\langle \frac{1}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s}) \right\rangle &= \langle m_s \mathcal{D}_{\parallel s} v_{G\parallel} \rangle \\ V_{\parallel s,\text{coll}} \langle m_s^2 \mathcal{D}_{\parallel s} v_{G\parallel} \rangle + T_{s,\text{coll}} \left\langle \frac{m_s}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s} v_{G\parallel}) + \frac{1}{B B_{\parallel s}^*} \partial_{\mu} (B_{\parallel s}^* \mathcal{D}_{\perp s}) \right\rangle &= \langle m_s^2 \mathcal{D}_{\parallel s} v_{G\parallel}^2 + \mathcal{D}_{\perp s} \rangle \end{aligned}$$

Solving this system, we find that the conservation constraints are verified if the profiles $V_{\parallel s,\text{coll}}(r, \theta, \varphi)$ and $T_{s,\text{coll}}(r, \theta, \varphi)$ for the collision operator are defined as follows

$$\begin{aligned} m_s P V_{\parallel s,\text{coll}} &= \left\langle \frac{m_s}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s} v_{G\parallel}) + \frac{1}{B B_{\parallel s}^*} \partial_{\mu} (B_{\parallel s}^* \mathcal{D}_{\perp s}) \right\rangle \langle m_s \mathcal{D}_{\parallel s} v_{G\parallel} \rangle \\ &\quad - \left\langle \frac{1}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s}) \right\rangle \langle m_s^2 \mathcal{D}_{\parallel s} v_{G\parallel}^2 + \mathcal{D}_{\perp s} \rangle \\ P T_{s,\text{coll}} &= \langle \mathcal{D}_{\parallel s} \rangle \langle m_s^2 \mathcal{D}_{\parallel s} v_{G\parallel}^2 + \mathcal{D}_{\perp s} \rangle - \langle m_s \mathcal{D}_{\parallel s} v_{G\parallel} \rangle^2 \end{aligned}$$

where

$$P = \langle \mathcal{D}_{\parallel s} \rangle \left\langle \frac{m_s}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s} v_{G\parallel}) + \frac{1}{B B_{\parallel s}^*} \partial_{\mu} (B_{\parallel s}^* \mathcal{D}_{\perp s}) \right\rangle - \langle m_s \mathcal{D}_{\parallel s} v_{G\parallel} \rangle \left\langle \frac{1}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s}) \right\rangle$$

and as a reminder $\langle \cdot \rangle = \int J_v d\mu_s dv_{G\parallel} \cdot \bar{F}_s$ with $J_v = 2\pi B_{\parallel s}^*/m_s$.

Let us remind that only collisions in the parallel direction are taken into account at the moment in the code, *i.e.* $\mathcal{D}_{\perp s} = 0$ in above expressions. Then, let us define the 5 following integrals:

$$\begin{aligned} I_0 &= \langle \mathcal{D}_{\parallel s} \rangle \quad ; \quad I_1 = \langle m_s \mathcal{D}_{\parallel s} v_{G\parallel} \rangle \quad ; \quad I_2 = \langle m_s^2 \mathcal{D}_{\parallel s} v_{G\parallel}^2 \rangle \\ I_3 &= \left\langle \frac{1}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* \mathcal{D}_{\parallel s}) \right\rangle \quad \text{and} \quad I_4 = \left\langle \frac{m_s}{B_{\parallel s}^*} \partial_{v_{G\parallel}} (B_{\parallel s}^* v_{G\parallel} \mathcal{D}_{\parallel s}) \right\rangle \end{aligned}$$

Therefore the mean velocity and mean temperature can be simply expressed as:

$$m_s V_{\parallel s,\text{coll}} = P^{-1} (I_4 \times I_1 - I_2 \times I_3), \quad T_{s,\text{coll}} = P^{-1} (I_0 \times I_2 - I_1^2) \quad \text{and} \quad P = I_0 \times I_4 - I_1 \times I_3$$

Appendix B.2. Crank-Nicolson scheme for collision operator solving

In the following, the semi-implicit second-order Crank-Nicolson scheme which is implemented in GYSELA to take into account the collisional effects as

$$\partial_t \bar{F}_s = \mathcal{C}_{ss}(\bar{F}_s) = \frac{1}{B_{\parallel s}^*} \partial_{v_{G\parallel}} \left\{ B_{\parallel s}^* \mathcal{D} \partial_{v_{G\parallel}} \bar{F}_s - B_{\parallel s}^* \mathcal{V} \bar{F}_s \right\} \quad (\text{B.8})$$

is detailed. The diffusion term $\mathcal{D}$ is defined as $\mathcal{D} = \mathcal{D}_{\parallel s}(r, v) = \mathcal{A}(r) \left(\frac{\Phi(v) - G(v)}{2v} \right)$ with $\mathcal{A}(r) = 3 \frac{\sqrt{\pi}}{2} \frac{v_{T_{s0}}^3 \epsilon^{3/2}}{q R_0} \nu_*$ while the expression of the drag term $\mathcal{V}$ is given by $\mathcal{V} = \mathcal{V}_{\parallel s}(r, v) = - \frac{(v_{G\parallel} - V_{\parallel s,\text{coll}})}{v_{T_{s,\text{coll}}}^2} \mathcal{D}_{\parallel s}(r, v)$. Classically, let us write: $f_j^n \equiv \bar{F}_s(r, \theta, \varphi, v_{G\parallel,j}, t_n)$, where subscript $j \in \{0, \dots, N\}$ refers to the discretised index in parallel

velocity space and superscript n refers to the time. For each value of μ_s , each term in Eq. (B.8) reads:

$$B_{\parallel s}^* \partial_t f \rightarrow B_{\parallel s, j}^* \frac{f_j^{n+1} - f_j^n}{\Delta t} \quad (\text{B.9})$$

$$\partial_{v_{G\parallel}} \left(B_{\parallel s}^* \mathcal{D} \partial_{v_{G\parallel}} f \right) \rightarrow \frac{1}{2\Delta v_{G\parallel}^2} \left\{ B_{\parallel s, j+\frac{1}{2}}^* \mathcal{D}_{j+\frac{1}{2}} [f_{j+1}^{n+1} - f_j^{n+1} + f_{j+1}^n - f_j^n] \right. \\ \left. - B_{\parallel s, j-\frac{1}{2}}^* \mathcal{D}_{j-\frac{1}{2}} [f_j^{n+1} - f_{j-1}^{n+1} + f_j^n - f_{j-1}^n] \right\} \quad (\text{B.10})$$

$$\partial_{v_{G\parallel}} \left(B_{\parallel s}^* \mathcal{V} f \right) \rightarrow \frac{B_{\parallel s, j+1}^* \mathcal{V}_{j+1} [f_{j+1}^{n+1} + f_{j+1}^n] - B_{\parallel s, j-1}^* \mathcal{V}_{j-1} [f_{j-1}^{n+1} + f_{j-1}^n]}{4\Delta v_{G\parallel}} \quad (\text{B.11})$$

where we use the fact that $B_{\parallel s}^*$ is linear in $v_{G\parallel}$, *i.e* the expression $B_{\parallel s, j\pm\frac{1}{2}}^* = \frac{1}{2} (B_{\parallel s, j}^* + B_{\parallel s, j\pm 1}^*)$ is exact. Then, using expressions (B.9), (B.10) and (B.11) in the equation (B.8) gives, for all $j = 0, \dots, N$:

$$A_j f_{j-1}^{n+1} + B_j f_j^{n+1} + C_j f_{j+1}^{n+1} = -A_j f_{j-1}^n + \left(\frac{2B_{\parallel s, j}^*}{\Delta t} - B_j \right) f_j^n - C_j f_{j+1}^n \quad (\text{B.12})$$

where the coefficients A_j , B_j and C_j are defined as:

$$A_j = -\frac{\alpha_{j-\frac{1}{2}}}{2\Delta v_{G\parallel}^2} - \frac{\beta_{j-1}}{4\Delta v_{G\parallel}} \quad ; \quad B_j = \frac{B_{\parallel s, j}^*}{\Delta t} + \frac{\alpha_{j+\frac{1}{2}} + \alpha_{j-\frac{1}{2}}}{2\Delta v_{G\parallel}^2} \quad ; \quad C_j = -\frac{\alpha_{j+\frac{1}{2}}}{2\Delta v_{G\parallel}^2} + \frac{\beta_{j+1}}{4\Delta v_{G\parallel}} \quad (\text{B.13})$$

with

$$\alpha_j = B_{\parallel s, j}^* \mathcal{D}_j \quad \text{and} \quad \beta_j = B_{\parallel s, j}^* \mathcal{V}_j \quad (\text{B.14})$$

To compute the terms $\alpha_{-1/2}$, β_{-1} , $\alpha_{N+1/2}$ and β_{N+1} required for A_0 , B_0 , B_N and C_N calculations, the boundary conditions $\partial^2 \mathcal{D} / \partial v_{G\parallel}^2 = 0$ and $\partial^2 \mathcal{V} / \partial v_{G\parallel}^2 = 0$ are imposed, *i.e* :

$$\mathcal{D}_{-\frac{1}{2}} = 2\mathcal{D}_0 - \mathcal{D}_{\frac{1}{2}} \quad \text{and} \quad \mathcal{V}_{-1} = 2\mathcal{V}_0 - \mathcal{V}_1 \quad (\text{B.15})$$

$$\mathcal{D}_{N+\frac{1}{2}} = 2\mathcal{D}_N - \mathcal{D}_{N-\frac{1}{2}} \quad \text{and} \quad \mathcal{V}_{N+1} = 2\mathcal{V}_N - \mathcal{V}_{N-1} \quad (\text{B.16})$$

and the fact that $B_{\parallel s}^*$ is linear in $v_{G\parallel}$ is used, *i.e* :

$$B_{\parallel s, -\frac{1}{2}}^* = \frac{1}{2} (3B_{\parallel s, 0}^* - B_{\parallel s, \frac{1}{2}}^*) \quad ; \quad B_{\parallel s, -1}^* = 2B_{\parallel s, 0}^* - B_{\parallel s, 1}^* \quad (\text{B.17})$$

$$B_{\parallel s, N+\frac{1}{2}}^* = \frac{1}{2} (3B_{\parallel s, N}^* - B_{\parallel s, N-\frac{1}{2}}^*) \quad ; \quad B_{\parallel s, N+1}^* = 2B_{\parallel s, N}^* - B_{\parallel s, N-1}^* \quad (\text{B.18})$$

Finally, let us consider that the second derivative of f is vanishing at domain boundaries, *i.e* $f_{-1} = 2f_0 - f_1$ and $f_{N+1} = 2f_N - f_{N-1}$ then

$$(B_0 + 2A_0)f_0^{n+1} + (C_0 - A_0)f_1^{n+1} = \left(\frac{2B_{\parallel s, 0}^*}{\Delta t} - B_0 - 2A_0 \right) f_0^n - (C_0 - A_0)f_1^n \\ (A_N - C_N)f_{N-1}^{n+1} + (B_N + 2A_N)f_N^{n+1} = -(A_N - C_N)f_{N-1}^n + \left(\frac{2B_{\parallel s, N}^*}{\Delta t} - B_N - 2C_N \right) f_N^n$$

The system can be rewritten in the compact tridiagonal form:

$$\begin{pmatrix} B_0 + 2A_0 & C_0 - A_0 & & & \\ A_1 & B_1 & & \ddots & \\ & \ddots & \ddots & \ddots & \\ & & A_N - C_N & B_N + 2A_N & \end{pmatrix} \begin{pmatrix} f_0^{n+1} \\ \vdots \\ \vdots \\ f_N^{n+1} \end{pmatrix} = \begin{pmatrix} R_0^n \\ \vdots \\ \vdots \\ R_N^n \end{pmatrix} \quad (\text{B.19})$$

and $(R_0^n, \dots, R_N^n)^t = \mathcal{R}(f_0^n, \dots, f_N^n)^t$ with the matrix $\mathcal{R}$ defined as

$$\mathcal{R} = \begin{pmatrix} -2A_0 - B_0 + \frac{2B_{\parallel s,0}^*}{\Delta t} & -(C_0 - A_0) & & & \\ -A_1 & -B_1 + \frac{2B_{\parallel s,1}^*}{\Delta t} & & \ddots & \\ & \ddots & \ddots & \ddots & \\ & & & -C_{N-1} & \\ & & -(A_N - C_N) & -2C_N - B_N + \frac{2B_{\parallel s,N}^*}{\Delta t} & \end{pmatrix} \quad (\text{B.20})$$

The tridiagonal system (B.19)-(B.20) is solved by using a modified Thomas algorithm.

Appendix C. Expressions for simplified collisional transfer between two species

Appendix C.1. Conservation properties of collisional energy transfer

Let us consider the energy exchange between two species defined by equations (34)-(39), i.e. $d\bar{F}_s/dt = C_{ss'}^\mathcal{E}(\mathcal{E}_{ss'})$ and $d\bar{F}_{s'}/dt = C_{s's}^\mathcal{E}(\mathcal{E}_{s's})$ with

$$C_{ss'}^\mathcal{E}(\mathcal{E}_{ss'}) \equiv -\frac{\eta_{ss'}^\mathcal{E}}{(2\pi\bar{T}/m_s)^{3/2}} \frac{\Delta T_{ss'}}{\bar{T}} \left(\frac{\mathcal{E}_{ss'}}{\bar{T}} - \frac{3}{2} \right) \exp\left(-\frac{\mathcal{E}_{ss'}}{\bar{T}}\right)$$

$$C_{s's}^\mathcal{E}(\mathcal{E}_{s's}) \equiv -\frac{\eta_{ss'}^\mathcal{E}}{(2\pi\bar{T}/m_{s'})^{3/2}} \frac{\Delta T_{s's}}{\bar{T}} \left(\frac{\mathcal{E}_{s's}}{\bar{T}} - \frac{3}{2} \right) \exp\left(-\frac{\mathcal{E}_{s's}}{\bar{T}}\right)$$

with $\mathcal{E}_{ss'}$, $\mathcal{E}_{s's}$ defined by equation (38) and $\bar{T}$, $\Delta T_{ss'}$ defined by equation (36). Let us consider the brackets $\langle \cdot \rangle_s$ defined as

$$\langle \cdot \rangle_{v,T} = \frac{1}{(2\pi T/m_s)^{3/2}} \int \int \cdot \exp\left(-\frac{v^2}{2T} - \frac{\mu B}{T}\right) \frac{2\pi}{m_s} B_{\parallel s}^* d\mu_s dv \quad (\text{C.1})$$

Let us define $\bar{v}_{\parallel} = v_{G\parallel} - \bar{V}$. Then, using the fact that $\langle 1 \rangle_{\bar{v}_{\parallel}, \bar{T}} = 1$ and $\langle m_s \bar{v}_{\parallel}^2 \rangle_{\bar{v}_{\parallel}, \bar{T}} = \langle \mu_s B \rangle_{\bar{v}_{\parallel}, \bar{T}} = \bar{T}$, it is straightforward to show that such a collision operator conserves both particles and parallel momentum, i.e. that

$$\langle C_{ss'}^\mathcal{E}(\mathcal{E}_s) \rangle = \langle C_{s's}^\mathcal{E}(\mathcal{E}_{s'}) \rangle = 0 \quad \text{and} \quad \langle v_{G\parallel} C_{ss'}^\mathcal{E}(\mathcal{E}_s) \rangle + \langle v_{G\parallel} C_{s's}^\mathcal{E}(\mathcal{E}_{s'}) \rangle = 0 \quad \text{where} \quad \langle \cdot \rangle = \int \cdot d^3\mathbf{v}$$

The parameters $\eta_{ss'}^\mathcal{E}$ and $\eta_{s's}^\mathcal{E}$ are designed such that the collisional energy exchange between species satisfy $\langle \mathcal{E}_s C_{ss'}^\mathcal{E}(\mathcal{E}_s) \rangle = -\langle \mathcal{E}_{s'} C_{s's}^\mathcal{E}(\mathcal{E}_{s'}) \rangle$. So let us compute the following integral in velocity space,

$$\langle \mathcal{E}_s C_{ss'}^\mathcal{E} \rangle = \int \mathcal{E}_s C_{ss'}^\mathcal{E} d^3\mathbf{v} = -\eta_{ss'}^\mathcal{E} \frac{\Delta T_{ss'}}{\bar{T}} \left\langle \mathcal{E}_s \left(\frac{\mathcal{E}_{ss'}}{\bar{T}} - \frac{3}{2} \right) \right\rangle_{\bar{v}_{\parallel}, \bar{T}}$$

which is equivalent to $\langle \mathcal{E}_s C_{ss'}^{\mathcal{E}} \rangle = -\eta_{ss'}^{\mathcal{E}} (\Delta T_{ss'}/\bar{T}) \mathcal{I}$, where

$$\begin{aligned} \mathcal{I} &= \left\langle \left(m_s \frac{\bar{v}_{\parallel}^2}{2} + m_s \frac{(V_{\parallel s}^2 - \bar{V}^2)}{2} - m_s (\bar{v}_{\parallel} + \bar{V}) \frac{\Delta V_{ss'}}{2} + \mu_s B \right) \left(\frac{1}{2} m_s \frac{\bar{v}_{\parallel}^2}{\bar{T}} + \frac{\mu_s B}{\bar{T}} - \frac{3}{2} \right) \right\rangle_{\bar{v}_{\parallel}, \bar{T}} \\ &= \left\langle \left(m_s \frac{\bar{v}_{\parallel}^2}{2} - m_s \frac{\Delta V_{ss'}}{2} \bar{v}_{\parallel} + m_s \frac{\Delta V_{ss'}^2}{8} + \mu_s B \right) \left(\frac{1}{2} m_s \frac{\bar{v}_{\parallel}^2}{\bar{T}} + \frac{\mu_s B}{\bar{T}} - \frac{3}{2} \right) \right\rangle_{\bar{v}_{\parallel}, \bar{T}} \\ &= \left\langle \frac{m_s^2 \bar{v}_{\parallel}^4}{4\bar{T}} + \left(m_s \frac{\bar{v}_{\parallel}^2}{2} + \mu_s B \right) \left(m_s \frac{\Delta V_{ss'}^2}{8\bar{T}} - \frac{3}{2} \right) + \frac{m_s \bar{v}_{\parallel}^2 \mu_s B}{\bar{T}} + \frac{\mu_s^2 B^2}{\bar{T}} - m_s \Delta V_{ss'} \frac{3}{16} \right\rangle_{\bar{v}_{\parallel}, \bar{T}} \end{aligned}$$

Then, using the properties, $\langle m_s^2 \bar{v}_{\parallel}^4 \rangle_{\bar{v}_{\parallel}, \bar{T}} = 3\bar{T}^2$ and $\langle \mu_s^2 B^2 \rangle_{\bar{v}_{\parallel}, \bar{T}} = 2\bar{T}^2$ and $\langle m_s \bar{v}_{\parallel}^2 \mu_s B \rangle_{\bar{v}_{\parallel}, \bar{T}} = \bar{T}^2$, the collisional energy exchange between species occurs at the following rate

$$\langle \mathcal{E}_s C_{ss'}^{\mathcal{E}} \rangle = -\eta_{ss'}^{\mathcal{E}} \frac{3\Delta T_{ss'}}{2} \left(1 - \frac{\Delta V_{ss'}^2}{8\bar{T}} \right) \quad (\text{C.2})$$

Let us use the property that the energy exchange term between two species is of the form (cf. [51], p.184)

$$W_{ss'} = -\frac{4}{\sqrt{\pi}} \frac{\gamma_{ss'}}{m_{s'}} \frac{n_s n_{s'} (T_s - T_{s'})}{(v_{Ts}^2 + v_{Ts'}^2)^{3/2}} \equiv \int d^3 \mathbf{v} \frac{1}{2} m_s v^2 C_{ss'} \quad \text{with} \quad \gamma_{ss'} \sim \frac{m_s v_{T>} v_{Ts}^2}{n_{s'}} \nu_{ss'}$$

where $v_{T>} = \max(v_{Ts}, v_{Ts'})$ and the collision frequency $\nu_{ss'}$ is deduced from the ion-ion collision frequency ν_{ss} as

$$\nu_{ss'} = \left(\frac{Z_{s'}}{Z_s} \right)^2 \frac{n_{s'}}{n_s} \frac{v_{Ts}}{v_{T>}} \nu_{ss} \quad \text{and} \quad \nu_{ss} = \frac{\varepsilon^{3/2}}{qR_0} \left(\frac{T_{s,\text{coll}}}{m_s} \right)^{1/2} \nu_{*s} \quad (\text{C.3})$$

Therefore

$$W_{ss'} = -\frac{4}{\sqrt{\pi}} \nu_{ss'} \frac{v_{T>}}{v_{Ts}} n_s \frac{m_s}{m_{s'}} \Delta T_{ss'} \left(1 + \frac{v_{Ts'}^2}{v_{Ts}^2} \right)^{-3/2} \quad (\text{C.4})$$

Finally, by analogy between (C.2) and eq. (C.4) (considering $\Delta V_{ss'}^2/8\bar{T} \ll 1$), this provides for the effective collision frequency $\eta_{ss'}^{\mathcal{E}}$ the following expression,

$$\eta_{ss'}^{\mathcal{E}} = \frac{8 \varepsilon^{3/2}}{3\sqrt{\pi}} n_{s'} \frac{m_s}{m_{s'}} \left(\frac{Z_{s'}}{Z_s} \right)^2 \left(1 + \frac{v_{Ts'}^2}{v_{Ts}^2} \right)^{-3/2} \frac{\sqrt{T_{s,\text{coll}}/m_s}}{qR_0} \nu_{*s}$$

which is equivalent to the one given by equation (39).

Appendix C.2. Conservation properties of collisional parallel momentum transfer

Let us consider the parallel momentum transfer defined by equations (40)-(42), i.e. $d\bar{F}_s/dt = C_{ss'}^{v_{G\parallel}}(E_s)$ and $d\bar{F}_{s'}/dt = C_{s's}^{v_{G\parallel}}(E_{s'})$ with

$$C_{ss'}^{v_{G\parallel}}(E_s) \equiv -\eta_{ss'}^{v_{G\parallel}} \Delta V_{ss'} v_{G\parallel} \exp\left(-\frac{E_s}{\bar{T}}\right) \quad \text{and} \quad C_{s's}^{v_{G\parallel}}(E_{s'}) \equiv -\eta_{s's}^{v_{G\parallel}} \Delta V_{s's} v_{G\parallel} \exp\left(-\frac{E_{s'}}{\bar{T}}\right)$$

Considering this approximation, the momentum exchanges between two species reads ⁹

$$\langle m_s v_{G\parallel} C_{ss'}^{v_{G\parallel}}(E_s) \rangle = \int d^3\mathbf{v} m_s v_{G\parallel} \frac{d\bar{F}_s}{dt} = -\eta_{ss'}^{v_{G\parallel}} \Delta V_{ss'} \left(\frac{2\pi}{m_s} \right)^{3/2} \bar{T}^{5/2} \quad (\text{C.5})$$

$$\langle m_s v_{G\parallel} C_{s's}^{v_{G\parallel}}(E_{s'}) \rangle = \int d^3\mathbf{v} m_{s'} v_{G\parallel} \frac{d\bar{F}_{s'}}{dt} = -\eta_{s's}^{v_{G\parallel}} \Delta V_{s's} \left(\frac{2\pi}{m_{s'}} \right)^{3/2} \bar{T}^{5/2} \quad (\text{C.6})$$

Therefore, the action-reaction principle $\langle m_s v_{G\parallel} C_{ss'}^{v_{G\parallel}} \rangle = -\langle m_s v_{G\parallel} C_{s's}^{v_{G\parallel}} \rangle$, leads to the first constraint $\eta_{s's}^{v_{G\parallel}} = (m_{s'}/m_s)^{3/2} \eta_{ss'}^{v_{G\parallel}}$. The second constraint comes from the neoclassical friction relation,

$$-\int d^3\mathbf{v} m_{s'} v_{G\parallel} \frac{d\bar{F}_{s'}}{dt} = m_s n_s \nu_{ss'} \Delta V_{ss'} = -m_{s'} n_{s'} \nu_{s's} \Delta V_{s's}$$

then using equation (C.5) leads to

$$\eta_{ss'}^{v_{G\parallel}} = m_s n_s \left(\frac{2\pi}{m_s} \right)^{-3/2} \bar{T}^{-5/2} \nu_{ss'}$$

Finally, using the relation (C.3) for $\nu_{ss'}$ expression, $\eta_{ss'}^{v_{G\parallel}}$ can be expressed as

$$\eta_{ss'}^{v_{G\parallel}} = n_{s'} m_s^2 \left(\frac{\varepsilon}{2\pi} \right)^{3/2} \frac{1}{qR_0} \frac{\sqrt{\bar{T}_{s,\text{coll}}}}{\bar{T}^{5/2}} \left(\frac{Z_{s'}}{Z_s} \right)^2 \frac{v_{T_s}}{v_{T_{>}}} \nu_{*s}$$

which is equivalent to equation (42).

AppendixD. Source terms

Focussing on the source term, the gyrokinetic equation reads:

$$\frac{d\bar{F}_s}{dt} = S_E(r, \theta, v_{G\parallel}, \mu_s) S_r(r) \quad (\text{D.1})$$

The prescribed radial source profile $S_r(r)$ is defined as the sum of two hyperbolic tangents

$$S_r(r) = -\frac{1}{2} \left[\tanh \left(\frac{\rho - \rho_S - 3L_S}{L_S} \right) + \tanh \left(\frac{\rho_S - 3L_S - \rho}{L_S} \right) \right] \quad (\text{D.2})$$

where ρ_S and L_S are input data and $\rho = (r - r_{\min})/L_r$ where L_r is the length of the radial box. The energy dependent part of the source is decomposed on the basis of orthogonal Hermite and Laguerre polynomials (cf. next section AppendixD.1 as a reminder):

$$S_E(r, \theta, v_{G\parallel}, \mu_s) = \sum_{\ell=0}^{+\infty} \sum_{h=0}^{+\infty} c_{h\ell} H_h(\bar{v}_{G\parallel s}) L_\ell(\bar{\mu}_s) e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s} \quad (\text{D.3})$$

where the $c_{h\ell}$ coefficients depend on the space coordinates only. The following definitions have been introduced:

$$\bar{\mu}_s \equiv \frac{\mu_s B}{T_{s,\text{srce}}} \quad ; \quad \bar{v}_{G\parallel s} \equiv \frac{v_{G\parallel}}{\sqrt{2T_{s,\text{srce}}/m_s}} \quad (\text{D.4})$$

⁹ $\langle m_s v_{G\parallel} C_{ss'}^{v_{G\parallel}} \rangle = -\eta_{ss'}^{v_{G\parallel}} \Delta V_{ss'} \left(\frac{2\pi\bar{T}}{m_s} \right)^{3/2} \langle m_s v_{G\parallel}^2 \rangle_{v_{G\parallel}, \bar{T}} = -\eta_{ss'}^{v_{G\parallel}} \Delta V_{ss'} \left(\frac{2\pi\bar{T}}{m_s} \right)^{3/2} \bar{T}$

with $T_{s,\text{srce}}$ the normalized source temperature.

Appendix D.1. Hermite and Laguerre polynomials

The Hermite and Laguerre polynomials form the set of orthogonal basis with respect to the following scalar products:

$$\text{Laguerre } L_\ell(x) : \int_0^{+\infty} L_\ell L_{\ell'} e^{-x} dx = \delta_{\ell\ell'} |L_\ell|^2 \quad (\text{D.5})$$

$$\text{Hermite } H_h(x) : \int_{-\infty}^{+\infty} H_h H_{h'} e^{-x^2} dx = \delta_{hh'} |H_h|^2 \quad (\text{D.6})$$

The Laguerre polynomials are normalized: $|L_\ell|^2 = 1$. The norm of the Hermite polynomials is:

$$|H_h|^2 \equiv \int_{-\infty}^{+\infty} H_h^2 e^{-x^2} dx = \sqrt{\pi} 2^h h! \quad (\text{D.7})$$

The five first Laguerre and Hermite polynomials are:

$$\begin{array}{llll} L_0(x) = 1 & H_0(x) = 1 & \rightarrow & |H_0|^2 = \sqrt{\pi} \\ L_1(x) = 1 - x & H_1(x) = 2x & \rightarrow & |H_1|^2 = 2\sqrt{\pi} \\ L_2(x) = \frac{1}{2}(2 - 4x + x^2) & H_2(x) = -2 + 4x^2 & \rightarrow & |H_2|^2 = 8\sqrt{\pi} \\ L_3(x) = \frac{1}{6}(6 - 18x + 9x^2 - x^3) & H_3(x) = -12x + 8x^3 & \rightarrow & |H_3|^2 = 48\sqrt{\pi} \\ L_4(x) = \frac{1}{24}(24 - 96x + 72x^2 - 16x^3 + x^4) & H_4(x) = 12 - 48x^2 + 16x^4 & \rightarrow & |H_4|^2 = 384\sqrt{\pi} \end{array} \quad (\text{D.8})$$

Appendix D.2. Corresponding sources for fluid moments

Let us derive the corresponding source of matter $\mathbb{S}_n$, of parallel momentum $\mathbb{S}_{v_{G\parallel}}$, of energy $\mathbb{S}_\mathcal{E}$ and of vorticity $\mathbb{S}_\Omega$. With the adopted definitions,

$$B_{\parallel s}^* = B(1 + J_{\parallel B} \bar{v}_{G\parallel s}) \quad \text{and} \quad J_{\parallel B} \equiv \frac{\sqrt{m_s} \sqrt{2T_{s,\text{srce}}}}{q_s B^2} \mu_0 J_{\parallel} \quad (\text{D.9})$$

The parallel current $\mu_0 J_{\parallel} = \mu_0 \mathbf{b} \cdot \mathbf{J}$ is prescribed in GYSELA. Then, the integral over the velocity space reads as follows:

$$\begin{aligned} \int d^3\mathbf{v} &\equiv \int_{-\infty}^{+\infty} dv_{G\parallel} \int_0^{+\infty} \frac{2\pi B_{\parallel s}^*}{m_s} d\mu_s \\ &= \frac{1}{\sqrt{\pi}} \left(\frac{2\pi T_{s,\text{srce}}}{m_s} \right)^{3/2} \int_{-\infty}^{+\infty} (1 + J_{\parallel B} \bar{v}_{G\parallel s}) dv_{G\parallel s} \int_0^{+\infty} d\bar{\mu}_s \end{aligned}$$

Notice that $(1 + J_{\parallel B} \bar{v}_{G\parallel s}) = H_0(\bar{v}_{G\parallel s}) + \frac{J_{\parallel B}}{2} H_1(\bar{v}_{G\parallel s})$ and $L_0(\bar{\mu}_s) = 1$. The fluid source of matter $\mathbb{S}_n$ is simply $\mathbb{S}_n \equiv \int d^3\mathbf{v} S_E S_r$. Using the decomposition of S_E on the basis of orthogonal polynomials (D.3), the source of matter becomes

$$\mathbb{S}_n = \left(\frac{2\pi T_{s,\text{srce}}}{m_s} \right)^{3/2} (c_{00} + J_{\parallel B} c_{10}) S_r \quad (\text{D.10})$$

The fluid source of parallel momentum $\mathbb{S}_{v_{G\parallel}}$ reads as follows: $\mathbb{S}_{v_{G\parallel}} \equiv \int d^3\mathbf{v} v_{G\parallel} S_E S_r$. Following the same procedure than for the density source, one finally obtains:

$$\mathbb{S}_{v_{G\parallel}} = 2\pi^{3/2} \left(\frac{T_{s,\text{srce}}}{m_s} \right)^2 [2c_{10} + J_{\parallel B} (c_{00} + 4c_{20})] S_r \quad (\text{D.11})$$

The fluid source of energy $\mathbb{S}_{\mathcal{E}}$ is defined as follows: $\mathbb{S}_{\mathcal{E}} \equiv \int d^3\mathbf{v} \left(m_s \frac{v_{G\parallel}^2}{2} + \mu_s B \right) S_E S_r$. Notice that $(m_s v_{G\parallel}^2/2 + \mu_s B) = T_{s,\text{srce}}(\bar{v}_{G\parallel}^2 + \bar{\mu}_s)$. Again, the energy source can also be recast in terms of the $c_{h\ell}$ coefficients:

$$\mathbb{S}_{\mathcal{E}} = \left(\frac{2\pi T_{s,\text{srce}}}{m_s} \right)^{3/2} T_{s,\text{srce}} \left[2 c_{20} + \frac{3}{2} c_{00} - c_{01} + \frac{5}{2} J_{\parallel B} c_{10} + 6 J_{\parallel B} c_{30} - J_{\parallel B} c_{11} \right] S_r \quad (\text{D.12})$$

Appendix D.2.1. Source of vorticity

The fluid source of vorticity $\mathbb{S}_{\Omega}$ is simply: $\mathbb{S}_{\Omega} \equiv \int d^3\mathbf{v} J_0 (S_E S_r)$, where J_0 is the gyro-average operator. We use the Padé approximation:

$$J_0 \approx 1 + \frac{m_s \mu_s}{q_s^2 2B} \nabla_{\perp}^2 = 1 + \alpha \bar{\mu}_s \nabla_{\perp}^2 \quad \text{with} \quad \alpha = \frac{m_s T_{s,\text{srce}}}{q_s^2 2B^2}$$

Again, the vorticity source can be recast in terms of the $c_{h\ell}$ coefficients as

$$\mathbb{S}_{\Omega} = \mathbb{S}_n + \alpha \left(\frac{2\pi T_{s,\text{srce}}}{m_s} \right)^{3/2} \left[\nabla_{\perp}^2 ((c_{00} - c_{01}) S_r) + J_{\parallel B} \nabla_{\perp}^2 ((c_{10} - c_{11}) S_r) \right] \quad (\text{D.13})$$

Appendix D.3. Pure sources of momentum, energy and vorticity

The expressions of $\mathbb{S}_n$ (eq. (D.10)), $\mathbb{S}_{v_{G\parallel}}$ (eq. D.11), $\mathbb{S}_{\mathcal{E}}$ (eq. (D.12)) and $\mathbb{S}_{\Omega}$ (eq. (D.13)) provide the constraints on the $c_{h\ell}$ coefficients in order to impose independently zero source of density, of momentum, of energy or of vorticity. Let us consider three cases:

- Non vanishing source of energy, with no injection of particles nor of momentum.
- Non vanishing source of momentum, with no injection of particles nor of energy.
- Non vanishing source of vorticity, with no injection of particles, of momentum nor of energy.

These three cases are considered hereafter. Imposing zero source of matter, the fluid sources of parallel momentum, of energy and of vorticity are proportional to:

$$\mathbb{S}_n = 0 \Rightarrow c_{00} + J_{\parallel B} c_{10} = 0 \quad (\text{D.14})$$

$$\mathbb{S}_{v_{G\parallel}} \propto \left(2 - J_{\parallel B}^2 \right) c_{10} + 4 J_{\parallel B} c_{20} \quad (\text{D.15})$$

$$\mathbb{S}_{\mathcal{E}} \propto 2c_{20} - c_{00} + 6J_{\parallel B} c_{30} - c_{01} - J_{\parallel B} c_{11} \quad (\text{D.16})$$

$$\mathbb{S}_{\Omega} \propto -\nabla_{\perp}^2 (J_{\parallel B} c_{10} S_r) + J_{\parallel B} \nabla_{\perp}^2 (c_{10} S_r) - \nabla_{\perp}^2 (c_{01} S_r) - J_{\parallel B} \nabla_{\perp}^2 (c_{11} S_r) \quad (\text{D.17})$$

Appendix D.3.1. Pure source of energy

Killing the fluid sources of particles, momentum and vorticity, while keeping finite the fluid source of energy, imposes eq. (D.15) and eq (D.17) to vanish. Several solutions can be envisaged. Let's choose $c_{30} = 0$ and $(c_{01} + J_{\parallel B} c_{11}) = 0$, with $(2c_{20} - c_{00}) \neq 0$. Then, the fluid source of vorticity trivially vanishes for $c_{10} = c_{11}$. To summarize, we propose the following set of coefficients for a pure source of energy,

$$c_{11} = c_{10} = -\frac{4J_{\parallel B}}{2 - J_{\parallel B}^2} c_{20} \quad ; \quad c_{00} = c_{01} = \frac{4J_{\parallel B}^2}{2 - J_{\parallel B}^2} c_{20} \quad ; \quad c_{30} = 0 \text{ and } c_{20} \neq 0 \quad (\text{D.18})$$

In order to inject solely energy into the system, the source term that should appear in the right hand side

of the gyrokinetic equation would then take the following form:

$$\begin{aligned} S_{\mathcal{E}} &= S_r (c_{00} + c_{01}L_1 + c_{10}H_1 + c_{11}H_1L_1 + c_{20}H_2) e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s} \\ &= 4 c_{20} S_r \left[\bar{v}_{G\parallel s}^2 - \frac{1}{2} + \frac{J_{\parallel B}}{2 - J_{\parallel B}^2} (2 - \bar{\mu}_s) (J_{\parallel B} - 2\bar{v}_{G\parallel s}) \right] e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s} \quad \text{due to (D.18)} \end{aligned}$$

while the fluid source of energy $\mathbb{S}_{\mathcal{E}}$ would have the following magnitude:

$$\begin{aligned} \mathbb{S}_{\mathcal{E}} &= \left(\frac{2\pi T_{s,\text{srce}}}{m_s} \right)^{3/2} T_{s,\text{srce}} (2 c_{20} - c_{00}) S_r \quad \text{according to (D.12) and (D.18)} \\ &= 2 \left(\frac{2\pi T_{s,\text{srce}}}{m_s} \right)^{3/2} T_{s,\text{srce}} \left(1 - \frac{2J_{\parallel B}^2}{2 - J_{\parallel B}^2} \right) c_{20} S_r \end{aligned}$$

Let us introduce the normalized intensity $S_0^{\mathcal{E}} \equiv 2 c_{20} \left(\frac{2\pi T_{s,\text{srce}}}{m_s} \right)^{3/2} T_{s,\text{srce}}$ then the previous equality becomes

$$\mathbb{S}_{\mathcal{E}} = \left(1 - \frac{2J_{\parallel B}^2}{2 - J_{\parallel B}^2} \right) S_0^{\mathcal{E}} S_r \quad (\text{D.19})$$

Then, up to small terms proportional to $J_{\parallel B}$, S_r –which is normalized such that its volume integral is equal to unity– provides the radial shape of the energy source, while $S_0^{\mathcal{E}}$ gives its magnitude. Finally, $S_{\mathcal{E}}$ can be expressed as:

$$S_{\mathcal{E}} = \frac{S_0^{\mathcal{E}} S_r}{\sqrt{2} \left(\frac{\pi T_{s,\text{srce}}}{m_s} \right)^{3/2} T_{s,\text{srce}}} \left[\bar{v}_{G\parallel s}^2 - \frac{1}{2} - \frac{J_{\parallel B}}{2 - J_{\parallel B}^2} (2 - \bar{\mu}_s) (2\bar{v}_{G\parallel s} - J_{\parallel B}) \right] e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s} \quad (\text{D.20})$$

with $J_{\parallel B}$ defined by (D.9).

Appendix D.3.2. Pure parallel momentum

Killing the fluid sources of particles, energy and vorticity, while keeping finite the fluid source of parallel momentum, imposes equations (D.16)-(D.17) to vanish. Again, several options could be considered. Following the same strategy as for the energy, namely $c_{11} = c_{10}$ and $c_{00} = c_{01} = -J_{\parallel B} c_{10}$, then the source of energy vanishes if $c_{30} = 0$ and $2 c_{20} = c_{00}$. Consistently, in order to inject only parallel momentum, the following source is proposed

$$S_{v_{G\parallel}} = \frac{S_0^{v_{G\parallel}} S_r}{4\pi^{3/2} \left(\frac{T_{s,\text{srce}}}{m_s} \right)^2} \left[2\bar{v}_{G\parallel s} (2 - \bar{\mu}_s) - J_{\parallel B} \left(1 + 2\bar{v}_{G\parallel s}^2 - \bar{\mu}_s \right) \right] e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s} \quad (\text{D.21})$$

where the normalized intensity $S_0^{v_{G\parallel}}$ is defined as $S_0^{v_{G\parallel}} \equiv 4\pi^{3/2} \left(\frac{T_{s,\text{srce}}}{m_s} \right)^2 c_{10}$. The corresponding fluid source of momentum is (according to (D.11))

$$\mathbb{S}_{v_{G\parallel}} = \left(1 - \frac{3J_{\parallel B}^2}{2} \right) S_0^{v_{G\parallel}} S_r \quad (\text{D.22})$$

Then, at leading order in $J_{\parallel B}$, S_r corresponds to the radial shape of the momentum source, and $S_0^{v_{G\parallel}}$ to its magnitude.

Appendix D.3.3. Pure source of vorticity

So as to inject vorticity only, the simplest choice appears to be: $c_{00} = c_{10} = c_{11} = 0$ and $c_{20} = 0$. Then the source of vorticity is governed by the c_{01} coefficient only: $\mathbb{S}_\Omega \propto -\nabla_\perp^2 (c_{01} S_r)$, while that of momentum eq. (D.15) is set to zero. The source of energy eq. (D.16) vanishes provided that $c_{30} = (1/6 J_{\parallel B}) c_{01}$. Obviously, such a constraint is invalid for those simulations performed at vanishing parallel current. Alternatively, one decides to allow for some parallel momentum injection by taking $c_{20} = c_{01}/2 \neq 0$ and $c_{30} = 0$. Then, the source term to be considered is the following

$$S_\Omega = -\frac{S_0^\Omega S_r}{\alpha \left(\frac{2\pi T_{s,\text{srce}}}{m_s} \right)^{3/2}} \left[2\bar{v}_{G\parallel s}^2 - \bar{\mu}_s \right] e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s} \quad \text{with } \alpha = \frac{m_s}{q_s^2} \frac{T_{s,\text{srce}}}{2B^2} \quad (\text{D.23})$$

where the normalized intensity S_0^Ω is defined by $S_0^\Omega \equiv -c_{01} \alpha \left(\frac{2\pi T_{s,\text{srce}}}{m_s} \right)^{3/2}$ and the resulting fluid source of vorticity is

$$\mathbb{S}_\Omega = S_0^\Omega \nabla_\perp^2 (S_r) \quad (\text{D.24})$$

We recall that such a source does inject some momentum as well. However, its magnitude remains small, and equal to $(-\frac{1}{\alpha} \sqrt{\frac{2m_s}{T_s}} J_{\parallel B} S_0^\Omega S_r)$.

Appendix D.3.4. Another possibility for the heating source

There exists a simplified version of the heating source possible in the code which is not exactly a pure source of heating. In this case the energy dependent part of the source is defined as:

$$S_\mathcal{E}(r, \theta, v_{G\parallel}, \mu_s) = \frac{S_0^\mathcal{E}}{3\sqrt{2} \left(\frac{\pi T_{s,\text{srce}}}{m_s} \right)^{3/2} T_{s,\text{srce}}} \left(\frac{E}{T_{s,\text{srce}}} - \frac{3}{2} \right) \exp \left(-\frac{E}{T_{s,\text{srce}}} \right) \quad (\text{D.25})$$

where E the energy is equal to $E = \frac{1}{2} m_s v_{G\parallel}^2 + \mu_s B$. Therefore,

$$\int_{-\infty}^{\infty} dv_{G\parallel} \int_0^\infty \frac{2\pi B_{\parallel s}^*}{m_s} d\mu_s E S_E = S_0^\mathcal{E}$$

Appendix D.4. A source of impurities

According to equation (D.13), it is clear that in this source formalism it is complicate to generate a source of matter without injecting a source of vorticity. So as first choice, we take $c_{00} = c_{01}$ and $c_{10} = c_{11} = 0$, such that $\mathbb{S}_\Omega = \mathbb{S}_n$. The injection of momentum and energy can be avoided by imposing in equations (D.11) and (D.12), $c_{00} + 4c_{20} = 0$ and $2c_{20} + (3/2)c_{00} - c_{01} + 6J_{\parallel B} c_{30} = 0$. Due to the previous assumptions, this leads to $c_{30} = 0$. Therefore, using the fact that $c_{01} = c_{00}$ and $c_{20} = -(1/4)c_{00}$, the matter source term take the form

$$S_n = S_r c_{00} \left(\frac{5}{2} - \bar{\mu}_s - \bar{v}_{G\parallel s}^2 \right) e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s}$$

Then, according to equation (D.10) and considering the normalized matter source $S_0^n = (2\pi T_{s,\text{srce}}/m_s)^{3/2} c_{00}$,

$$S_n = \frac{S_0^n S_r}{\left(\frac{2\pi T_{s,\text{srce}}}{m_s} \right)^{3/2}} \left(\frac{5}{2} - \bar{\mu}_s - \bar{v}_{G\parallel s}^2 \right) e^{-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s} \quad (\text{D.26})$$

Let us denote by S_{n_s} (resp. $S_{n_{s'}}$) the source of matter associated to the majority species s (resp. to the impurity species s'). The injection of impurity must be compensated by the injection of the majority species such that

$$Z_s \int S_{n_s} J_{v_s} dv_{G\parallel} d\mu_s + Z_{s'} \int S_{n_{s'}} J_{v_{s'}} dv_{G\parallel} d\mu_{s'} = 0$$

Appendix D.5. Numerical treatment of the source terms

The source terms are taken into account by solving $d_t \bar{F}_s = S_{\mathcal{E}} + S_{v_{G\parallel}} + S_{\Omega} + S_n$ with $S_{\mathcal{E}}$ defined by eq. (D.20), $S_{v_{G\parallel}}$ by eq. (D.21), S_{Ω} by eq. (D.23) and S_n by eq. (D.26). For the following, let us use the fact that each source is of the form $S_E(r(t), \theta(t), v_{G\parallel}(t), \mu_s) S_r(r(t))$. For more readability let us consider one unique source knowing that the numerical method described below can be generalized to a sum of sources. Let us integrate in time between t and $t + \Delta t$ the equation

$$\frac{d\bar{F}_s}{dt}(r, \theta, \varphi, v_{G\parallel}, \mu) = S(t) \quad \text{with} \quad S(t) = S_E(r(t), \theta(t), v_{G\parallel}(t), \mu_s) S_r(r(t))$$

then

$$\bar{F}_s(t + \Delta t) - \bar{F}_s(t) = \int_t^{t+\Delta t} S(t') dt'$$

$$\begin{aligned} S_r(r(t')) &= S_r(r(t + \Delta t)) + [r(t') - r(t + \Delta t)] \frac{\partial S_r}{\partial r}(r(t + \Delta t)) \\ &= S_r(r(t + \Delta t)) + (t' - t - \Delta t) \frac{dr}{dt}|_{t+\Delta t} \frac{\partial S_r}{\partial r}(r(t + \Delta t)) + \mathcal{O}(\Delta t^2) \end{aligned}$$

Besides, let us define $S_E(t') = S_E(r(t'), \theta(t'), v_{G\parallel}(t'), \mu_s)$, then

$$\begin{aligned} S_E(t') &= S_E(r(t + \Delta t), \theta(t'), v_{G\parallel}(t'), \mu_s) + \\ &\quad (t' - t - \Delta t) \frac{dr}{dt}|_{t+\Delta t} \frac{\partial S_E}{\partial r}(r(t + \Delta t), \theta(t'), v_{G\parallel}(t'), \mu_s) + \mathcal{O}(\Delta t^2) \\ &\approx S_E(r(t + \Delta t), \theta(t + \Delta t), v_{G\parallel}(t + \Delta t), \mu_s) + \\ &\quad (t' - t - \Delta t) \left[\frac{dr}{dt}|_{t+\Delta t} \frac{\partial S_E}{\partial r}(r(t + \Delta t), \theta(t + \Delta t), v_{G\parallel}(t + \Delta t), \mu_s) + \right. \\ &\quad \left. \frac{d\theta}{dt}|_{t+\Delta t} \frac{\partial S_E}{\partial \theta}(r(t + \Delta t), \theta(t + \Delta t), v_{G\parallel}(t + \Delta t), \mu_s) + \right. \\ &\quad \left. \frac{dv_{G\parallel}}{dt}|_{t+\Delta t} \frac{\partial S_E}{\partial v_{G\parallel}}(r(t + \Delta t), \theta(t + \Delta t), v_{G\parallel}(t + \Delta t), \mu_s) \right] + \mathcal{O}(\Delta t^2) \end{aligned}$$

Therefore, the distribution function $\bar{F}_s$ at time $t + \Delta t$ is given at second order in time by

$$\begin{aligned} \bar{F}_s(t + \Delta t) &= \bar{F}_s(t) + \Delta t S_E(t + \Delta t) S_r(t + \Delta t) - \\ &\quad \frac{\Delta t^2}{2} \left[S_E(t + \Delta t) \frac{dr}{dt}|_{t+\Delta t} \frac{\partial S_r}{\partial r}|_{t+\Delta t} + S_r(t + \Delta t) \left(\frac{dr}{dt}|_{t+\Delta t} \frac{\partial S_E}{\partial r}|_{t+\Delta t} + \right. \right. \\ &\quad \left. \left. \frac{d\theta}{dt}|_{t+\Delta t} \frac{\partial S_E}{\partial \theta}|_{t+\Delta t} + \frac{dv_{G\parallel}}{dt}|_{t+\Delta t} \frac{\partial S_E}{\partial v_{G\parallel}}|_{t+\Delta t} \right) \right] \end{aligned}$$

In the code, the radial derivatives of the sources of: (i) energy $\partial_r S_{\mathcal{E}}$, (ii) momentum $\partial_r S_{v_{G\parallel}}$, (iii) vorticity $\partial_r S_{\Omega}$ and (iv) density $\partial_r S_n$ are computed numerically as well as the poloidal derivatives. On the other hand, the derivatives in $v_{G\parallel}$ direction are expressed analytically as

(i) for the pure source of energy:

$$\begin{aligned} \frac{\partial S_{\mathcal{E}}}{\partial v_{G\parallel}} &= \frac{S_0^{\mathcal{E}} S_r}{\sqrt{2} \left(\frac{\pi T_{s, \text{srce}}}{m_s} \right)^{3/2} T_{s, \text{srce}}} \frac{1}{\sqrt{2 T_{s, \text{srce}}}} \exp \left(-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s \right) \times \\ &\quad \left[-2K_h(2 - \bar{\mu}_s) - 2\bar{v}_{G\parallel s} \left\{ \bar{v}_{G\parallel s}^2 - \frac{3}{2} - K_h(2 - \bar{\mu}_s) (2\bar{v}_{G\parallel s} - J_{\parallel B}) \right\} \right] \end{aligned}$$

(ii) for the pure source of momentum:

$$\frac{\partial S_{v_{G\parallel}}}{\partial v_{G\parallel}} = \frac{S_0^{v_{G\parallel}} S_r}{4\pi^{3/2} \left(\frac{T_{s,\text{srce}}}{m_s}\right)^2} \frac{1}{\sqrt{2T_{s,\text{srce}}}} \exp\left(-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s\right) \times \\ \left[2(2 - \bar{\mu}_s) \left(1 - 2\bar{v}_{G\parallel s}^2\right) + 2J_{\parallel B} \bar{v}_{G\parallel s} \left(-1 + 2\bar{v}_{G\parallel s}^2 - \bar{\mu}_s\right)\right]$$

(iii) for the pure source of vorticity (according to eq. (D.23)):

$$\frac{\partial S_{\Omega}}{\partial v_{G\parallel}} = -\frac{S_0^{\Omega} S_r}{\alpha \left(\frac{2\pi T_{s,\text{srce}}}{m_s}\right)^{3/2}} \frac{1}{\sqrt{2T_{s,\text{srce}}}} \exp\left(-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s\right) \left[4\bar{v}_{G\parallel s} - 2\bar{v}_{G\parallel s} \left(2\bar{v}_{G\parallel s}^2 - \bar{\mu}_s\right)\right]$$

(iv) for the source of impurity (according to eq. (D.26)):

$$\frac{\partial S_n}{\partial v_{G\parallel}} = \frac{S_0^n S_r}{\left(\frac{2\pi T_{s,\text{srce}}}{m_s}\right)^{3/2}} \frac{1}{\sqrt{2T_{s,\text{srce}}}} \exp\left(-\bar{v}_{G\parallel s}^2 - \bar{\mu}_s\right) \left[-2\bar{v}_{G\parallel s} \left(\frac{7}{2} - \bar{\mu}_s - \bar{v}_{G\parallel s}^2\right)\right]$$

Appendix E. A Crank-Nicolson scheme for diffusion terms

In this appendix the semi-implicit Crank-Nicolson scheme, which is implemented in GYSELA to take into account a diffusion term of the form

$$\partial_t \bar{F}_s = \frac{1}{r} \partial_r (r \chi(r) \partial_r \bar{F}_s) \quad (\text{E.1})$$

is detailed. Classically, let us write: $f_i^n \equiv \bar{F}_s(r_i, \theta, \varphi, v_{G\parallel}, t_n)$, where subscript $j \in \{0, \dots, N\}$ refers to the discretised index in radial direction and superscript n refers to the time. Let us also define $\mathcal{D} \equiv r \chi(r)$ with $\chi(r) = \chi_0 H_{\text{buff}}(r)$. The diffusion amplitude χ_0 is modulated by a radial profile H_{buff} which is equal to 1 in the buffer region and equal to 0 elsewhere. For each value of μ_s , each term in Eq. (E.1) reads for all $i = 0, \dots, N$:

$$f_i^{n+1} = f_i^n + \frac{\Delta t}{2 \Delta r^2 r_i} \left\{ \mathcal{D}_{i+\frac{1}{2}} [(f_{i+1}^{n+1} - f_i^{n+1}) + (f_{i+1}^n - f_i^n)] \right. \\ \left. - \mathcal{D}_{i-\frac{1}{2}} [(f_i^{n+1} - f_{i-1}^{n+1}) + (f_i^n - f_{i-1}^n)] \right\} \quad (\text{E.2})$$

Let us introduce $\alpha_i = \frac{\Delta t}{2 \Delta r^2 r_i}$, then

$$-\alpha_i \mathcal{D}_{i-\frac{1}{2}} f_{i-1}^{n+1} + \left[1 + \alpha_i \left(\mathcal{D}_{i-\frac{1}{2}} + \mathcal{D}_{i+\frac{1}{2}}\right)\right] f_i^{n+1} - \alpha_i \mathcal{D}_{i+\frac{1}{2}} f_{i+1}^{n+1} = R_i^n$$

with

$$R_i^n = \alpha_i \mathcal{D}_{i-\frac{1}{2}} f_{i-1}^n + \left[1 - \alpha_i \left(\mathcal{D}_{i-\frac{1}{2}} + \mathcal{D}_{i+\frac{1}{2}}\right)\right] f_i^n + \alpha_i \mathcal{D}_{i+\frac{1}{2}} f_{i+1}^n$$

Therefore, for all $i = 0, \dots, N$,

$$A_i f_{i-1}^{n+1} + B_i f_i^{n+1} + C_i f_{i+1}^{n+1} = -A_i f_{i-1}^n + (2 - B_i) f_i^n - C_i f_{i+1}^n \quad (\text{E.3})$$

where the coefficients A_i , B_i and C_i are defined as:

$$A_i = -\alpha_i \mathcal{D}_{i-\frac{1}{2}} \quad ; \quad B_i = 1 + \alpha_i \left(\mathcal{D}_{i-\frac{1}{2}} + \mathcal{D}_{i+\frac{1}{2}}\right) \quad ; \quad C_i = -\alpha_i \mathcal{D}_{i+\frac{1}{2}}$$

Using the fact that $r_{i\pm 1/2} = 0.5 (r_i + r_{i\pm 1})$, then $\mathcal{D}_{i\pm \frac{1}{2}} = \frac{1}{4} (r_i \chi_i + r_{i\pm 1} \chi_i + r_i \chi_{i\pm 1} + r_{i\pm 1} \chi_{i\pm 1})$. Concerning the boundary conditions, let us consider that the second derivative of the radial profile χ is equal to 0 at the boundaries (*i.e* $\frac{\partial^2 \chi}{\partial r^2} |_{r_{\min}} = \frac{\partial^2 \chi}{\partial r^2} |_{r_{\max}} = 0$), then

$$\mathcal{D}_{-\frac{1}{2}} = \frac{1}{2} \left(r_0 - \frac{\Delta r}{2} \right) (3\chi_0 - \chi_1) \quad \text{and} \quad \mathcal{D}_{N+\frac{1}{2}} = \frac{1}{2} \left(r_N + \frac{\Delta r}{2} \right) (3\chi_N - \chi_{N-1})$$

Besides, a non-homogeneous Dirichlet boundary condition is applied to the outer radial boundary such that $\bar{F}_s(r = r_{\max}, \theta, \varphi, v_{G\parallel}, \mu_s) = \bar{F}_{s,\text{eq}}(r = r_{\max}, \theta, v_{G\parallel}, \mu_s)$. Concerning the inner radial boundary, there are two possibility: (i) non-homogeneous Dirichlet boundary condition $\bar{F}_s(r = r_{\min}, \theta, \varphi, v_{G\parallel}, \mu_s) = \bar{F}_{s,\text{eq}}(r = r_{\min}, \theta, v_{G\parallel}, \mu_s)$ in the case of thermal bath or (ii) Neumann boundary conditions (*i.e* $f_{-1} = f_1$) in the case of flux-driven simulation. Let us consider the boolean κ which is equal to 1 if Neumann boundary conditions are applied (*i.e* in the case of flux driven boundary conditions) and is equal to 0 otherwise. The system can be rewritten in the compact tridiagonal form:

$$\begin{pmatrix} (1-\kappa) + \kappa B_0 & \kappa(A_0 + C_0) & & \\ A_1 & B_1 & \ddots & \\ & \ddots & \ddots & C_{N-2} \\ & A_{N-1} & B_{N-1} & C_{N-1} \\ & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} f_0^{n+1} \\ \vdots \\ f_N^{n+1} \end{pmatrix} = \begin{pmatrix} R_0^n \\ \vdots \\ R_N^n \end{pmatrix} \quad (\text{E.4})$$

with

$$\begin{pmatrix} R_0^n \\ \vdots \\ R_N^n \end{pmatrix} = \begin{pmatrix} (1-\kappa) + \kappa(2-B_0) & -\kappa(A_0 + C_0) & & \\ -A_1 & (2-B_1) & \ddots & \\ & \ddots & \ddots & -C_{N-2} \\ & -A_{N-1} & (2-B_{N-1}) & -C_{N-1} \\ & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} f_0^n \\ \vdots \\ f_N^n \end{pmatrix}$$

The tridiagonal system (E.4) is solved by using a modified Thomas algorithm.

Appendix F. GYSELA normalized equations

This appendix contains the expression of all the normalized quantities and normalized equations implemented in the code GYSELA .

Appendix F.1. GYSELA normalization

Let consider $m_0 = A_0 m_p$ (kilograms) a reference ionic mass, $q_0 = Z_0 e$ (coulombs) a reference ionic charge, B_0 (tesla) a reference magnetic induction and T_0 (electron volts) a reference temperature. A_0 and Z_0 are the (dimensionless) mass number and charge state of the principal ionic species and e the modulus of the electron charge. These quantities are used to define the reference ion cyclotron frequency Ω_{c_0} , the reference thermal speed v_{T_0} and the reference Larmor-radius ρ_0 as:

$$\Omega_{c_0} = \frac{Z_0 e B_0}{m_0} \quad ; \quad v_{T_0} = \sqrt{\frac{T_0}{m_0}} \quad ; \quad \rho_0 = \frac{v_{T_0}}{\Omega_{c_0}} = \frac{\sqrt{T_0 m_0}}{Z_0 e B_0} \quad (\text{F.1})$$

Hereafter, $\hat{X}$ refers to the dimensionless expression of the X quantity. The normalizations used in the code are summarized in the table F.12.

$ \begin{aligned} m_s &= m_0 \hat{A}_s \\ q_s &= Z_0 e \hat{Z}_s \\ l &= \rho_0 \hat{l} \\ t &= \hat{t} / \Omega_{c_0} \\ n_s &= n_0 \hat{n}_s \\ T_s &= T_0 \hat{T}_s \\ B &= B_0 \hat{B} \end{aligned} $	$\Rightarrow$	$ \begin{aligned} v_{G\parallel} &= v_{T_{s0}} \hat{v}_{G\parallel s} \\ &= (v_{T_0} / \sqrt{\hat{A}_s}) \hat{v}_{G\parallel s} \\ v_{E \times B_s} &= v_{T_0} \hat{v}_{E \times B_s} \\ v_{D_s} &= v_{T_0} \hat{v}_{D_s} \\ \mu_s &= (T_0 / B_0) \hat{\mu}_s \\ U &= [T_0 / (Z_0 e)] \hat{\phi} \\ E &= (v_{T_0} B_0) \hat{E} \\ F_s &= (n_0 / v_{T_{s0}}^3) \hat{F}_s \end{aligned} $	and	$ \begin{aligned} \chi &= \rho_0^2 \Omega_{c_0} \hat{\chi} \\ \nu &= \Omega_{c_0} \hat{\nu} \\ \mathcal{D}_{\parallel s} &= v_{T_{s0}}^2 \Omega_{c_0} \hat{\mathcal{D}}_{\parallel s} \\ \mathcal{V}_{\parallel s} &= v_{T_{s0}} \Omega_{c_0} \hat{\mathcal{V}}_{\parallel s} \\ S_0^{\mathcal{E}} &= n_0 \Omega_{C_0} T_0 \hat{S}_0^{\mathcal{E}} \\ S_0^{v_{G\parallel}} &= n_0 \Omega_{C_0} v_{T_{s0}} \hat{S}_0^{v_{G\parallel}} \\ S_0^{\Omega} &= \frac{n_0 \Omega_{C_0}}{\rho_0^2} \hat{S}_0^{\Omega} \end{aligned} $
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Table F.12: Links between physical and normalized quantities.

Let us notice that the parallel velocity is normalized to $v_{T_{s0}}$ while all drift velocities are normalized to v_{T_0} . This appendix is divided into two parts. The first part contains a summarize of all the normalized equations used in GYSELA. The second part contains the details of the computation for obtaining these normalized expressions.

Appendix F.2. Summary of equations implemented in the GYSELA code

The equation system –a 5D Vlasov equation self-consistently coupled to a 3D quasi-neutrality equation– treated in the code is summarized in this section. As a reminder, let us define $\mathbf{b} = \mathbf{B} / \|\mathbf{B}\|$ and its covariant components b_k as J_x the jacobian in space of the system, then the Poisson brackets $[\cdot, \cdot]$ are defined as

$$[F, G] = \mathbf{b} \cdot (\nabla F \times \nabla G) = J_x^{-1} \epsilon^{ijk} \partial_i F \partial_j G b_k \quad (\text{F.2})$$

Appendix F.3. Normalized Boltzmann equation

The evolution of the guiding-center distribution function $\hat{F}_s$ is governed by the following normalized Boltzmann equation:

$$\frac{\partial \hat{F}_s}{\partial \hat{t}} + \frac{1}{\hat{B}_{\parallel s}^*} \hat{\nabla} \cdot \left(\hat{B}_{\parallel s}^* \frac{d\hat{\mathbf{x}}_G}{d\hat{t}} \hat{F}_s \right) + \frac{1}{\hat{B}_{\parallel s}^*} \frac{\partial}{\partial \hat{v}_{G\parallel s}} \left(\hat{B}_{\parallel s}^* \frac{d\hat{v}_{G\parallel s}}{d\hat{t}} \hat{F}_s \right) = \hat{\mathcal{D}}_r(\hat{F}_s) + \hat{\mathcal{K}}(\hat{F}_s) + \hat{\mathcal{C}}(\hat{F}_s) + \hat{\mathcal{S}} \quad (\text{F.3})$$

where $\hat{\mathcal{D}}_r$ and $\hat{\mathcal{K}}$ are respectively a diffusion term and a Krook operator applied on a radial buffer region (see section [Appendix F.5](#)), while $\hat{\mathcal{C}}$ refers to a collision operator (see section [Appendix F.6](#)) and $\hat{\mathcal{S}}$ corresponds to source terms (see section [Appendix F.8](#) for detailed expressions). The evolution of the gyro-center coordinates $(\mathbf{x}_G, v_{G\parallel}, \mu_s)$ of species s is given by:

$$\frac{d\hat{x}_G^i}{d\hat{t}} = \frac{1}{\sqrt{\hat{A}_s}} \hat{v}_{G\parallel s} \hat{\mathbf{b}}_s^* \cdot \hat{\nabla} x^i + \hat{\mathbf{v}}_{E \times B_s} \cdot \hat{\nabla} x^i + \hat{\mathbf{v}}_{D_s} \cdot \hat{\nabla} x^i \quad (\text{F.4})$$

$$\frac{d\hat{v}_{G\parallel s}}{d\hat{t}} = -\frac{1}{\sqrt{\hat{A}_s}} \hat{\mu}_s \hat{\mathbf{b}}_s^* \cdot \hat{\nabla} \hat{B} - \frac{\hat{Z}_s}{\sqrt{\hat{A}_s}} \hat{\mathbf{b}}_s^* \cdot \hat{\nabla} \hat{\phi} + K_{\nabla B} \hat{v}_{G\parallel s} \hat{\mathbf{v}}_{E \times B_s} \cdot \frac{\hat{\nabla} \hat{B}}{\hat{B}} \quad (\text{F.5})$$

The i -th covariant coordinates of the normalized drift velocities are given by:

$$\hat{\mathbf{v}}_{E \times B_s} \cdot \hat{\nabla} x^i = \hat{\mathbf{v}}_{E \times B_s}^i = \frac{1}{\hat{B}_{\parallel s}^*} \left[\hat{\phi}, \hat{x}^i \right] \quad (\text{F.6})$$

$$\hat{\mathbf{v}}_{D_s} \cdot \hat{\nabla} x^i = \hat{\mathbf{v}}_{D_s}^i = K_{\nabla B} \left(\frac{\hat{v}_{G\parallel s}^2 + \hat{\mu}_s \hat{B}}{\hat{Z}_s \hat{B}_{\parallel s}^* \hat{B}} \right) \left[\hat{B}, \hat{x}^i \right] \quad (\text{F.7})$$

while $\hat{\mathbf{b}}_s^*$ and $\hat{B}_{\parallel s}^*$ are defined as:

$$\hat{\mathbf{b}}_s^* = \frac{1}{\hat{B}_{\parallel s}^*} \left(\hat{\mathbf{B}} + \frac{\sqrt{\hat{A}_s}}{\hat{Z}_s} \frac{\hat{v}_{G\parallel s}}{\hat{B}} \mu_0 \hat{\mathbf{J}} \right) \quad \text{and} \quad \hat{B}_{\parallel s}^* = \hat{B} + \frac{\sqrt{\hat{A}_s}}{\hat{Z}_s} \frac{\hat{v}_{G\parallel s}}{\hat{B}} \mu_0 \mathbf{b} \cdot \hat{\mathbf{J}} \quad (\text{F.8})$$

The normalized magnetic field $\hat{B}$ is defined as

$$\hat{\mathbf{B}} = \frac{R_0}{\hat{R}(r, \theta)} [\zeta(\hat{r}) \mathbf{e}_\theta + \mathbf{e}_\varphi] \quad \text{with} \quad \zeta(\hat{r}) = \frac{\hat{r}}{qR_0} \quad (\text{F.9})$$

while the normalized current reads

$$\mu_0 \hat{\mathbf{J}} = \mu_0 \hat{J}_T \hat{R} \hat{\nabla} \varphi \quad \text{with} \quad \mu_0 \hat{J}_T = \frac{R_0}{\hat{R}} \frac{\zeta}{\hat{r}} \left(1 + \frac{\hat{r}}{\zeta} \frac{d\zeta}{d\hat{r}} - \frac{\hat{r}}{\hat{R}} \cos \theta \right) \quad (\text{F.10})$$

The constant $K_{\nabla B}$ has been added for tests in the code. $K_{\nabla B}$ is equal to 1 if the curvature of the magnetic field is taken into account and 0 otherwise. In equations (F.4) and (F.5) the parallel projection $\hat{\nabla}_{\parallel}^* F$ is computed as $\hat{\mathbf{b}}_s^* \cdot \hat{\nabla} F = \hat{b}_s^{*i} \partial_i F$. In equations (F.6) and (F.7), the explicit expressions of the Poisson brackets $[\bar{\phi}, x^i]$ and $[B, x^i]$ are given by:

$$\left\{ \begin{array}{l} [\bar{\phi}, x^1] = \frac{1}{J_x B} \{ B_2 \partial_{x^3} \bar{\phi} - B_3 \partial_{x^2} \bar{\phi} \} \\ [\bar{\phi}, x^2] = \frac{1}{J_x B} \{ -B_1 \partial_{x^3} \bar{\phi} + B_3 \partial_{x^1} \bar{\phi} \} \\ [\bar{\phi}, x^3] = \frac{1}{J_x B} \{ B_1 \partial_{x^2} \bar{\phi} - B_2 \partial_{x^1} \bar{\phi} \} \end{array} \right. \quad \text{and} \quad \left\{ \begin{array}{l} [B, x^1] = -\frac{1}{J_x B} \times B_3 \partial_{x^2} B \\ [B, x^2] = \frac{1}{J_x B} \times B_3 \partial_{x^1} B \\ [B, x^3] = \frac{1}{J_x B} (B_1 \partial_{x^2} B - B_2 \partial_{x^1} B) \end{array} \right.$$

Appendix F.4. Normalized quasi-neutrality equation

The evolution of $\hat{F}_s$ is self-consistently coupled to the normalized quasi-neutrality equation ¹⁰:

$$-\frac{1}{\hat{n}_{e0}} \sum_s \hat{A}_s \hat{\nabla}_{\perp} \cdot \left(\frac{\hat{n}_{s0}}{\hat{B}} \hat{\nabla}_{\perp} \hat{\phi} \right) + \frac{1}{Z_0^2 \hat{T}_e} [\hat{\phi} - \lambda \langle \hat{\phi} \rangle_{\text{FS}}] = \frac{1}{\hat{n}_{e0}} \sum_s \hat{Z}_s [\hat{n}_{G_s} - \hat{n}_{G_s, \text{eq}}] \quad (\text{F.11})$$

where the normalized electron density $\hat{n}_{e0}$ is defined as:

$$\hat{n}_{e0} = \sum_s Z_0 \hat{Z}_s \hat{n}_{s0} \approx \sum_s Z_0 \hat{Z}_s \hat{n}_{s, \text{eq}} \approx \sum_s Z_0 \hat{Z}_s \hat{n}_{G_s, \text{eq}} \quad (\text{F.12})$$

The integral $\langle \hat{\phi} \rangle_{\text{FS}} = \int \hat{\phi} \hat{J}_x d\theta d\varphi / \int \hat{J}_x d\theta d\varphi$ represents the flux surface average of $\hat{\phi}$ (with $\hat{J}_x$ being the normalized jacobian space). The parameter λ has been added for test. It can be chosen equal to 1 or 0. The normalized guiding-center density $\hat{n}_{G_s}$ of species s is given by:

$$\hat{n}_{G_s} = \int d\hat{\mu}_s \int \hat{J}_v d\hat{v}_{G\parallel s} \hat{J}_0 \cdot \hat{\hat{F}}_s \quad (\text{F.13})$$

¹⁰In the code, to avoid the expensive gyroaverage operation for each value of $v_{G\parallel}$ —which occur for the computation of the RHS of the quasi-neutrality equation (F.11)—we use the definition (F.8) of $\hat{B}_{\parallel s}^*$ and therefore the fact the term $\hat{n}_{G_s} - \hat{n}_{G_s, \text{eq}}$ can be expressed as:

$$\hat{n}_{G_s} - \hat{n}_{G_s, \text{eq}} = 2\pi \int d\hat{\mu}_s \left(\hat{B} \hat{J}_0 \cdot I_0(\hat{r}, \theta, \varphi, \hat{\mu}_s) + \frac{\sqrt{\hat{A}_s}}{\hat{Z}_s} \frac{\hat{v}_{G\parallel s}}{\hat{B}} \mu_0 \mathbf{b} \cdot \hat{\mathbf{J}} \hat{J}_0 \cdot I_1(\hat{r}, \theta, \varphi, \hat{\mu}_s) \right)$$

where the integrals I_0 and I_1 are defined by:

$$I_0(\hat{r}, \theta, \varphi, \hat{\mu}_s) = \int (\hat{F}_s - \hat{F}_{s, \text{eq}}) d\hat{v}_{G\parallel s} \quad \text{and} \quad I_1(\hat{r}, \theta, \varphi, \hat{\mu}_s) = \int \hat{v}_{G\parallel s} (\hat{F}_s - \hat{F}_{s, \text{eq}}) d\hat{v}_{G\parallel s}$$

with the normalized jacobian in velocity equal to $\hat{J}_v = 2\pi\hat{B}_{\parallel s}^*$. The correction term $\hat{n}_{G_s, \text{eq}}$ in the right hand side is defined such that:

$$\hat{n}_{G_s, \text{eq}} = \int d\hat{\mu}_s \int \hat{J}_v d\hat{v}_{G\parallel s} \hat{J}_0 \cdot \hat{\hat{F}}_{s, \text{eq}} \quad (\text{F.14})$$

where

$$\hat{\hat{F}}_{s, \text{eq}} = c_s \frac{\hat{n}_{s0}}{(2\pi\hat{T}_s)^{3/2}} \exp \left[-\frac{(\hat{v}_{G\parallel s}^2/2 + \hat{\mu}_s \hat{B})}{\hat{T}_s} \right] \quad (\text{F.15})$$

represents the equilibrium part of the distribution function. The concentration c_s is defined such that $\sum_s c_s Z_0 \hat{Z}_s = 1$. The normalized gyroaverage operator $\hat{J}$ approximated by Padé corresponds to $\hat{J}_0 \approx 1 + \frac{1}{2} \frac{\hat{A}_s}{\hat{Z}_s^2} \frac{\hat{\mu}_s}{\hat{B}} \hat{\nabla}_\perp^2$.

Appendix F.5. Normalized diffusion terms and Krook operator in buffer regions

A radial diffusion and an artificial damping are applied in buffer regions. These buffer regions are defined at each extremities of the radial domain $\hat{r} \in [\hat{r}_{\min}, \hat{r}_{\max}]$ as a sum of hyperbolic tangents:

$$\hat{H}_{\text{buff}}(\hat{r}) = 1 + \frac{1}{2} \left[\tanh \left(\frac{\hat{r} - \hat{r}_{\max} + B_L \hat{L}_r}{B_S \hat{L}_r} \right) - \tanh \left(\frac{\hat{r} - \hat{r}_{\min} - B_L \hat{L}_r}{B_S \hat{L}_r} \right) \right] \quad (\text{F.16})$$

where $\hat{L}_r$ is the normalized length of the radial domain. B_L and B_S are respectively the location and the stiffness of the buffer regions. The function $\hat{H}_{\text{buff}}$ plays the role of a mask which is equal to 1 in the buffer regions and equal to 0 elsewhere.

A radial diffusion ¹¹ is applied on this buffer region by solving the following equation

$$\frac{\partial \hat{\hat{F}}_s}{\partial \hat{t}} = \hat{\mathcal{D}}_r(\hat{\hat{F}}_s) \quad \text{with} \quad \hat{\mathcal{D}}_r(\hat{\hat{F}}_s) = \frac{1}{\hat{r} \hat{B}_{\parallel s}^*} \frac{\partial}{\partial \hat{r}} \left(\hat{r} \hat{\chi}(\hat{r}) \frac{\partial}{\partial \hat{r}} (\hat{B}_{\parallel s}^* \hat{\hat{F}}_s) \right) \quad (\text{F.17})$$

where $\hat{\chi}(\hat{r}) = \chi_0 \hat{H}_{\text{buff}}(\hat{r})$, i.e $\hat{\chi}(\hat{r}) = \chi(r)/\rho_0^2 \Omega_{c0}$.

A Krook operator is applied by solving $\partial_{\hat{t}} \hat{\hat{F}}_s = \hat{\mathcal{K}}_r(\hat{\hat{F}}_s)$ with $\hat{\mathcal{K}}_r(\hat{\hat{F}}_s) = -\hat{\nu}(\hat{r})(\hat{\hat{F}}_s - \hat{\hat{F}}_{s, \text{eq}})$ where $\hat{\nu}(\hat{r}) = \nu_0 \hat{H}_{\text{buff}}(\hat{r})$ such that $\hat{\nu}(\hat{r}) = \nu(r)/\Omega_{c0}$.

Appendix F.6. Normalized collision operator

For taking into account the collision effects, a collision operator $\mathcal{C}_s$ is added to the normalized Vlasov equation (F.3), such that the normalized Boltzmann equation reads:

$$\frac{d\hat{\hat{F}}_s}{d\hat{t}} = \hat{\mathcal{C}}_s(\hat{\hat{F}}_s) \quad \text{with} \quad \hat{\mathcal{C}}_s(\hat{\hat{F}}_s) = \frac{1}{\hat{B}_{\parallel s}^*} \frac{\partial}{\partial \hat{v}_{G\parallel s}} \left[\hat{B}_{\parallel s}^* \left(\hat{\mathcal{D}}_{\parallel s} \frac{\partial \hat{\hat{F}}_s}{\partial \hat{v}_{G\parallel s}} - \hat{\mathcal{V}}_{\parallel s} \hat{\hat{F}}_s \right) \right] \quad (\text{F.18})$$

where the diffusion term $\hat{\mathcal{D}}_{\parallel s}$ and the drag term $\hat{\mathcal{V}}_{\parallel s}$ in parallel direction are defined as:

$$\hat{\mathcal{D}}_{\parallel s}(\hat{r}, \hat{v}) = \hat{\mathcal{A}}_s(\hat{r}) \left(\frac{\Phi(\hat{v}) - G(\hat{v})}{2\hat{v}} \right) \quad ; \quad \hat{\mathcal{V}}_{\parallel s}(\hat{r}, \hat{v}) = -\frac{(\hat{v}_{G\parallel s} - \hat{V}_{\parallel s, \text{coll}})}{\hat{v}_{T_s, \text{coll}}^2} \hat{\mathcal{D}}_{\parallel s}(\hat{r}, \hat{v}) \quad (\text{F.19})$$

¹¹In the code, there exists the possibility to apply the diffusion term to $(\hat{\hat{F}}_s - \hat{\hat{F}}_{s, \text{eq}})$ instead of $\hat{\hat{F}}_s$.

where $\hat{v} = \sqrt{\left(\frac{1}{2}\hat{v}_{G\parallel s}^2 + \hat{\mu}_s \hat{B}_{\text{axis}}\right) / \hat{T}_{s,\text{coll}}}$ while the Chandrasekhar function G is defined

$$G(\hat{v}) = \frac{\Phi(\hat{v}) - \hat{v}\Phi'(\hat{v})}{2\hat{v}^2} \quad ; \quad \Phi(\hat{v}) = \frac{2}{\sqrt{\pi}} \int_0^{\hat{v}} e^{-x^2} dx \quad ; \quad \Phi'(\hat{v}) = \frac{2}{\sqrt{\pi}} e^{-\hat{v}^2} \quad (\text{F.20})$$

and the radial profile $\hat{\mathcal{A}}_s$ takes the form

$$\hat{\mathcal{A}}_s(\hat{r}) = \frac{1}{\sqrt{\hat{\mathcal{A}}_s}} \left(3 \frac{\sqrt{\pi}}{2} \frac{\hat{v}_{T_{s,\text{coll}}}^3 \hat{\epsilon}^{3/2}}{q(\hat{r}) \hat{R}_0} \right) \hat{v}_{*s} \quad (\text{F.21})$$

The normalized collisionality $\hat{v}_{*s}$ is given as input data, and the normalized collisionality frequency is deduced from this scalar as:

$$\hat{v}_{ss} = \frac{1}{\sqrt{\hat{\mathcal{A}}_s}} \left(\frac{\hat{v}_{T_{s,\text{coll}}} \hat{\epsilon}^{3/2}}{q(\hat{r}) \hat{R}_0} \right) \hat{v}_{*s} \quad (\text{F.22})$$

Considering that the species s is the major ion species, the collisionality $\hat{v}_{*s'}$ of each minority ion species s' is determined as

$$\hat{v}_{*s'} = \left(\frac{\hat{n}_{s'}}{\hat{n}_s} \right) \left(\frac{\hat{Z}_{s'}}{\hat{Z}_s} \right)^4 \left(\frac{\hat{T}_{s,\text{coll}}}{\hat{T}_{s',\text{coll}}} \right)^2 \hat{v}_{*s} \quad (\text{F.23})$$

In practice, $\hat{T}_{s,\text{coll}}/\hat{T}_{s',\text{coll}}$ is approximated by $\hat{T}_s/\hat{T}_{s'}$. The ratio $\hat{n}_{s'}/\hat{n}_s$ and $\hat{T}_{s,\text{coll}}/\hat{T}_{s',\text{coll}}$ are computed at the radial point $\hat{r}_p$ which usually corresponds to the middle of the simulation radial domain. The normalized mean velocity $\hat{V}_{\parallel s,\text{coll}}$ and the normalized mean temperature $\hat{T}_{s,\text{coll}}$ can be calculated as:

$$\hat{V}_{\parallel s,\text{coll}} = \hat{P}^{-1} \left(\hat{I}_4 \times \hat{I}_1 - \hat{I}_2 \times \hat{I}_3 \right) \quad ; \quad \hat{T}_{s,\text{coll}} = \hat{P}^{-1} \left(\hat{I}_0 \times \hat{I}_2 - \hat{I}_1^2 \right) \quad (\text{F.24})$$

with the normalized mean pressure defined as $\hat{P} = \hat{I}_0 \times \hat{I}_4 - \hat{I}_1 \times \hat{I}_3$ and where the five integrals $\hat{I}_0$, $\hat{I}_1$, $\hat{I}_2$, $\hat{I}_3$ and $\hat{I}_4$ are

$$\hat{I}_0 = \langle \hat{\mathcal{D}}_{\parallel s} \rangle \quad ; \quad \hat{I}_1 = \langle \hat{\mathcal{D}}_{\parallel s} \hat{v}_{G\parallel s} \rangle \quad ; \quad \hat{I}_2 = \langle \hat{\mathcal{D}}_{\parallel s} \hat{v}_{G\parallel s}^2 \rangle \quad (\text{F.25})$$

$$\hat{I}_3 = \left\langle \frac{1}{\hat{B}_{\parallel s}^*} \partial_{\hat{v}_{G\parallel s}} \left(\hat{B}_{\parallel s}^* \hat{\mathcal{D}}_{\parallel s} \right) \right\rangle \quad \text{and} \quad \hat{I}_4 = \left\langle \frac{1}{\hat{B}_{\parallel s}^*} \partial_{\hat{v}_{G\parallel s}} \left(\hat{B}_{\parallel s}^* \hat{v}_{G\parallel s} \hat{\mathcal{D}}_{\parallel s} \right) \right\rangle \quad (\text{F.26})$$

with the brackets $\langle \cdot \rangle = \int \cdot \hat{\tilde{F}}_s \hat{J}_v d\hat{v}_{G\parallel s} d\hat{\mu}_s$.

Appendix F.7. Normalized collisional transfers between two species

Normalized collisional energy transfer. Energy exchange between two species can be expressed by the following reduced collision operator:

$$\frac{d\hat{F}_s}{d\hat{t}} = - \frac{\hat{\eta}_{ss'}^{\mathcal{E}}}{\left(2\pi \langle \hat{T} \rangle_{ss'} \right)^{3/2}} \frac{\Delta \hat{T}_{ss'}}{\langle \hat{T} \rangle_{ss'}} \left(\frac{\hat{\mathcal{E}}_{ss'}}{\langle \hat{T} \rangle_{ss'}} - \frac{3}{2} \right) \exp \left(- \frac{\hat{\mathcal{E}}_{ss'}}{\langle \hat{T} \rangle_{ss'}} \right) \quad (\text{F.27})$$

$$\frac{d\hat{F}_{s'}}{d\hat{t}} = - \frac{\hat{\eta}_{ss'}^{\mathcal{E}}}{\left(2\pi \langle \hat{T} \rangle_{s's} \right)^{3/2}} \frac{\Delta \hat{T}_{s's}}{\langle \hat{T} \rangle_{s's}} \left(\frac{\hat{\mathcal{E}}_{s's}}{\langle \hat{T} \rangle_{s's}} - \frac{3}{2} \right) \exp \left(- \frac{\hat{\mathcal{E}}_{s's}}{\langle \hat{T} \rangle_{s's}} \right) \quad (\text{F.28})$$

with

$$\begin{aligned}\langle \hat{T} \rangle_{ss'} &= \frac{\hat{T}_s + \hat{T}_{s'}}{2} = \langle \hat{T} \rangle_{s's} \quad ; \quad \Delta \hat{T}_{ss'} = \hat{T}_s - \hat{T}_{s'} = -\Delta \hat{T}_{s's} \quad ; \quad \langle \hat{V} \rangle_{ss'} = \frac{\hat{V}_{\parallel s} + \hat{V}_{\parallel s'}}{2} = \langle \hat{V} \rangle_{s's} \\ \hat{\mathcal{E}}_{ss'} &= \frac{(\hat{v}_{G\parallel s} - \langle \hat{V} \rangle_{ss'})^2}{2} + \hat{\mu}_s \hat{B} \quad \text{and} \quad \hat{\mathcal{E}}_{s's} = \frac{(\hat{v}_{G\parallel s'} - \langle \hat{V} \rangle_{s's})^2}{2} + \hat{\mu}_{s'} \hat{B}\end{aligned}$$

and

$$\hat{\eta}_{ss'}^{\mathcal{E}}(\hat{r}) = \frac{8 \hat{\varepsilon}^{3/2}}{3\sqrt{\pi}} \frac{\sqrt{\hat{A}_s}}{\hat{A}_{s'}} \left(\frac{\hat{Z}_{s'}}{\hat{Z}_s} \right)^2 \frac{\hat{n}_{s'} \sqrt{\langle \hat{T}_{s,\text{coll}} \rangle_{\text{FS}}}}{q \hat{R}_0} \left(1 + \frac{\hat{A}_s}{\hat{A}_{s'}} \frac{\hat{T}_{s'}^2}{\hat{T}_s^2} \right)^{-3/2} \hat{\nu}_{*s} \quad (\text{F.29})$$

Let us notice that in practice, the flux surface average of $\hat{T}_{s,\text{coll}}$ is used (compare to the general expression given by (42)), to consider only the radial dependency of $\hat{\eta}_{ss'}^{\mathcal{E}}$. The normalized fluid moments $\hat{V}_{\parallel s}$ and $\hat{T}_s$ are computed as $3\hat{T}_s(r)/2 = \langle \int \hat{\mathcal{E}}_s \hat{\hat{F}}_s d^3\hat{\mathbf{v}} \rangle_{\text{FS}} / \hat{N}_s(r)$ and $\hat{V}_{\parallel s}(r) = \langle \int \hat{v}_{G\parallel s} \hat{\hat{F}}_s d^3\hat{\mathbf{v}} \rangle_{\text{FS}} / \hat{N}_s(r)$ where $\hat{\mathcal{E}}_s$ is defined as $\hat{\mathcal{E}}_s = \frac{1}{2} (\hat{v}_{G\parallel s} - \hat{V}_{\parallel s})^2 + \hat{\mu}_s \hat{B}$ while $\hat{N}_s(r) = \langle \int \hat{\hat{F}}_s d^3\hat{\mathbf{v}} \rangle_{\text{FS}}$.

Normalized collisional momentum transfer. Momentum exchange between two species can be expressed by the following reduced collision operator

$$\frac{d\hat{\hat{F}}_s}{d\hat{t}} = -\hat{\eta}^{v_{G\parallel}} \hat{A}_s^{-1/2} \Delta \hat{V}_{ss'} \hat{v}_{G\parallel s} \exp\left(-\frac{\hat{E}_s}{\langle \hat{T} \rangle_{ss'}}\right) \quad (\text{F.30})$$

$$\frac{d\hat{\hat{F}}_{s'}}{d\hat{t}} = -\hat{\eta}^{v_{G\parallel}} \hat{A}_{s'}^{-1/2} \Delta \hat{V}_{s's} \hat{v}_{G\parallel s'} \exp\left(-\frac{\hat{E}_{s'}}{\langle \hat{T} \rangle_{ss'}}\right) \quad (\text{F.31})$$

with

$$\Delta \hat{V}_{ss'} = \hat{V}_{\parallel s} \hat{A}_s^{-1/2} - \hat{V}_{\parallel s'} \hat{A}_{s'}^{-1/2} \quad (\text{F.32})$$

The averages $\langle \hat{T} \rangle_{ss'}$ and $\langle \hat{V} \rangle_{ss'}$ are respectively defined as previously. The energy $\hat{E}_s$ is equal to $\frac{1}{2} \hat{v}_{G\parallel s}^2 + \hat{\mu}_s \hat{B}$. The radial profile $\hat{\eta}^{v_{G\parallel}}$ reads:

$$\hat{\eta}^{v_{G\parallel}}(r) = \left(\frac{\varepsilon}{2\pi} \right)^{3/2} \frac{1}{q \hat{R}_0} \frac{\hat{n}_{s'}}{\hat{v}_{T_{>}}} \frac{\sqrt{\hat{T}_s} \sqrt{\langle \hat{T}_{s,\text{coll}} \rangle_{\text{FS}}}}{\langle \hat{T} \rangle_{ss'}^{5/2}} \left(\frac{\hat{Z}_{s'}}{\hat{Z}_s} \right)^2 \hat{\nu}_{*s} \quad (\text{F.33})$$

where the velocity $\hat{v}_{T_{>}}$ is calculated as $\hat{v}_{T_{>}} = \max(\hat{A}_s^{-1/2} \hat{v}_{T_s}, \hat{A}_{s'}^{-1/2} \hat{v}_{T_{s'}})$ where $\hat{v}_{T_s} = \sqrt{\hat{T}_s}$ and $\hat{v}_{T_{s'}} = \sqrt{\hat{T}_{s'}}$ and the collision frequencies $\hat{\nu}_{*s}$, $\hat{\nu}_{*s'}$ are linked by the relation (F.23).

Appendix F.8. Normalized source terms

Focusing on the source term, the gyrokinetic equation is:

$$\frac{d\hat{\hat{F}}_s}{d\hat{t}} = \hat{S}_{\text{heat}}(\hat{r}, \theta, \hat{v}_{G\parallel s}, \hat{\mu}_s) + \hat{S}_{\text{moment}}(\hat{r}, \theta, \hat{v}_{G\parallel s}, \hat{\mu}_s) + \hat{S}_{\text{vorticity}}(\hat{r}, \theta, \hat{v}_{G\parallel s}, \hat{\mu}_s) \quad (\text{F.34})$$

where the heat source can be defined as

$$\hat{S}_{\text{heat}} = \frac{\hat{S}_0^{\mathcal{E}}}{3\sqrt{2} \left(\pi \hat{T}_{s,\text{srce}} \right)^{3/2} \hat{T}_{s,\text{srce}}} \left(\frac{\hat{E}_s}{\hat{T}_{s,\text{srce}}} - \frac{3}{2} \right) \exp\left(-\frac{\hat{E}_s}{\hat{T}_{s,\text{srce}}}\right) \hat{S}_r^{\mathcal{E}} \quad (\text{F.35})$$

with $\hat{E}_s = \frac{1}{2}\hat{v}_{G\parallel s} + \hat{\mu}_s\hat{B}$ or as a pure source of heating

$$\hat{S}_{\text{heat}} = \left\{ \left[\hat{v}_{G\parallel s}^2 - \frac{1}{2} - \frac{\hat{J}_{\parallel B}}{2 - \hat{J}_{\parallel B}^2} (2 - \hat{\mu}_s) (2\hat{v}_{G\parallel s} - \hat{J}_{\parallel B}) \right] \frac{2\hat{S}_0^\mathcal{E}}{(2\pi\hat{T}_{s,\text{srce}})^{3/2} \hat{T}_{s,\text{srce}}} \hat{S}_r^\mathcal{E} \right\} e^{-\hat{v}_{G\parallel s}^2 - \hat{\mu}_s} \quad (\text{F.36})$$

with $\hat{\mu}_s = \frac{\hat{\mu}_s\hat{B}}{\hat{T}_{s,\text{srce}}}$, $\hat{v}_{G\parallel s} = \frac{\hat{v}_{G\parallel s}}{\sqrt{2\hat{T}_{s,\text{srce}}}}$ and $\hat{J}_{\parallel B} \equiv \frac{\sqrt{\hat{A}_s}}{\hat{Z}_s} \frac{\sqrt{2\hat{T}_{s,\text{srce}}}}{\hat{B}^2} \mu_0 \hat{J}_{\parallel}$. The source $\hat{S}_{\text{moment}}$ is a pure source of momentum expressed as

$$\hat{S}_{\text{moment}} = \left\{ \left[2\hat{v}_{G\parallel s}(2 - \hat{\mu}_s) - \hat{J}_{\parallel B} (1 + 2\hat{v}_{G\parallel s} - \hat{\mu}_s) \right] \frac{\hat{S}_0^{v_{G\parallel}}}{4\pi^{3/2}\hat{T}_{s,\text{srce}}^2} \hat{S}_r^{v_{G\parallel}} \right\} e^{-\hat{v}_{G\parallel s}^2 - \hat{\mu}_s} \quad (\text{F.37})$$

and the $\hat{S}_{\text{vorticity}}$ is a pure source of vorticity is defined as

$$\hat{S}_{\text{vorticity}} = - \left\{ \left[2\hat{v}_{G\parallel s}^2 - \hat{\mu}_s \right] \frac{\hat{S}_0^\Omega}{\alpha (2\pi\hat{T}_{s,\text{srce}})^{3/2}} \hat{S}_r^\Omega \right\} e^{-\hat{v}_{G\parallel s}^2 - \hat{\mu}_s} \quad \text{with } \alpha = \frac{\hat{A}_s}{\hat{Z}_s^2} \frac{\hat{T}_{s,\text{srce}}}{2\hat{B}^2} \quad (\text{F.38})$$

The radial components of the sources (*i.e.* $\hat{S}_r^\mathcal{E}$, $\hat{S}_r^{v_{G\parallel}}$ and $\hat{S}_r^\Omega$) are defined as:

$$\hat{S}_r^x(\hat{r}) = -\frac{1}{2} \left[\tanh \left(\frac{\hat{\rho} - (\hat{\rho}_S^x + 3\hat{L}_S^x)}{\hat{L}_S^x} \right) + \tanh \left(-\frac{\hat{\rho} - (\hat{\rho}_S^x - 3\hat{L}_S^x)}{\hat{L}_S^x} \right) \right] \quad (\text{F.39})$$

where $\hat{\rho} = (\hat{r} - \hat{r}_{\min})/\hat{L}_r$. $\hat{\rho}_S^x$ and $\hat{L}_S^x$ are input data corresponding to the radial position and the normalized width of the different sources. These radial parts are normalized such that

$$\int_{\hat{r}_{\min}}^{\hat{r}_{\max}} \hat{r} \, d\hat{r} \hat{S}_r^x(\hat{r}) = 1$$

Appendix F.9. Normalized source of impurity

Let considers the species s as the majority species and the species s' as an impurity species. Then, it is possible to add a source of impurities, by adding to the right hand side of equation (F.34) a source of matter. The source $\hat{S}_{n_{s'}}(\hat{r}, \theta, \hat{v}_{G\parallel s}, \hat{\mu}_s)$ of impurity s' is defined as

$$\hat{S}_{n_{s'}} = \frac{\hat{S}_0^{n_{s'}} \hat{S}_r}{(2\pi\hat{T}_{s',\text{srce}})^{3/2}} \left(\frac{5}{2} - \hat{\mu}_{s'} - \hat{v}_{G\parallel s'}^2 \right) \exp \left(-\hat{v}_{G\parallel s'}^2 - \hat{\mu}_{s'} \right) \quad (\text{F.40})$$

where $\hat{v}_{G\parallel s'} = \hat{v}_{G\parallel s}/\sqrt{2\hat{T}_{s',\text{srce}}}$ and $\hat{\mu}_{s'} = \hat{\mu}_s\hat{B}/\hat{T}_{s',\text{srce}}$. $\hat{S}_0^{n_{s'}}$ corresponds to the normalized intensity of the source. These injection of impurities must be compensated by the injection of source of matter for the majority species such that

$$Z_s \int S_{n_s} J_{v_s} \, dv_{G\parallel} \, d\mu_s + Z_{s'} \int S_{n_{s'}} J_{v_{s'}} \, dv_{G\parallel} \, d\mu_{s'} = 0$$

with J_{V_s} (resp. $J_{V_{s'}}$) the jacobian in velocity space for species s (resp. species s'). The normalized radial profiles of the sources are assumed equal to $\hat{S}_r$. Therefore,

$$\hat{S}_{n_s} = \frac{\hat{S}_0^{n_s} \hat{S}_r}{\left(2\pi \hat{T}_{s,\text{srce}}\right)^{3/2}} \left(\frac{5}{2} - \hat{\mu}_s - \hat{v}_{G\parallel s}^2\right) e^{-\hat{v}_{G\parallel s}^2 - \hat{\mu}_s} \text{ with } \hat{S}_0^{n_s} = -\frac{\hat{Z}_{s'}}{\hat{Z}_s} \left(\frac{\hat{T}_{s',\text{srce}}}{\hat{T}_{s,\text{srce}}}\right)^{3/2} \hat{S}_0^{n_{s'}} \quad (\text{F.41})$$

In the code, the temperature profile of the density sources are imposed equal, *i.e.* $\hat{T}_{s,\text{srce}} = \hat{T}_{s',\text{srce}}$.

Appendix G. Invariance

Let us consider a reference simulation without source terms. Let us also define a second simulation similar to the reference one but where the mass and charge have been respectively multiplied by a factor α_A and α_Z . Then, it is shown in this appendix that 7 control parameters (α_x , α_t , α_{T_e} , α_ϕ , α_{coll} , α_{diff} and α_{Krook} that respectively rescale the space, the time, the electron temperature, the electrostatic potential, the collision operator, the diffusion term and the Krook operator) are sufficient to ensure that both simulations –reference and scaled– are equivalent. The constraints on the different control parameters are summarized in table G.13 and the proof follows.

	Charge	Mass	Length	Time	T_e	Elec. potential
Scaling	α_Z	α_A	$\sqrt{\alpha_A}/\alpha_Z$	α_A/α_Z	α_Z^{-1}	α_Z^{-1}
	Collision	Krook	Diffusion			
Scaling	1	α_Z/α_A	α_Z^{-1}			

Table G.13: Scaling which must be applied to the different characteristic quantities to obtain an equivalence between any reference simulation and a scaled simulation

For the proof let us first consider the following normalized reference Vlasov equation (deduced from eqs. (F.3) to (F.8)) for a species of charge $\hat{Z}_s$ and mass $\hat{A}_s$

$$\begin{aligned} & \frac{\partial \hat{F}_s}{\partial t} + \left\{ \frac{1}{\sqrt{\hat{A}_s}} \hat{v}_{G\parallel} \hat{\mathbf{b}}_s^* + \left(\frac{\hat{v}_{G\parallel}^2 + \hat{\mu}_s \hat{B}}{\hat{Z}_s \hat{B}_{\parallel s}^* \hat{B}} \right) (\mathbf{b} \times \hat{\nabla} \hat{B}) + \frac{1}{\hat{B}_{\parallel s}^*} (\mathbf{b} \times \hat{\nabla} \hat{\phi}) \right\} \cdot \hat{\nabla} \hat{F}_s \\ & + \left(-\frac{1}{\sqrt{\hat{A}_s}} \hat{\mu}_s \hat{\mathbf{b}}_s^* \cdot \hat{\nabla} \hat{B} - \frac{\hat{Z}_s}{\sqrt{\hat{A}_s}} \hat{\mathbf{b}}_s^* \cdot \hat{\nabla} \hat{\phi} + \hat{v}_{G\parallel} \hat{\mathbf{v}}_{E \times B_s} \cdot \hat{\nabla} \ln \hat{B} \right) \frac{\partial \hat{F}_s}{\partial \hat{v}_{G\parallel}} \\ & = \frac{1}{\hat{B}_{\parallel s}^*} \frac{\partial}{\partial \hat{v}_{G\parallel}} \left[\hat{B}_{\parallel s}^* \hat{\mathcal{D}}_{\parallel s} \left(\frac{\partial \hat{F}_s}{\partial \hat{v}_{G\parallel}} + \frac{(\hat{v}_{G\parallel} - \hat{V}_{\parallel s, \text{coll}})}{\hat{v}_{T_{s, \text{coll}}}^2} \hat{F}_s \right) \right] + \frac{1}{\hat{B}_{\parallel s}^*} \left[\frac{1}{\hat{r}} \frac{\partial}{\partial \hat{r}} \left(\hat{r} \hat{\chi} \hat{B}_{\parallel s}^* \frac{\partial \hat{F}_s}{\partial \hat{r}} \right) \right] \\ & - \hat{\nu}(\hat{r}) (\hat{F}_s - \hat{F}_{s, \text{eq}}) \end{aligned}$$

where

$$\hat{\mathbf{b}}_s^* = \frac{1}{\hat{B}_{\parallel s}^*} \left(\hat{\mathbf{B}} + \sqrt{\hat{A}_s} \frac{\hat{v}_{G\parallel}}{\hat{Z}_s \hat{B}} \hat{\nabla} \times \hat{\mathbf{B}} \right)$$

and

$$\hat{\mathcal{D}}_{\parallel s}(\hat{r}, \hat{v}) = \hat{\mathcal{A}}_s(\hat{r}) \left(\frac{\Phi(\hat{v}) - G(\hat{v})}{2\hat{v}} \right) \quad \text{with} \quad \hat{\mathcal{A}}_s(\hat{r}) = \frac{1}{\sqrt{\hat{A}_s}} \left(3 \frac{\sqrt{\pi}}{2} \frac{\hat{v}_{T_{s, \text{coll}}}^3}{q(\hat{r}) \hat{R}_0} \hat{\epsilon}^{3/2} \right) \hat{\nu}_{*s}$$

The Chandrasekhar function G and the function Φ are defined by equation (F.20). Let us apply a scaling

with the control parameters defined below, this leads to the following new Vlasov equation

$$\begin{aligned}
& \frac{1}{\alpha_t} \frac{\partial \hat{F}_s}{\partial \hat{t}} + \left\{ \frac{1}{\sqrt{\alpha_A \hat{A}_s}} \hat{v}_{G\parallel} \hat{\mathbf{b}}_s^* + \left(\frac{\hat{v}_{G\parallel}^2 + \hat{\mu}_s \hat{B}}{\alpha_Z \hat{Z}_s \hat{B}_{\parallel s}^* \hat{B}} \right) \left(\mathbf{b} \times \frac{1}{\alpha_x} \hat{\nabla} \hat{B} \right) + \frac{1}{\hat{B}_{\parallel s}^*} \left(\mathbf{b} \times \frac{1}{\alpha_x} \hat{\nabla} (\alpha_\phi \hat{\phi}) \right) \right\} \cdot \frac{\hat{\nabla} \hat{F}_s}{\alpha_x} \\
& + \left(-\frac{1}{\sqrt{\alpha_A \hat{A}_s}} \hat{\mu}_s \hat{\mathbf{b}}_s^* \cdot \frac{\hat{\nabla} \hat{B}}{\alpha_x} - \frac{\alpha_Z \hat{Z}_s}{\sqrt{\alpha_A \hat{A}_s}} \hat{\mathbf{b}}_s^* \cdot \frac{\hat{\nabla} (\alpha_\phi \hat{\phi})}{\alpha_x} + \hat{v}_{G\parallel} \hat{\mathbf{v}}_{E \times B_s} \cdot \frac{1}{\alpha_x} \hat{\nabla} \ln \hat{B} \right) \frac{\partial \hat{F}_s}{\partial \hat{v}_{G\parallel}} \\
& = \frac{1}{\hat{B}_{\parallel s}^*} \frac{\partial}{\partial \hat{v}_{G\parallel}} \left[\hat{B}_{\parallel s}^* \frac{\alpha_{\text{coll}}}{\sqrt{\alpha_A} \alpha_x} \hat{\mathcal{D}}_{\parallel s} \left(\frac{\partial \hat{F}_s}{\partial \hat{v}_{G\parallel}} + \frac{(\hat{v}_{G\parallel} - \hat{V}_{\parallel s, \text{coll}})}{\hat{v}_{T_s, \text{coll}}^2} \hat{F}_s \right) \right] \\
& + \frac{1}{\hat{B}_{\parallel s}^*} \left[\frac{1}{\alpha_x \hat{r}} \frac{\partial}{\partial \hat{r}} \left(\hat{r} \alpha_{\text{diff}} \hat{\chi} \hat{B}_{\parallel s}^* \frac{1}{\alpha_x} \frac{\partial \hat{F}_s}{\partial \hat{r}} \right) \right] - \alpha_{\text{Krook}} \hat{\nu}(\hat{r}) (\hat{F}_s - \hat{F}_{s, \text{eq}})
\end{aligned}$$

with

$$\hat{\mathbf{b}}_s^* = \frac{1}{\hat{B}_{\parallel s}^*} \left(\hat{\mathbf{B}} + \sqrt{\alpha_A \hat{A}_s} \frac{\hat{v}_{G\parallel}}{\alpha_Z \hat{Z}_s \hat{B}} \frac{\hat{\nabla} \times \hat{\mathbf{B}}}{\alpha_x} \right)$$

Concerning the left hand side, this equation is equivalent to the reference one, if and only if:

$$\frac{1}{\alpha_t} = \frac{1}{\sqrt{\alpha_A} \alpha_x} = \frac{\sqrt{\alpha_A}}{\alpha_Z \alpha_x^2} = \frac{1}{\alpha_Z \alpha_x^2} = \frac{\alpha_\phi}{\alpha_x^2} \quad (\text{G.1})$$

$$\frac{1}{\sqrt{\alpha_A} \alpha_x} = \frac{1}{\alpha_Z \alpha_x^2} = \frac{\alpha_Z \alpha_\phi}{\sqrt{\alpha_A} \alpha_x} = \frac{\alpha_\phi}{\alpha_x^2} \quad (\text{G.2})$$

Therefore, according to the last equality of equation (G.2), $\alpha_x = \sqrt{\alpha_A}/\alpha_Z$. Besides, due to the last equality of equation (G.1), $\alpha_\phi = \alpha_Z^{-1}$. And finally, due to the first equality of equation (G.1), $\alpha_t = \sqrt{\alpha_A}/\alpha_x$, so $\alpha_t = \alpha_A/\alpha_Z$. Taking into account the collision operator $1/\alpha_t = \alpha_{\text{coll}}/(\sqrt{\alpha_A} \alpha_x)$ which gives $\alpha_{\text{coll}} = 1$. In addition, the equivalence condition for the Krook operator leads to $1/\alpha_t = \alpha_{\text{Krook}}$ thus $\alpha_{\text{Krook}} = \alpha_Z/\alpha_A$ while for the diffusion term the condition reads $1/\alpha_t = \alpha_{\text{diff}}/\alpha_x^2$, *i.e.* $\alpha_{\text{diff}} = \alpha_Z^{-1}$.

Concerning the equivalence for the quasi-neutrality let us consider normalized equation

$$-\sum_s \hat{A}_s \hat{\nabla} \left(\frac{\hat{n}_{s, \text{eq}}}{\hat{B}} \hat{\nabla}_\perp \hat{\phi} \right) + \frac{1}{Z_0 \hat{T}_e} \sum_s \hat{Z}_s \hat{n}_{s, \text{eq}} [\hat{\phi} - \langle \hat{\phi} \rangle_{\text{FS}}] = \sum_s \hat{Z}_s \int \hat{J}_0 (\hat{F}_s - \hat{F}_{s, \text{eq}}) d^3 \hat{\mathbf{v}}$$

Therefore, the rescaled equation reads:

$$-\frac{\alpha_A \alpha_\phi}{\alpha_x^2} \sum_s \hat{A}_s \hat{\nabla} \left(\frac{\hat{n}_{s, \text{eq}}}{\hat{B}} \hat{\nabla}_\perp \hat{\phi} \right) + \frac{\alpha_Z \alpha_\phi}{\alpha_{T_e}} \frac{1}{Z_0 \hat{T}_e} \sum_s \hat{Z}_s \hat{n}_{s, \text{eq}} [\hat{\phi} - \langle \hat{\phi} \rangle_{\text{FS}}] = \sum_s \alpha_Z \hat{Z}_s \int \hat{J}_0 (\hat{F}_s - \hat{F}_{s, \text{eq}}) d^3 \hat{\mathbf{v}}$$

The equivalence between reference and rescaled QN equations is ensured if and only if:

$$\frac{\alpha_A \alpha_\phi}{\alpha_x^2} = \frac{\alpha_Z \alpha_\phi}{\alpha_{T_e}} \Rightarrow \alpha_{T_e} = \frac{\alpha_x^2 \alpha_Z}{\alpha_A} \quad \text{and} \quad \alpha_Z = \frac{\alpha_Z \alpha_\phi}{\alpha_{T_e}} \Rightarrow \alpha_\phi = \alpha_{T_e}$$

Therefore $\alpha_{T_e} = 1/\alpha_Z$ which confirms the previous equality $\alpha_\phi = \alpha_Z^{-1}$.

Finally, regarding the gyroaverage operator in the case of a Padé approximation, *i.e.* $\hat{J}_0 \sim 1 + \frac{1}{2} \frac{\hat{A}_s}{\hat{Z}_s^2} \frac{\hat{\mu}_s}{\hat{B}} \hat{\nabla}_\perp^2$, its invariance is conserved if $\alpha_A = \alpha_Z^2 \alpha_x^2$.

AppendixH. Input data files

This appendix contains the input data files of all simulations used in this paper.

15/01/2015 A150_FD_allwcoil2_semitore 1	15/01/2015 A150_FD_allwcoil2_semitore 2
<pre> ***** # job parameters ***** NB RESTART = 5 TIME = 10:00:00 JOBNAME = A75_A1Z1 ***** # Variables used to specify if the # simulation is a restart from # another one ***** SEARCH SP0 = 0 ! 0 if no restart, 1 otherwise DIR SP0 = ! if SEARCH SP0=1, name of the directory ! containing the major species restart files ***** # Variables for parallelization ***** NSPECIES = 1 NPROC_R = 2 NPROC_TH = 4 NPROC_MU = 16 NTHREAD = 4 ***** # Variables for radial profile input files ***** ns0_sp0_filename = ns0_sp0_test.dat ns0_spl_filename = ns0_spl_test.dat Ts0_sp0_filename = Ts0_sp0_test.dat Ts0_spl_filename = Ts0_spl_test.dat q_filename = safety_factor_test.dat ***** # gysela input datas ***** &PARALLEL Nspecies = \$NSPECIES Nbproc_r = \$NPROC_R Nbproc_theta = \$NPROC_TH Nbproc_mu = \$NPROC_MU bloc_phi = \$NTHREAD Nbthread = \$NTHREAD large_platform = .true. distrib3D = .true., / &MESH CylindricalGeometry = .false. Nr = 255 Ntheta = 256 Nphi = 128 Nvpar = 127 Nmu = \$(NPROC_MU-1) a = 150. rhomin = 0.15 rhomax = 0.85 </pre>	<pre> Ltheta = 6.283185307179586476925286766559005768394 Lphi = 3.141592653589793238462643383279502884197 Z0 = 1. nb_vth0 = 7. Lmu = 12. mumin = 0., / &EQUIL canonical_profile = 0 tau0 = 1. A0 = 1. Z0 = 1. As = 1. Zs = 1. species_frac = 1. read_ns0 = .false. read_Ts0 = .false. read_q = .false. read_vpar0 = .false. profile_choice = 2 rpeak = 0.5 kappa_ns0 = 2.2 kappa_Ts0 = 8. deltar_ns0 = 0.04 deltar_Ts0 = 0.04 q_profile = 2 q_param1 = 1.5 q_param2 = 1.7 q_param3 = 2.8 q_param4 = 1. magnetic_drift = .true. perturb_amplitude = 0.001 perturb_choice = 2 m = 28 n = 10 FFTPotential_filter = 0 FFTPotential_filter_deltam = 0. !--> Gyroaverage tunable input parameters gyroaverage = .true. ! gyro strategy ! All possible choices for gyro strategy (by default "PADE") ! ("PADE", "LIBGYROAVG_HERMITE", "LIBGYROAVG_HERMITE_C1", ! "LIBGYROAVG_HERMITE_C1_PRE_COMP") ! gyro Nbpnts = 32, ! usefull only if gyro_strategy=LIBGYROAVG_* TF_ripple = .false. delta_ripple = 0 N_ripple = 0, / &COLLISION_OPERATOR collisions = .true. version_coll = 2 energy_transfer = .false. moment_transfer = .false. nustar = 0.28 coll_rhrefresh = 10 coll_nbstep = 1, </pre>
15/01/2015 A150_FD_allwcoil2_semitore 3	15/01/2015 A150_FD_allwcoil2_semitore 4
<pre> / &BUFFER REGION coef_krook = 0.01 coef_diffr = 0.015 coef_diffrh = 0.0 coef_nustar = 20. buffer_asymmetry = .false. right_buffer_only = .true. buffer_location = 0.06 buffer_stiffness = 0.017635, / &ALGORITHM Vlasov_scheme = "BSL" limiter = 0 deltat = 5. nbiter = 250 dt_diag = 50., / &SOURCE gradient_driven = .false. grad_version = 1 flux_driven = .false. heat_source = .false. heat_version = 0 Sce0 = 0.0217 TS = 1.5 rhoS = 0. LS = 0.12 dens_source = .false. Sce_dens = 0.01 TS_dens = 1 rhoS_dens = 0.1 LS_dens = 0.01 mom_source = .false. Sce_mom = -0.001 TS_mom = 1.5 rhoS_mom = 0. LS_mom = 0.1 vor_source = .false. Sce_vor = 0.03 TS_vor = 1 rhoS_vor = 0.1 LS_vor = 0.01, / &TEST !--> Tests for general numerics integral_vperp = .true. integration_scheme = 2 !--> Tests for QN solver solve_QN = .true. QN_version = 5 QN_comm_version = 1 QN_coef_polarization = 1. QN_coef_Phi00 = 1 Phi00_BCrmin_Neumann = .false. </pre>	<pre> QN_species_impact = 1 !--> Tests for Vlasov solver adaptive_deltat = .false. advec2D_deltat = 0. reduced_deltat = 0 reduced_begin = 800000. reduced_end = 400000. ! hfilter_strategy = " ! ("DEFAULT", "NONE", "FFT", "WAVELET") ! hfilterfreq = 0 ! hfilterfref = 0 deltaF_in_diff = .true. avoidtaylor = .false. deltat_in_advec = .false. split_linear = .false. clampv = .false. cancel_nonlin = .false. diagf = .false. !--> Tests for RHS of Boltzmann equation RHS_only = .false. !--> Tests for equilibrium definition feq_choice = 1 canonical_vpar0 = 4 hvar in fperturb = .true. !--> Tests with simplified physics BstareqB = .false. single_m = .false. single_n = .false. Rosenbluth_Hinton = .false. !--> Tests for computer sciences ask_for_checksum = 1 modulo_restart_iter = 150 ! compact_rstfile = .false. ! checkoverflow = .false., / &OUTPUT integration_CS = .true. diag_level = 7 Phi3D_saving = .false. Moments3D_saving = .false. fluxes3D_GC_saving = .false. f5D_saving = .false. Phi3D_dt_diag = 500. Moments3D_dt_diag = 500. fluxes3D_GC_dt_diag = 500. f5D_dt_diag = 500. spectra_saving = .false. spectra_dt_diag = 10. r_diags = 31 doubleSP_m = 2 doubleSP_n = 0 iter_beg_write_trace = 1 iter_end_write_trace = 3, / </pre>

Figure H.14: Input data file used in the code for the first simulation for invariance tests (see section 4.2).

<pre> 26/01/2015 test_RH_q1.9_Nr64Ntheta64Nvpar128_dt5 1 ***** # job parameters # simulation is a restart from # another one ***** NB RESTART = 0 TIME = 08:00:00 MACHINE_FILE = JOBNAME = gysela5D # Variables used to specify if the # simulation is a restart from # another one ***** SEARCH_SP0 = 0 ! 0 if no restart, 1 otherwise DIR_SP0 = ! if SEARCH_SP0=1, name of the directory ! containing the major species restart files ***** # Variables for parallelization ***** NSPECIES = 1 NPROC_R = 2 NPROC_TH = 1 NPROC_MU = 16 NTHREAD = 4 # Variables for radial profile input files ***** ns0_sp0_filename = ns0_sp0_test.dat ns0_sp1_filename = ns0_sp1_test.dat Ts0_sp0_filename = Ts0_sp0_test.dat Ts0_sp1_filename = Ts0_sp1_test.dat q_filename = safety_factor_test.dat Vpar0_filename = Vpar0_test.dat ripple_filename = ripple_test.dat ***** # gysela input datas ***** &PARALLEL Nspecies = \$NSPECIES Nbproc_r = \$NPROC_R Nbproc_theta = \$NPROC_TH Nbproc_mu = \$NPROC_MU bloc_phi = \$NTHREAD Nbthread = \$NTHREAD large_platform = .true. distrib3D = .false., / &MESH CylindricalGeometry = .false. Nr = 63 Ntheta = 64 Nphi = 32 Nvpar = 127 Nmu = \$((NPROC_MU-1)) / 26/01/2015 test_RH_q1.9_Nr64Ntheta64Nvpar128_dt5 3 moment transfer = .false. nustar = 0.1 coll_FHrefresh = 2 coll_nbstep = 1, / &BUFFER_REGION coef_krook = 0.0 coef_diffrr = 0.0 coef_diffth = 0.0 coef_nustar = 0 buffer_asymmetry = .false. buffer_location = 0.1 buffer_stiffness = 0.1, / &ALGORITHM Vlasov_scheme = "BSL" limiter = 0 deltat = 5 nbiter = 4000 dt_diag = 50., / &SOURCE gradient_driven = .false. grad_version = 1 flux_driven = .false. heat_source = .false. heat_version = 0 Sce0 = 0.01 TS = 1 rhoS = 0.1 LS = 0.01 dens_source = .false. Sce_dens = 0.01 TS_dens = 1 rhoS_dens = 0.1 LS_dens = 0.01 mom_source = .false. Sce_mom = 0.02 TS_mom = 1 rhoS_mom = 0.1 LS_mom = 0.02 vor_source = .false. Sce_vor = 0.03 TS_vor = 1 rhoS_vor = 0.1 LS_vor = 0.03 fp_source = .false. fast_part = .false. Sce_fp = 0.01 Sce_fpplus = 0.6 TS_fp = 0.5 v0_fp = 2 rhoS_fp = 0.5 LS_fp = 0.015 alpha_trapp = 1., / </pre>	<pre> 26/01/2015 test_RH_q1.9_Nr64Ntheta64Nvpar128_dt5 2 a = 100. rhomin = 0.2 rhomax = 0.8 Ltheta = 6.283185307179586476925286766559005768394 Lphi = 6.283185307179586476925286766559005768394 aspect_ratio = 2.78 nb_vth0 = 7 Lmu = 12. Lmunin = 0., / &EQUIL canonical_profile = 1 tau0 = 1. A0 = 1. Z0 = 1. As = 1. Zs = 1. species_frac = 1. read_ns0 = .false. read_Ts0 = .false. read_q = .false. read_Vpar0 = .false. profile_choice = 1 rpeak = 0.5 kappa_ns0 = 0.0000001 kappa_Ts0 = 0.0000001 deltar_ns0 = 0.2 deltar_Ts0 = 0.1 q_profile = 2 q_param1 = 1.9 q_param2 = 0.2 q_param3 = 1.9 q_param4 = 0 magnetic_drift = .true. perturb_amplitude = 0.001 perturb_choice = 1 m = 0 n = 0 FFTpotential_filter = 0 FFTpotential_filter_deltam = 0. ! --> Gyroaverage tunable input parameters gyroaverage = .true. ! gyro_strategy = "" ! All possible choices for gyro_strategy (by default "PADE") ! ("PADE", "LIBGYROAVG_HERMITE", "LIBGYROAVG_HERMITE_C1", ! "LIBGYROAVG_HERMITE_C1_PRE_COMP") ! gyro_Nbpoints = 32, ! usefull only if gyro_strategy=LIBGYROAVG_* TF_ripple = .false. read_ripple = .false. delta_ripple = 0 N_ripple = 0, / &COLLISION_OPERATOR collisions = .false. version_coll = 2 energy_transfer = .false. / 26/01/2015 test_RH_q1.9_Nr64Ntheta64Nvpar128_dt5 4 &FAST_PARTICLES bump_on_tail = .false. fp_rshape = .false. fp_lambda = .true. fp_mom = .false. fp_energy = .true. fp_momgauss = .false. Eb_fp = 50. Emin_fp = 0. n_fp = 0 dn_fp = 0 E_critfp = 5. rho_0fp = 0.5 deltarho_fp = 0.01 lambda_0fp = 0.5 deltalambdafp = 0.2 psi_0fp = 0.5 deltapstfp = 0.01 zetabarf = 4. / &TEST ! --> Tests for general numerics integral_vperp = .true. integration_scheme = 2 ! --> Tests for QN solver solve_QN = .true. QN_version = 3 QN_comm_version = 1 QN_coef_polarization = 1. QN_coef_Phi00 = 1 Phi00_BCmin_Neumann = .false. QN_species_impact = 1. ! --> Tests for Vlasov solver adaptive_deltat = .false. advec2D_deltat = 0. reduced_deltat = 0 reduced_begin = 800000. reduced_end = 120000. hfilterfreq = 0 ! hfilter_strategy = "" ! ("DEFAULT", "NONE", "FFT", "WAVELET") hfilterfref = 0 deltaF_in_diff = .false. avoidtaylor = .false. deltaF_in_advec = .false. split_linear = .false. clampv = .false. cancel_nonlin = .false. diagf = .false. ! --> Tests for RHS of Boltzmann equation RHS_only = .false. ! --> Tests for equilibrium definition feq_choice = 1 canonical_vpar0 = 3 hvpar_in_fperturb = .true. ! --> Tests with simplified physics BstareqB = .false. single_m = .false. / </pre>
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Figure H.15: Example of input data file used for Rosenbluth-Hinton test (see section 4.3.1).

<pre> 27/01/2015 A180_cyclone5D_NEWPROF_SINGLEN_m7n5 1 ***** # job parameters # simulation is a restart from # another one ***** NB RESTART = 8 TIME = 08:00:00 MACHINE_FILE = JOBNAME = A180_m7n5 # Variables used to specify if the # simulation is a restart from # another one ***** SEARCH_SP0 = 0 ! 0 if no restart, 1 otherwise DIR_SP0 = ! if SEARCH_SP0=1, name of the directory ! containing the major species restart files ***** # Variables for parallelization ***** NSPECIES = 1 NPROC_R = 2 !150:4 NPROC_TH = 4 !150:4 NPROC_MU = 16 !150:32 NTHREAD = 4 !150:4 # Variables for radial profile input files ***** ns0_sp0_filename = ns0_sp0_test.dat ns0_sp1_filename = ns0_sp1_test.dat Ts0_sp0_filename = Ts0_sp0_test.dat Ts0_sp1_filename = Ts0_sp1_test.dat Te0_filename = Te0_test.dat q_filename = safety_factor_test.dat # gysela input datas ***** &PARALLEL Nspecies = \$NSPECIES Nbproc_r = \$NPROC_R Nbproc_theta = \$NPROC_TH Nbproc_mu = \$NPROC_MU bloc_phi = \$NTHREAD Nbthread = \$NTHREAD transpose4D = .true. large_platform = .true. distrib3D = .false., / &MESH CylindricalGeometry = .false. Nr = 255 Ntheta = 256 Nphi = 128 Nvpar = 64 Nmu = \$((NPROC_MU-1)) / 27/01/2015 A180_cyclone5D_NEWPROF_SINGLEN_m7n5 3 coef_krook = 0. coef_diffrr = 0.0 !150: 0.04 coef_diffth = 0.0 !150: 0.04 coef_diffh = 0.0 !75: 100 coef_nustar = 0.05 !.true. buffer_asymmetry = .false. !.true. buffer_location = 0.05 !75: 0.1 !150: 0.08 buffer_stiffness = 0.025, !75: 0.02, !150: 0.018 / &ALGORITHM Vlasov_scheme = "BSL" limiter = 0 deltat = 40. nbiter = 200 dt_diag = 240., / &SOURCE gradient_driven = .false. flux_driven = .false. heat_source = .false. heat_version = 0 Sce0 = 0.009 TS = 1. rhoS = 0.1 LS = 0.01 dens_source = .false. Sce_dens = 0.009 TS_dens = 1. rhoS_dens = 0.1 LS_dens = 0.01 mom_source = .false. Sce_mom = 0.018 TS_mom = 1. rhoS_mom = 0.1 LS_mom = 0.02 vor_source = .false. Sce_vor = 0.009 TS_vor = 1. rhoS_vor = 0.1 LS_vor = 0.01, / &TEST gyroaverage = .true. integral_vperp = .true. BstareqB = .false. Rosenbluth_Hinton = .false. solve_QN = .true. QN_version = 4 QN_comm_version = 1 QN_coef_polarization = 1. QN_coef_Phi00 = 0 Phi00_BCrmin_Neumann = .false. single_m = .false. single_n = .true. RHS_only = .false. feq_choice = 1 canonical_vpar0 = 3 </pre>	<pre> 27/01/2015 A180_cyclone5D_NEWPROF_SINGLEN_m7n5 2 a = 180. rhomin = 0.1 rhomax = 0.9 Ltheta = 6.283185307179586476925286766559005768394 Lphi = 6.283185307179586476925286766559005768394 aspect_ratio = 2.78 nb_vth0 = 6. Lmu = 7. Lmun = 0., / &EQUIL canonical_profile = 0 tau0 = 1. A0 = 1. Z0 = 1. As = 1. Zs = 1. species_frac = 1. read_ns0 = .false. read_Ts0 = .false. read_Te0 = .false. read_q = .false. profile_choice = 5 rpeak = 0.5 kappa_ns0 = 2.23 kappa_Ts0 = 6.96 kappa_Te0 = 6.96 deltar_ns0 = 0.3 deltar_Ts0 = 0.3 deltar_Te0 = 0.3 q_profile = 4 q_param1 = 0.86 q_param2 = -0.16 q_param3 = 2.52 q_param4 = 0. magnetic_drift = .true. perturb_amplitude = 0.001 perturb_choice = 2 m = 7 n = 5 FFTpotential_filter = 0 FFTpotential_filter_deltam = 0. TF_ripple = .false. delta_ripple = 0 N_ripple = 0, / &COLLISION_OPERATOR collisions = .false. version_coll = 3 energy_transfer = .false. moment_transfer = .false. nustar = 1. coll_Fmrefresh = 15 !150:10 coll_nbstep = 4, / &BUFFER_REGION </pre>
<pre> 27/01/2015 A180_cyclone5D_NEWPROF_SINGLEN_m7n5 4 hvar in fperturb = .true. adaptive_deltat = .false. advec2D_deltat = 0. reduced_deltat = 0. reduced_begin = 8000000. reduced_end = 40000000. hffilterfreq = 0 hffiltertype = 0 hffilterref = 0 deltat in diff = .true. asynchronous_writing = .false. ask_for_checksum = 1 modulo_restart_iter = 100. checkoverflow = .false. avoidtaylor = .false. deltaF in advec = .false. split_linear = .false. clampv = .false. cancel_nonlin = .false. diagf = .false., / &OUTPUT integration_CS = .true. diag_level = 6 Phi3D_saving = .false. Moments3D_saving = .false. f5D_saving = .false. Phi3D_dt_diag = 1800. Moments3D_dt_diag = 1800. f5D_dt_diag = 1800. spectra_saving = .false. spectra_dt_diag = 10. r_diags = 31 doubleSP_m = 2 doubleSP_n = 0 rst_saving = .true. iter_beg_write_trace = 1 iter_end_write_trace = 3, / </pre>	

Figure H.16: Example of input data file used for the Cyclone Base Case benchmark with GENE code (see section 4.3.2).

Appendix I. Detailed computations for local conservation laws

In this appendix, we present several detailed calculations useful for the derivation of the gyrokinetic conservation laws in section 5

Appendix I.1. Useful integrals

Let us consider two arbitrary fields X , Y and Z . The aim of this paragraph is to compute the general integral

$$\mathcal{I} \equiv \int J_x d\chi d\theta d\varphi X \nabla \cdot (Y \nabla_{\perp} Z) \quad (\text{I.1})$$

where J_x is the jacobian in space, *i.e.* $J_x = \sqrt{g}$ with g representing the determinant of the metric tensor. In order to perform this integration it is useful to write the operator $\nabla \cdot (Y \nabla_{\perp} Z)$ by using Einstein notations. Given that $\nabla \varphi \cdot \nabla \theta = \nabla \varphi \cdot \nabla \chi = 0$, we can write this operator as

$$\nabla \cdot (Y \nabla_{\perp} Z) = \frac{1}{\sqrt{g}} \partial_i (\sqrt{g} Y g^{ij} \partial_j Z) + \frac{1}{\sqrt{g}} \partial_{\varphi} (\sqrt{g} Y g^{\varphi\varphi} \partial_{\varphi} Z)$$

where i and j correspond to χ or θ . With these covariant notations, the perpendicular Laplacian-type operator we consider is equivalent to ¹²

$$\nabla \cdot (Y \nabla_{\perp} Z) = \nabla_{\perp} \cdot (Y \nabla_{\perp} Z) = \frac{1}{\sqrt{g}} \partial_i (\sqrt{g} Y g^{ij} \partial_j Z)$$

Then, the previous integral (I.1) can be written as

$$\mathcal{I} = \int J_x d\chi d\theta d\varphi X \left\{ \frac{1}{J_x} \partial_i (Y J_x g^{ij} \partial_j Z) \right\}$$

Then, using an integration by parts on the coordinate i

$$\mathcal{I} = - \int d\chi d\theta d\varphi (\partial_i X) Y J_x g^{ij} (\partial_j Z) + \left[\int d\theta d\varphi X Y J_x g^{\chi j} (\partial_j Z) \right]_{\partial L_x} + \left[\int d\theta d\varphi X Y J_x g^{\theta j} (\partial_j Z) \right]_0^{2\pi}$$

where i and j are still θ or χ . Because of the periodicity in θ , the surface term resulting from the integration by parts on $i = \theta$ (*i.e.* last term of previous equation) is equal to 0. Therefore,

$$\mathcal{I} = - \int d\chi d\theta d\varphi (\partial_i X) Y J_x g^{ij} (\partial_j Z) + \left[\int d\theta d\varphi X Y J_x g^{\chi j} (\partial_j Z) \right]_{\partial L_x} \quad (\text{I.2})$$

Using the fact that $g^{ij} = \nabla x^i \cdot \nabla x^j$ and $\nabla X = (\partial_i X) \nabla x^i$ for all field X , then

$$\mathcal{I} = - \int J_x d\chi d\theta d\varphi [Y (\partial_i X) \nabla x^i \cdot \nabla x^j (\partial_j Z)] - \int J_x d\theta d\varphi X Y \nabla_{\chi} \cdot \nabla x^j (\partial_j Z)$$

therefore,

$$\mathcal{I} = - \int J_x d\chi d\theta d\varphi [Y \nabla_{\perp} X \cdot \nabla_{\perp} Z] + \int J_x d\theta d\varphi X Y \nabla_{\chi} \cdot \nabla_{\perp} Z \quad (\text{I.3})$$

¹²The direction $\perp$ corresponds to an approximation of the poloidal plane because it is actually perpendicular to $\nabla \varphi$ rather than $\mathbf{B}$.

Then, using again an integration by parts on the coordinate j for equation (I.2) and the fact that the surface term $\left[\int d\chi d\varphi (\partial_i X) Y J_x g^{ij} Z \right]_0^{2\pi} = 0$ due to periodicity in θ ,

$$\mathcal{I} = \int d\chi d\theta d\varphi \partial_j (Y J_x g^{ij} \partial_i X) Z - \left[\int d\theta d\varphi (\partial_i X) Y J_x g^{ix} Z \right]_{\partial L_x} + \left[\int d\theta d\varphi X Y J_x g^{xj} (\partial_j Z) \right]_{\partial L_x}$$

Finally, previous equation gives

$$\int J_x d\chi d\theta d\varphi X \nabla_\perp \cdot (Y \nabla_\perp Z) = \int J_x d\chi d\theta d\varphi \nabla_\perp \cdot (Y \nabla_\perp X) Z - \left[\int J_x d\theta d\varphi Y g^{xj} \{ (\partial_j X) Z - X (\partial_j Z) \} \right]_{\partial L_x} \quad (\text{I.4})$$

Appendix I.2. Expression of $\sum_s q_s \int d\tau^ \bar{U} \partial_t \bar{F}_s$ for local energy conservation*

The opposite of the total energy source $-W = \sum_s q_s \int d\tau^* \bar{U} \partial_t \bar{F}_s$ can be divided into two parts

$$\sum_s q_s \int d\tau^* \bar{U} \partial_t \bar{F}_s = \sum_s q_s \int d\tau^* U (J_0 \cdot \partial_t \bar{F}_s) + \sum_s q_s \int d\tau^* \{ (J_0 \cdot U) \partial_t \bar{F}_s - U (J_0 \cdot \partial_t \bar{F}_s) \} \quad (\text{I.5})$$

Then, using the quasi-neutrality equation (I.2), the first term in Eq. (I.5) can be expressed as

$$\sum_{species} q_s \int d\tau^* U (J_0 \cdot \partial_t \bar{F}_s) = - \sum_{species} q_s \int J_x d\theta d\varphi U \nabla_\perp \cdot \left(\frac{n_{s,eq}}{B \Omega_s} \nabla_\perp \partial_t U \right)$$

Besides, using the general following equality (I.3) the previous equation is equivalent to

$$\sum_{species} q_s \int d\tau^* U (J_0 \cdot \partial_t \bar{F}_s) = -\frac{1}{2} \sum_{species} q_s \partial_t \int J_x d\theta d\varphi U \nabla_\perp \cdot \left(\frac{n_{s,eq}}{B \Omega_s} \nabla_\perp U \right)$$

Let us define the potential energy as $E_p \equiv \frac{1}{2} \sum_{species} q_s \int d\tau^* U (J_0 \cdot \bar{F}_s)$ this term can be identified to the time derivative of potential energy,

$$\sum_{species} q_s \int d\tau^* U (J_0 \cdot \partial_t \bar{F}_s) = \partial_t \left[\frac{1}{2} \sum_{species} q_s \int d\tau^* U (J_0 \cdot \bar{F}_s) \right] \equiv \partial_t E_p$$

Let us consider the low wavenumber approximation of the gyroaverage operator used in the code, *i.e* the Padé approximation $J_0 \simeq 1 + \frac{1}{2} \nabla \cdot \left(\frac{m_s \mu_s}{q_s^2 B} \nabla_\perp \right)$ (see section 2.5.1 for more details) then for each species

$$\mathcal{I}_E = q_s \int d\tau^* \{ (J_0 \cdot U) \partial_t \bar{F}_s - U (J_0 \cdot \partial_t \bar{F}_s) \} = \frac{m_s}{2q_s} \int d\tau^* \left\{ \partial_t \bar{F}_s \nabla \cdot \left(\frac{\mu_s}{B} \nabla_\perp U \right) - U \nabla \cdot \left(\frac{\mu_s}{B} \nabla_\perp \partial_t \bar{F}_s \right) \right\}$$

and using the result (I.3) of Appendix Appendix I

$$\begin{aligned} \mathcal{I}_E &= \frac{m_s}{2q_s} \partial_\chi \left[\int d\tau \left\{ \partial_t \bar{F}_s \nabla \cdot \left(\frac{\mu_s}{B} \nabla_\perp U \right) - U \nabla \cdot \left(\frac{\mu_s}{B} \nabla_\perp \partial_t \bar{F}_s \right) \right\} \right] \quad \text{with } d\tau = d\chi d\tau^* \\ &= \frac{m_s}{2q_s} \partial_\chi \left[- \int J_x d\chi d\theta d\varphi d^3\mathbf{v} \frac{\mu}{B} \nabla_\perp (\partial_t \bar{F}_s) \cdot \nabla_\perp U + \int J_x d\theta d\varphi d^3\mathbf{v} \partial_t \bar{F}_s \frac{\mu}{B} \nabla_\chi \cdot \nabla_\perp U \right] \\ &\quad + \frac{m_s}{2q_s} \partial_\chi \left[\int J_x d\chi d\theta d\varphi d^3\mathbf{v} \frac{\mu}{B} \nabla_\perp U \cdot \nabla_\perp (\partial_t \bar{F}_s) - \int J_x d\theta d\varphi d^3\mathbf{v} U \frac{\mu}{B} \nabla_\chi \cdot \nabla_\perp (\partial_t \bar{F}_s) \right] \end{aligned}$$

The first and third terms cancel each other out. The remaining terms

$$\mathcal{I}_E = \frac{m_s}{2q_s} \partial_\chi \left[\int d\tau^* \partial_t \bar{F}_s \frac{\mu}{B} \nabla \chi \cdot \nabla_\perp U - \int d\tau^* U \frac{\mu}{B} \nabla \chi \cdot \nabla_\perp (\partial_t \bar{F}_s) \right]$$

So finally, using the gyrocenter perpendicular stress $P_{s,\perp} = \int d^3v \bar{F}_s \mu_s B$,

$$\sum_s q_s \int d\tau^* \bar{U} \partial_t \bar{F}_s = \partial_t E_p + \frac{m_s}{2q_s} \partial_\chi \left[\int d\tau^* \partial_t \bar{F}_s \frac{\mu}{B} \nabla \chi \cdot \nabla_\perp U - \int d\tau^* U \frac{\mu}{B} \nabla \chi \cdot \nabla_\perp (\partial_t \bar{F}_s) \right]$$

Appendix I.3. Effect of the electric potential on the toroidal canonical momentum

The objective of this section is to prove that

$$dt P_\varphi = -q_s \partial_\varphi \bar{U} \quad \text{with} \quad P_\varphi = -q_s \chi + \frac{m_s I}{B} v_{G\parallel}$$

Due to the fact that P_φ is an invariant of the equilibrium motion, this result is equivalent to proving that

$$[q_s \bar{U}, P_\varphi]_{\text{GC}} = q_s \partial_\varphi \bar{U} \quad (\text{I.6})$$

where $[\cdot, \cdot]_{\text{GC}}$ indicate the Poisson brackets in the gyro-center coordinates. In the gyrokinetic framework, we recall the expression of the Poisson brackets for two given fields X and Y

$$B_{\parallel s}^* [X, Y]_{\text{GC}} = \frac{\mathbf{B}_s^*}{m_s} \cdot (\nabla X \partial_{v_{G\parallel}} Y - \partial_{v_{G\parallel}} X \nabla Y) - \frac{\mathbf{b}}{q_s} \cdot (\nabla X \times \nabla Y)$$

where $\mathbf{b} = \mathbf{B}/B$ is the unit vector parallel to the magnetic field. The quantities $\mathbf{B}_s^*$ and $B_{\parallel s}^*$ are respectively defined by Eqs (3) and $B_{\parallel s}^* = \mathbf{B}_s^* \cdot \mathbf{B}/B$, i.e $\mathbf{B}_s^* \equiv \mathbf{B} + \frac{m_s}{q_s} v_{G\parallel} \nabla \times \mathbf{b}$ and $B_{\parallel s}^* \equiv B + \frac{m_s}{q_s} v_{G\parallel} \mathbf{b} \cdot (\nabla \times \mathbf{b})$.

$$-q_s B_{\parallel s}^* [\bar{U}, P_\varphi]_{\text{GC}} = -q_s \nabla \bar{U} \cdot \left(\frac{I}{B} \mathbf{B} - \mathbf{b} \times \nabla \chi + \frac{m_s v_{G\parallel}}{q_s} \left[\frac{I}{B} (\nabla \times \mathbf{b}) - \nabla \left(\frac{I}{B} \right) \times \mathbf{b} \right] \right) \quad (\text{I.7})$$

In the chosen coordinate system, we recall that the magnetic field $\mathbf{B}$ can be written as $\mathbf{B} = I(\chi) \nabla \varphi + \nabla \varphi \times \nabla \chi$ with I a flux function, then

$$\mathbf{B} \times \nabla \chi = I \mathbf{B} - I^2 \nabla \varphi = I \mathbf{B} - B^2 R^2 \nabla \varphi \quad (\text{I.8})$$

$$\nabla \times \mathbf{B} = -(\partial_\chi I) \mathbf{B} - R^2 \partial_\chi P \nabla \varphi \quad (\text{I.9})$$

Besides,

$$\begin{aligned} \nabla \left(\frac{I}{B} \right) \times \mathbf{b} &= -\frac{I}{B^2} (\nabla B \times \mathbf{b}) + \frac{1}{B} (\nabla I \times \mathbf{b}) = -\frac{I}{B^2} (\nabla B \times \mathbf{b}) + \frac{1}{B} (\partial_\chi I \nabla \chi) \times \mathbf{b} \\ &= \frac{I}{B^2} (\mathbf{b} \times \nabla B) - \frac{1}{B} (\partial_\chi I) \mathbf{b} \times \nabla \chi \end{aligned} \quad (\text{I.10})$$

Then according to (I.8) and (I.10), equation (I.7) becomes

$$-q_s B_{\parallel s}^* [\bar{U}, P_\varphi]_{\text{GC}} = -q_s \nabla \bar{U} \cdot \left[R^2 B \nabla \varphi + \frac{m_s v_{G\parallel}}{q_s} \left(\frac{I}{B} (\nabla \times \mathbf{b}) - \frac{I}{B^2} (\mathbf{b} \times \nabla B) + \frac{1}{B} (\partial_\chi I) \mathbf{b} \times \nabla \chi \right) \right]$$

Using the fact that $\nabla \times \mathbf{b} = \mathbf{b} \times \frac{\nabla B}{B} + \frac{1}{B} \nabla \times \mathbf{B}$, the previous equation reads

$$-q_s B_{\parallel s}^* [\bar{U}, P_\varphi]_{\text{GC}} = -q_s \nabla \bar{U} \cdot \left(R^2 B \nabla \varphi + \frac{m_s v_{G\parallel}}{q_s} \left[\frac{I}{B^2} \nabla \times \mathbf{B} + \frac{1}{B} (\partial_\chi I) \mathbf{b} \times \nabla \chi \right] \right) \quad (\text{I.11})$$

Therefore,

$$\frac{I}{B^2} \nabla \times \mathbf{B} + \frac{1}{B} (\partial_\chi I) \mathbf{b} \times \nabla \chi = -R^2 \nabla \varphi \left(\frac{I}{B^2} \partial_\chi P + \partial_\chi I \right)$$

So, using the fact that the volume element in guiding-center velocity space $B_{\parallel s}^*$ can be expressed as

$$B_{\parallel s}^* = B - \frac{m_s}{q_s} v_{G\parallel} \left(\partial_\chi I + \frac{I}{B^2} \partial_\chi P \right) \quad (\text{I.12})$$

equation (I.11) becomes

$$-q_s B_{\parallel s}^* [\bar{U}, P_\varphi]_{GC} = -q_s \nabla \bar{U} \cdot \left(R^2 B_{\parallel s}^* \nabla \varphi \right) = -q_s \partial_\varphi \bar{U} \nabla \varphi \cdot \nabla \varphi R^2 B_{\parallel s}^* = -q_s B_{\parallel s}^* \partial_\varphi \bar{U}$$

which is equivalent to equation (I.6).

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